OPTO Prelim Notes

Patrick Xiao

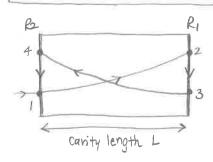
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Prelim Review

Gain condition in a laser



A mode is sustained in a can'ty if the round-trip gain is balanced exactly with loss:

$$e^{(\int gL - \alpha_i L)} R_i e^{(\int gL - \alpha_i L)} R_2 = 1$$

$$first pass & second pass & g = gain$$

$$reflection reflection $\alpha_i = intnisic loss$

$$\int = confinement$$

$$factor$$$$

What is the minimum gain required?

$$\ell^{2(\lceil g L - \alpha_i L)} = \frac{1}{R_1 R_2} \Rightarrow$$

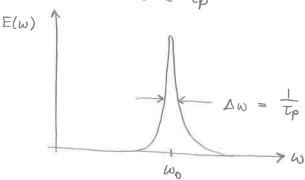
 $\Rightarrow g_{th} = \int (\alpha_i + \frac{1}{2L} \ln \frac{1}{R_1 R_2})$ gain at I'ntrinsic mimor loss threshold (usefulj gives output beam)

· Photon lifetime: consider a photon that is traveling inside a cavity. It is lost from the cavity after a characteristic time Tp

$$J(t) = I_0 e^{-t/Tp}$$

$$E(t) = E_0 e^{-i\omega_0 t} e^{-t/2Tp}$$

$$FT(E(t)) = \frac{1}{i(\omega - \omega_0) + 1/2Tp}$$
(Lorentzian) $\Delta \omega/2$



total absorption/

loss coefficient

The Q-factor is:

$$Q = \frac{\omega}{\Delta w} \Rightarrow Q = \omega \tau_p$$

 $Q = \frac{\omega}{\Delta w} \Rightarrow Q = w \tau_p$ and also note $\frac{1}{\tau_p} = \frac{c}{\sqrt{n}}$

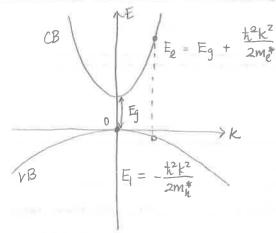
Express the threshold gain in terms of the photon lifetime:

$$\int_{0}^{\infty} g^{2} dh = \frac{1}{Tp}$$

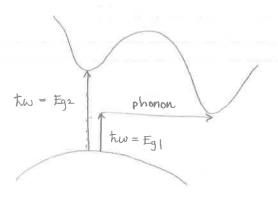
$$g_{th} = \frac{1}{V_g \int \tau_p} = \frac{1}{V_g \int Q}$$

The threshold gain is independent of diode physics, it is simply equating gain with loss





Indirect bandgap.



In an optical absorption/emission process,

- 1) Evergy conservation \Rightarrow $E_2 E_1 = \hbar \omega$
- 2) Momentum conservation:

$$k_2 - k_1 = \frac{2\pi}{\lambda}$$

- Electronic wave-number K1, K2 ~ 2t a, a~0.5nm
- Photon wavenumber $\frac{2\pi}{\lambda}$, $\lambda \sim 1000$ nm

Thus, the photon momentum is negligible to the Size of the Brillouin zone in k-space, and we have the condition $K_1 \approx K_2$

> optical transitions are vertical in E-K

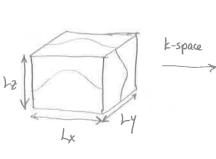
Indirect transitions are not vertical, and require the participation of a phonon to conserve momentum

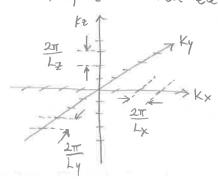
> much less likely lower absorption coefficient

Ge has a small indirect bandgap but can be alloyed with Sn to bring the direct bandgap below the indirect bandgap

Electronic density of states (3D)

Consider a box of volume V = Lx Ly Lz in which electronic wavefunctions exist:





The DOS in K-space is:

$$dN = \frac{d^3k}{\left(\frac{2\pi}{L_X}\right)\left(\frac{2\pi}{L_Z}\right)} \times 2$$

Integrate over a spherical shell be presenting a wavenumber $k = |\vec{k}|$

$$dN = \frac{4\pi k^2 dk}{\left(\frac{2\pi}{L_K}\right)\left(\frac{2\pi}{L_Y}\right)\left(\frac{2\pi}{L_z}\right)} \times 2 = \frac{k^2 dk}{\pi^2} V$$

Per rolume, n = N/V, we have

$$\frac{dn}{dk} = \frac{k^2}{\pi^2}$$

* kiretic energy
$$E' = \frac{\hbar^2 k^2}{2m} \Rightarrow \frac{dE'}{dk} = \frac{\hbar^2}{m}k$$
, $k = \frac{\sqrt{2mE'}}{\hbar}$, $dk = \frac{m}{\hbar^2 k} dE'$

$$dn = \frac{k^{2}}{\pi^{2}} dk = \frac{k^{2}}{\pi^{2}} \frac{m}{\hbar^{2}k} dE' = \frac{m}{\pi^{2}h^{2}} k dE'$$

$$\frac{dn}{dE'} = \frac{m}{\pi^{2}h^{2}} \frac{\sqrt{2mE'}}{\hbar} = \frac{1}{2\pi^{2}} \left(\frac{2m}{\hbar^{2}}\right)^{3/2} \sqrt{E'}$$

$$E' = E - E_c$$
 for electrons, $E' = E_r - E$ for holes

$$\Rightarrow P_{e}(E) = \frac{1}{2\pi^{2}} \left(\frac{2m_{e}^{*}}{\hbar^{2}}\right)^{3/2} \sqrt{E - E_{c}} \qquad P_{h}(E) = \frac{1}{2\pi^{2}} \left(\frac{2m_{h}^{*}}{\hbar^{2}}\right)^{3/2} \sqrt{E_{v} - E_{c}}$$

$$P_h(E) = \frac{1}{2\pi^2} \left(\frac{2m_h^*}{\hbar^2}\right)^{3/2} \sqrt{E_V - E}$$

Fermi-Dirac distribution: the occupancy of an electronic energy state is < 1 because of the Pauli exclusion principle

$$f_n(E) = \frac{1}{e^{(E-F_n)/kT}+1}$$
 \leftarrow electron occupancy
 $f_p(E) = \frac{1}{e^{(F_p-E)/kT}+1}$ \leftarrow hole occupancy

The hole occupancy is fp = 1 - fn(E) with Fp as the Fermi level, since a state that is occupied by a hole is a valence band state that is not occupied by an electron

Camer densities

$$n = \int_{E_c}^{\infty} f_e(E) f_n(E) dE = 2 \left(\frac{m_e^* kT}{2\pi h^2} \right)^{3/2} F_{V2} \left(\frac{F_n - E_c}{kT} \right)$$

$$\rho = \int_{-\infty}^{E_V} f_h(E) f_p(E) dE = 2 \left(\frac{m_h^* kT}{2\pi \hbar^2} \right)^{3/2} F_{V2} \left(\frac{E_V - F_p}{kT} \right)$$

where
$$F_{y_2}(\eta) \approx \begin{cases} e^{\eta} & \text{when } \eta < 1 \end{cases} \rightarrow \begin{cases} n = N_c e^{-\frac{E_c - F_n}{kT}} & \text{8oltzmann} \end{cases}$$

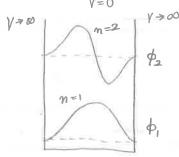
$$\begin{cases} \frac{4}{3} \left(\frac{\eta^3}{\pi}\right)^{1/2} & \eta > 1 \end{cases} \rightarrow \begin{cases} n = N_c \frac{4}{3\sqrt{\pi}} \left(\frac{F_n - E_c}{kT}\right)^{3/2} & \text{degenerate} \end{cases}$$

· Electronic density of states in 2D

Quantum well: structure in which electron is confined along one dimension, 2

$$\mathcal{Y} = \int_{A}^{1} e^{ikx} e^{iky} \phi(z) e^{-i\omega t}$$

$$E_{x}' = \frac{\hbar^{2}k_{x}^{2}}{2\pi} \quad E_{y}' = \frac{\hbar^{2}k_{y}^{2}}{2\pi}$$



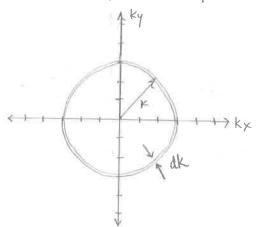
To find the energy along z, solve z=0 z=L Schnodinger's equation: $E\phi(z)=-\frac{\hbar^2}{2m}\frac{d^2\phi}{dz^2}+V(z)\phi(z)$ Let the well be infinite, so V(z)=0 for 0< z< L, $V(z)=\infty$ elsewhere The solution is:

$$\phi_{n}(z) = \sqrt{\frac{2}{L_{z}}} \sin\left(\frac{n\pi}{L_{z}}z\right) \qquad n = 1, 2, 3, \dots$$

$$E'_{nz} = \frac{\hbar^{2}}{2m} \left(\frac{n\pi}{L_{z}}\right)^{2}$$

The total kinetic energy is:
$$E_n' = \frac{h^2}{2m} \left[\frac{k_x^2 + k_y^2 + \left(\frac{n_T}{L_z} \right)^2}{k_x^2 + k_y^2 + \left(\frac{n_T}{L_z} \right)^2} \right]$$

Within a subband (specific n) the DOS is found by integrating over the K-space of the unconfined directions Kx, Ky



$$dn = \frac{1}{V} \frac{2\pi k \, dk}{\left(\frac{2\pi}{L_K}\right)\left(\frac{2\pi}{L_Y}\right)} \times 2 = \frac{k}{\pi L_Z} \, dk$$
- From dispersion relation, $dk = \frac{m}{\hbar^2 k} \, dE'$, $k = \frac{2mE'}{\hbar}$

$$dn = \frac{1}{\pi L_Z} k \frac{m}{\hbar^2 k} \, dE' \Rightarrow \frac{dn}{dE'} = \frac{m}{\pi \hbar^2 L_Z}$$
- Therefore, the density of states of each subband is:

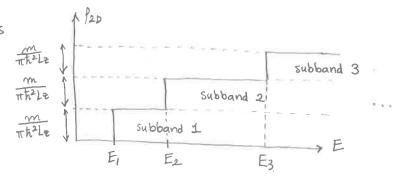
$$\beta_n^{2D}(E) = \frac{m}{\pi \hbar^2 L_z} \times H(E > E_n)$$

$$\int_{\text{Trund state energy of subband } n: E_n'(k_x = k_y = 0)$$

Sum over subbands:

$$\rho^{2D} = \sum_{n} \frac{m}{\pi \hbar^{2} L_{z}} \times H(E > E_{n})$$

The quantum well DOS consists of a series of steps:



· Electron concentration in a quantum mell

First, integrate only the first subband:

$$N_{\parallel} = \int_{E_{\parallel}}^{\infty} \frac{m}{\pi h^2 L_z} \cdot \frac{1}{e^{(E - F_c)/kT} + 1} dE$$

$$N_1 = \frac{mkT}{\pi \hbar^2 L_z} \ln \left(1 + e^{(F_c - E_I)/kT}\right)$$

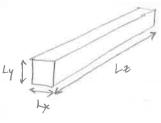
Then sum over all subbands:

$$N = \sum_{n} N_{n} = \frac{m_{r}^{*} kT}{\pi \hbar^{2} L_{z}} \sum_{n} \ln \left(1 + e^{(F_{c} - E_{n})/kT} \right)$$

$$P = \sum_{m} P_{m} = \frac{m_{h}^{*} kT}{\pi \hbar^{2} L_{z}} \sum_{m} \ln(1 + e^{(E_{m} - F_{v})/kT})$$

If $\frac{f_c - E_n}{kT} \gg 1$ (or $\frac{E_m - F_v}{kT} \gg 1$), the exponential term dominates inside the logarithm and the carrier density in the subband is linear with $F_c - E_{en}$ (or $E_m - F_v$)

· Quantum wire: confinement in two dimensions



The wire can be approximately decomposed into orthogonal guantum wells with energies

$$E_{nx} = \frac{\hbar^2}{2m} \left(\frac{n\pi}{Lx} \right)^2 , \quad E_{my} = \frac{\hbar^2}{2m} \left(\frac{m\pi}{Ly} \right)^2$$

$$\Rightarrow E' = \frac{\hbar^2}{2m} \left[\left(\frac{n\pi}{Lx} \right)^2 + \left(\frac{m\pi}{Ly} \right)^2 + k_z^2 \right]$$

There is one free dimension with dispersion relation $E_z' = \frac{\hbar^2}{2m} k_z^2$ Treat a 1D K-space

$$dn = \frac{1}{V} \frac{dk_{z}}{\left(\frac{2\pi}{L_{z}}\right)} \times 2 = \frac{2}{\pi L_{x} L_{y}} dk_{z}$$

$$= \frac{2}{\pi L_{x} L_{y}} \cdot \frac{m}{h^{2}} \frac{\hbar}{\sqrt{2mE'_{z}}} dE_{z}$$

$$\Rightarrow \frac{dn}{dE_{z}} = \left(\frac{2m}{\hbar^{2}}\right)^{1/2} \frac{1}{\pi L_{x} L_{y}} \sqrt{\frac{E'_{z} - E_{xx} - E_{my}}{E_{x}}}$$

Therefore,
$$\int_{mn}^{1D} = \frac{1}{\pi L_{x}L_{y}} \left(\frac{2m}{\hbar^{2}}\right)^{1/2} \frac{1}{\left[E' - E_{nx} - E_{my}\right]}$$

Time-dependent perturbation theory : Hamiltonian

- Conventional Hamiltonian.
$$H_0 = \frac{p^2}{2m} + V(F)$$

Upon interaction with an electromagnetic wave,
$$\vec{p} \rightarrow \vec{p} - g\vec{A}$$
 $\Rightarrow H = (\vec{p} - g\vec{A}) \cdot (\vec{p} - g\vec{A}) + V(\vec{r})$

of EM wave

$$\mathcal{H} = \frac{p^2}{2m} - \frac{3}{2m} (\vec{p} \cdot \vec{A} + \vec{A} \cdot \vec{p}) + \frac{3^2 A^2}{2m} + V(\vec{r}) = \mathcal{H}_0 + \mathcal{H}'$$

= Perturbed Hamiltonian is:

H' = =
$$\frac{9}{2m} (\vec{p} \cdot \vec{A} + \vec{A} \cdot \vec{p}) + \frac{9^2 A^3}{2m}$$
 since A is a small perturbation

Simplify

$$(\vec{p} \cdot \vec{A}) \gamma = -i\hbar \vec{\nabla} \cdot (\vec{A} \gamma) = -i\hbar [(\vec{\nabla} \vec{A}) \gamma + \vec{A} \cdot \nabla \gamma]$$

$$= \vec{A} \cdot (-i\hbar \vec{\nabla} \gamma) - (\vec{A} \cdot \vec{p}) \gamma$$

$$\Rightarrow \mathcal{H}' = -\frac{9}{m} \vec{A} \cdot \vec{p}$$

For an electromagnetic wave of a single frequency, set \overrightarrow{A} to be a plane wave:

$$\vec{A} = e^{\frac{A_0}{2}} e^{i(\vec{k} \cdot \vec{r} - \omega t)} + e^{\frac{A_0}{2}} e^{-i(\vec{k} \cdot \vec{r} - \omega t)}$$

$$e^{\text{polarization}}_{\text{vector}}$$

$$\Rightarrow \mathcal{H}' = \frac{9}{m} \frac{Ao}{2} \left[\hat{e} \cdot \vec{p} e^{i\vec{k} \cdot \vec{r}} e^{-i\omega t} + \hat{e} \cdot \vec{p} e^{-i\vec{k} \cdot \vec{r}} e^{i\omega t} \right]$$

$$= \mathcal{H}'(\vec{r}) e^{-i\omega t} + \mathcal{H}'^{\dagger}(\vec{r}) e^{i\omega t}$$

· Apply the Hamiltonian: use Schrodinger's equation

 $H\gamma(\vec{r},t) = i\hbar \frac{\partial}{\partial t} \gamma(\vec{r},t)$ where $\gamma(\vec{r},t) = \sum_{n} a_{n} \phi_{n}(\vec{r}) e^{-i\frac{E_{n}}{\hbar}t}$ eigenstates of \mathcal{H}_{i} .

 $(H_0 + H') \sum_n a_n \oint_n e^{-i\frac{E_n}{\hbar}t} = i\hbar \sum_n \frac{da_n}{dt} \oint_n e^{-i\frac{E_n}{\hbar}t} + i\hbar \sum_n a_n \oint_n \left(-i\frac{E_n}{\hbar}\right) e^{-i\frac{E_n}{\hbar}t}$

We know that $H_0 \sum_{n} a_n \phi_n e^{-iE_nt/\hbar} = i\hbar \sum_{n} a_n \phi_n (-i\frac{E_n}{\hbar}) e^{-iE_nt/\hbar}$ since we have used the eigenstates of H_0

 $\Rightarrow \mathcal{H}' \sum_{n} a_{n} \phi_{n} e^{-i\frac{E_{n}}{\hbar}t} = i\hbar \sum_{n} \frac{da_{n}}{dt} \phi_{n} e^{-i\frac{E_{n}}{\hbar}t}$

het $|m\rangle = \oint_n(\vec{r})$ and multiply both sides by $\langle m|$ to fum this into an overlap integral:

 $\sum_{n} a_{n} \langle m|H'|n \rangle e^{-i\frac{E_{n}}{\hbar}t} = i\hbar \sum_{n} \frac{da_{n}}{dt} \langle m|n \rangle e^{-i\frac{E_{n}}{\hbar}t} = i\hbar \frac{da_{m}}{dt} e^{-i\frac{E_{m}}{\hbar}t}$

 $\frac{da_{m}(t)}{dt} = \frac{1}{i\hbar} \sum_{n} a_{n} \langle m|H'|n \rangle e^{-i\frac{E_{n}}{\hbar}t} e^{+i\frac{E_{m}}{\hbar}t}$

 $\frac{0}{dt} = \frac{1}{i\hbar} \sum_{n} a_{n}(t) H'_{mn} e^{-i\omega_{mn}t}, \qquad \omega_{mn} = \frac{E_{m} - E_{n}}{\hbar}$

Use perturbation theory: $\mathcal{H}=\mathcal{H}_0+\lambda\mathcal{H}'$ $\mathcal{L}_{perturbation\ parameter}$ $a_n(t)=a_n^{(0)}(t)+\lambda a_n^{(1)}(t)+\lambda^2 a_n^{(2)}(t)+\dots$

In applying eg. (1), we use the eigenstate amplitudes an(t) of the lower order to get the dam(t)/dt of the higher order on the left side. So:

 $\frac{da_{n}^{(1)}}{dt} = \frac{1}{i\hbar} \sum_{n} a_{n}^{(2)}(t) \mathcal{H}'_{mn} e^{-i\omega_{mn}t}$

 $\frac{da_{m}^{(2)}}{dt} = \frac{1}{i\hbar} \sum_{n} a_{n}^{(1)}(t) \mathcal{H}_{mn} e^{-i\omega_{mn}t}$

$$\mathcal{H}_{fi} = \langle f | \mathcal{H}' | i \rangle$$

$$= \langle f | -\frac{gAo}{2m} \hat{e} \cdot \vec{p} e^{i\vec{k} \cdot \vec{r}} | i \rangle = -\frac{gAo}{2m} \langle f | \hat{e} \cdot \vec{p} e^{i\vec{k} \cdot \vec{r}} | i \rangle$$

. Invoke the dipole approximation: over the size of an electric dipole (or atom), the EM wave does not change much, again since $\lambda \gg a$ (lattice period) So we can say the wave has a constant phase across the dipole and Climinate the 2'E tem.

$$\mathcal{H}_{ba} = -\frac{9A_0}{2m} \langle b \mid \hat{e} \cdot \vec{p} \mid a \rangle = -\frac{9A_0}{2m} \hat{e} \cdot \vec{p}_{ba}$$

An alternative way to write this is in terms of the dipole moment:

$$\vec{p} = m \frac{d}{dt} \vec{F} = \frac{m}{i\hbar} [\vec{F}, H_0] = \frac{m}{i\hbar} (\vec{F} H_0 - H_0 \vec{F})$$

Ehnenfest's theorem

Ehvenfest's theorem
$$\langle b|\vec{p}|a\rangle = \frac{m}{i\pi} \langle b|\vec{r}H_0 - H_0\vec{r}|a\rangle = \frac{m}{i\pi} [\langle b|\vec{r}H_0|a\rangle - \langle b|H_0\vec{r}|a\rangle]$$

$$= \frac{m}{i\hbar} (E_a - E_b) \langle b| \vec{F} | a \rangle = \frac{m_o w}{i} \vec{F}_{ba}$$

$$\mathcal{H}'_{ba} = -\frac{g}{m_0} \vec{A} \cdot (-i m_0 \omega) \vec{r}_{ba} = g(i \omega \vec{A}) \cdot \vec{r}_{ba}$$

Since
$$\vec{E} = -\frac{\partial \vec{A}}{\partial t} = -iw\vec{A}$$
, we have

$$\mathcal{H}_{ba} = -\vec{E} \cdot g \vec{P}_{ba} = -\frac{gA_0}{2m} \stackrel{\wedge}{e} \cdot \vec{P}_{ba}$$

dipole moment

In a bulk semiconductor, the matrix element is isotropic (does not depend on polarization). An estimate of the magnitude from Kane's model gives 1 ~ 1 ≈ 0.5 nm

Use first-order perturbation theory: let the initial state be $a_m^{(0)}(t) = 1$ if m = i (i.e. initial wavefunction is an eigenstate of H_0) 0 if m ≠ i

$$\frac{da_{f}(t)}{dt} = \frac{1}{i\hbar} \mathcal{H}_{fi}(t) e^{i\omega f_{i}t} = \frac{1}{i\hbar} \left[\mathcal{H}_{fi}' e^{-i\omega t} + \mathcal{H}_{fi}'^{\dagger} e^{i\omega f_{i}t} \right]$$

$$= \frac{1}{i\hbar} \left[\mathcal{H}_{fi} e^{i(\omega f_{i}-\omega)t} + \mathcal{H}_{fi}^{\dagger} e^{i(\omega f_{i}+\omega)t} \right]$$

$$a_{f}(t) = -\frac{1}{\hbar} \left[\mathcal{H}_{fi} \frac{e^{i(\omega f_{i}-\omega)t}}{\omega f_{i}-\omega} + \mathcal{H}_{fi}^{\dagger} \frac{e^{i(\omega f_{i}+\omega)t}}{\omega f_{i}+\omega} \right]$$

$$|a_{f}(t)|^{2} = \frac{4|\mathcal{H}_{fi}|^{2}}{2} \sin^{2}(\frac{\omega f_{i}-\omega}{2}t) + \frac{4|\mathcal{H}_{fi}'|^{2}}{2} \sin^{2}(\frac{\omega f_{i}+\omega}{2}t)$$

 $\left|a_f^{(1)}(t)\right|^2 = \frac{4|\mathcal{H}_{fi}|^2}{\hbar} \frac{\sin^2(\frac{\omega_{fi}-\omega}{2}t)}{(\omega_{fi}-\omega)^2} + \frac{4|\mathcal{H}_{fi}^{fi}|^2}{\hbar} \frac{\sin^2(\frac{\omega_{fi}+\omega}{2}t)}{(\omega_{fi}+\omega)^2}$

Ignore the cross terms: they cancel when time-averaged

Consider the function
$$\frac{\sin^2(\frac{\omega f_i - \omega}{2} t)}{(\omega f_i - \omega)^2} = \frac{t^2}{4} \operatorname{sinc}^2(\frac{\omega f_i + \omega}{2} t)$$

As too, sinc > 0 unless Wf: > wi. In exact terms, the limit is $\lim_{t \to \infty} \frac{t^2}{t} \operatorname{sinc}^2\left(\frac{wf_i - w}{2} t\right) = \frac{\pi t}{2} S(wf_i - w)$

$$|a_{f}^{(i)}(t)|^{2} = \frac{4|\mathcal{H}_{fi}|^{2}}{\hbar} \frac{\pi t}{2} S(w_{fi} - \omega) + \frac{4|\mathcal{H}_{fi}^{t}|^{2}}{\hbar} \frac{\pi t}{2} S(w_{fi} + \omega)$$

Transition rate is:

$$\frac{d|a_f^{(i)}(t)|^2}{dt} = \frac{2\pi}{\hbar} |\mathcal{H}_f|^2 \delta(\omega_f - \omega) + \frac{2\pi}{\hbar} |\mathcal{H}_f|^2 \delta(\omega_f + \omega)$$

$$W_{i\rightarrow f} = \frac{2\pi}{\hbar} |\mathcal{H}_{fi}|^2 s(E_f - E_i - \hbar \omega) + \frac{2\pi}{\hbar} |\mathcal{H}_{fi}|^2 s(E_f - E_i + \hbar \omega)$$

Fermi's golden

$$E_f = E_i + \hbar \omega$$
optical absorption

$$Ef = E_i - tw$$
photon emission

- Consider only the absorption component of the transition rate, and use the density of final states

$$W_{ab} = \int \frac{2\pi}{\hbar} |\mathcal{H}_{ba}|^2 \delta(\xi_b - E_a - \hbar w) \, \rho(E_b) \, dE_b \qquad (i.e. \, T = 0K)$$

For simplicity, first consider a completely full valence band and a completely empty conduction band:

$$E_{b} = E_{c} + \frac{h^{2}k^{2}}{2m_{e}^{*}} \Rightarrow E_{b} - E_{a} = E_{c} - E_{v} + \frac{h^{2}k^{2}}{2} \left(\frac{1}{m_{e}^{*}} + \frac{1}{m_{h}^{*}}\right)$$

$$= E_{g} + \frac{h^{2}k^{2}}{2m_{h}^{*}} \quad \text{where} \quad m_{r}^{*} = \left(\frac{1}{m_{e}^{*}} + \frac{1}{m_{h}^{*}}\right)^{-1}$$

- Furthermore, due to the conservation of momentum $(k_a = k_b = k)$, we can also define a joint density of states that gives the effective number of possible electronic transitions when a photon of energy $\hbar w$ is absorbed

- To get the absorption coefficient, we normalize this to the incident photon flux

$$N_{ph} = \frac{1}{2} \mathcal{E}_{0} \mathcal{E}_{r} \mathcal{E}^{2} \cdot \frac{c}{m_{r}} \cdot \frac{1}{\hbar \omega} = \frac{1}{2} \mathcal{E}_{0} m_{r} \frac{c}{\hbar \omega} (-i\omega A_{0})^{2}$$
electromagnetic formula and the energy photon group energy photon energy velocity
$$\frac{1}{2} \mathcal{E}_{0} m_{r} \frac{c}{\hbar \omega} (-i\omega A_{0})^{2}$$

$$= \frac{1}{2} \mathcal{E}_{0} m_{r} \frac{c}{\hbar \omega} \omega^{2} A_{0}^{2} \Rightarrow N_{ph} = \frac{\mathcal{E}_{0} m_{r} c A_{0}^{2}}{2 \hbar} \omega$$

$$\frac{1}{2} \mathcal{E}_{0} m_{r} \frac{c}{\hbar \omega} \omega^{2} A_{0}^{2} \Rightarrow N_{ph} = \frac{\mathcal{E}_{0} m_{r} c A_{0}^{2}}{2 \hbar} \omega$$

$$\alpha_{0}(\hbar\omega) = \frac{W_{ab}}{N_{ph}} = \frac{2\pi}{\hbar} \left| -\frac{gA_{0}}{\epsilon_{m}} \hat{e} \cdot \vec{P}_{cv} \right|^{2} \cdot \frac{2\hbar}{\epsilon_{0} m_{r} c A_{0}^{2} \omega} P_{r} (\hbar\omega - E_{g})$$

$$= \frac{4\pi}{\epsilon_{0} m_{r} c A_{0}^{2} \omega} \cdot \frac{g^{2} A_{0}^{2}}{4m^{2}} \left| \hat{e} \cdot \vec{P}_{cv} \right|^{2} P_{r} (\hbar\omega - E_{g})$$

$$\mathcal{L}_{o}(\hbar\omega) = \frac{\pi q^{2}}{\varepsilon_{o} n_{r} cm^{2} \omega} |\hat{e}.\hat{p}_{cv}|^{2} \rho_{r}(\hbar\omega - E_{g}) \quad at \quad T = 0k$$

Include the occupancy probabilities of the upper and lower electronic states to obtain accurate results at finite temperature

$$W_{ab} = \int dk \frac{2\pi}{\hbar} |\mathcal{H}_{ba}|^2 S(E_g + \frac{\hbar^2 k^2}{2m_r^2} - \hbar \omega) \cdot \int_V (E_a) \cdot (1 - \int_C (E_b)) \cdot \frac{dn}{dk}$$
absorption
$$electron is occupied uno cupied electron at VB state state in CB$$

 $W_{ba} = \int dk \frac{2\pi}{\hbar} |\mathcal{H}_{ba}|^2 S(E_g + \frac{h^2 k^2}{2m_f^*} - \hbar w) \cdot f_c(E_b) \cdot (1 - f_v(E_a)) \cdot \frac{dn}{dk}$

* Note that both fc, for refer to electron occupancies here!

- Net transition rate

$$W(\hbar\omega) = \int dk \frac{2\pi}{\hbar} |\mathcal{H}_{ba}|^2 \delta(E_g + \frac{\hbar^2 k^2}{2m_r^*} - \hbar\omega) \frac{dn}{dk} \left[f_v(1 - f_c) - f_c(1 - f_v) \right]$$

$$= \int_V (E_a) - f_c(E_b)$$

Convert to an energy integral, then apply the delta function:

$$W(\hbar\omega) = \frac{2\pi}{\hbar} |\mathcal{H}_{ba}|^2 P_r(\hbar\omega - E_g) \cdot \left[-f_g(\hbar\omega - E_g) \right]$$

where
$$f_g(tw - E_g) = f_c(E_b) - f_v(E_a)$$

$$= f_c(E_g + \frac{t^2k^2}{2m_e^4}) = f_v(-\frac{t^2k^2}{2m_h^4})$$
het $E_c = E_g$, $E_v = 0$

From the delta function, we must have
$$E_g + \frac{\hbar^2 k^2}{2m_r^*} = \hbar \omega$$

$$\frac{\hbar^2 k^2}{2m_r^*} = \hbar \omega - E_g$$
So: $\frac{\hbar^2 k^2}{2m_e^*} = (\hbar \omega - E_g) \frac{m_r^*}{m_e^*}$, $\frac{\hbar^2 k^2}{2m_r^*} = (\hbar \omega - E_g) \frac{m_r^*}{m_h^*}$

So =

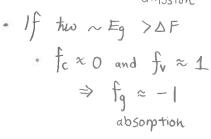
$$f_g(\hbar\omega - E_g) = f_c(E_g + \frac{m_r^*}{m_e^*}(\hbar\omega - E_g)) - f_v(-(\hbar\omega - E_g)\frac{m_r^*}{m_h^*})$$
where $f_c(E) = \frac{1}{e^{(E - F_c)/kT} + 1}$

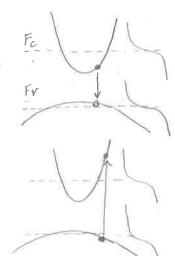
factor

 $f_{V}(E) = \frac{1}{e^{(E-F_{V})/kT}+1}$ (Make sure Fc, Fv are referenced to Ev = 0)

The shape of the Fermi inversion factor is: f_c-f_v population inversion T=0k $\Delta F=F_c-F_v$ $\Delta F=F_c-F_v$ $\Delta F=F_c-F_v$ population for inversion

• If $tw \sim Eg < \Delta F$ • $f_c \approx 1$ and $f_v \approx 0$ $\Rightarrow f_g \approx 1$ emission

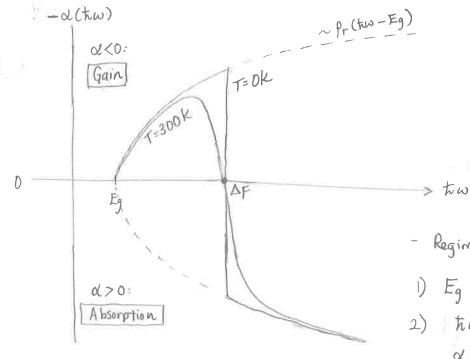




Finally, the expression for the absorption coefficient becomes modified to:

$$d(\hbar\omega) = c_0 |\hat{e} \cdot \vec{p}_{or}|^2 \rho_r (\hbar\omega - E_g) \cdot \left[-f_g (\hbar\omega - E_g) \right]$$

het's plot - a(tw)



For the $\langle \Delta F, f_g \rangle 0 \Rightarrow \alpha \langle 0 - d \rangle 0$ For the $\langle \Delta F, f_g \rangle 0 \Rightarrow \alpha \rangle 0$ $-\alpha \langle 0 \rangle$

- Regimes of bias:
- 1) Eg < tw < △F > Gain
- 2) $\hbar w = \Delta F$ \Rightarrow Transparency $\alpha = 0$: no gain or absorption
- 3) hw > △F > Absorption
- * ΔF arises from an applied voltage V. Unless the applied voltage is large ($\sim E_g$), the response mill always be absorption.
- * Once $\Delta F = Eg$, the transparency condition (Bernard Duraffourg condition) is reached: this is the onset of population inversion in the semiconductor

In our derivations, it was important that $\vec{k_a} = \vec{k_b} = \vec{k}$, Let us revisit this, and justify it more rigorously

- In a periodic crystal, the electron wavefunction has two components:
 - (1) A rapidly varying function, with periodicity given by the atomic spacing

(2) A slowly varying plane wave envelope

$$|a\rangle = u_{V}(\vec{r}) \frac{e^{i\vec{k}_{V}\cdot\vec{r}}}{\sqrt{V}}$$

$$|a\rangle = u_{C}(\vec{r}) \frac{e^{i\vec{k}_{V}\cdot\vec{r}}}{\sqrt{V}}$$

$$|a\rangle = u_{C}(\vec{r}) \frac{e^{i\vec{k}_{V}\cdot\vec{r}}}{\sqrt{V}}$$

- Evaluate the optical matrix element for a transition from 1a> to 16>

$$\mathcal{H}_{ba} = -\frac{gA_0}{2m} \hat{\ell} \cdot \langle b|\vec{p}e^{i\vec{k}op\cdot\vec{r}}|a\rangle \leftarrow \frac{\text{before applying dipole approximation}}{\vec{k}op} = \frac{gA_0}{2m} \hat{\ell} \cdot \langle b|\vec{p}e^{i\vec{k}op\cdot\vec{r}}|a\rangle \leftarrow \frac{\text{before applying dipole approximation}}{\vec{k}op} = \frac{gA_0}{2m} \hat{\ell} \cdot \int u_c(\vec{r}) \frac{e^{-i\vec{k}_c\cdot\vec{r}}}{|\vec{v}|} e^{-i\vec{k}_c\cdot\vec{r}} \hat{\ell} = \frac{gA_0}{2m} \hat{\ell} \cdot \int u_c(\vec{r}) e^{-i\vec{k}_c\cdot\vec{r}} \hat{\ell} = \frac{gA_0$$

Since u_v varies much more quickly than $e^{i\vec{k}_v \cdot \vec{r}}$, we can export ximate $\vec{\nabla} u_v \gg k_v$ (slowly varying envelope approximation)

$$\mathcal{H}_{ba}' = -\frac{gA_0}{2m} e^{i} \int u_c^*(\vec{r}) e^{-i\vec{k}_c \cdot \vec{r}} e^{i\vec{k}_0 \cdot \vec{r}} (-i\hbar \vec{\nabla} u_v(\vec{r})) e^{i\vec{k}_v \cdot \vec{r}} \frac{d^3 \vec{r}}{V}$$

Since \mathcal{U}_{c}^{*} and $\nabla \mathcal{U}_{v}$ vary on the scale of the unit cell, we can separate the above integral into two parts: one that is over the unit cell (volume Ω) and one over the whole volume V

$$\mathcal{H}_{ba} = -\frac{gA_0}{2m} \hat{e} \cdot \left(\int_{\Omega} u_c^*(\vec{r}) \frac{t}{i} \vec{\nabla} u_v(\vec{r}) \frac{d^3\vec{r}}{\Omega} \right) \left(\int_{V} e^{i(\vec{k}_V + \vec{k}_{op} - \vec{k}_c) \cdot \vec{r}} \frac{d^3\vec{r}}{V} \right)$$

$$= -\frac{gA_0}{2m} \hat{e} \cdot \vec{p}_{cV} \cdot \delta(\vec{k}_V + \vec{k}_{op} - \vec{k}_c)$$

momentum matrix element over one unit cell

Momentum conservation:

$$\hbar \vec{k}_r = \hbar \vec{k}_V + \hbar \vec{k}_{op}$$

1 Optical absorption in a quantum well

Interband optical transitions:

$$\frac{g_{m(2)}}{\sqrt{2}}$$

$$\frac{g_{m(2)}}{\sqrt{2}}$$

$$\frac{g_{m(2)}}{\sqrt{2}}$$

$$\frac{g_{m(2)}}{\sqrt{2}}$$

$$\frac{g_{m(2)}}{\sqrt{2}}$$

$$\frac{g_{m(2)}}{\sqrt{2}}$$

$$\frac{g_{m(2)}}{\sqrt{2}}$$

$$\frac{g_{m(2)}}{\sqrt{2}}$$

the quantum well wavefunction again has a component that follows the afornic-scale potential, and an envelope function that is a plane wave in (x, y) but is confined in Z:

$$|a\rangle = u_{\nu}(\vec{r}) \frac{e^{i\vec{k}\cdot\vec{p}}}{\sqrt{A}} g_{m}(z)$$

$$|b\rangle = u_{c}(\vec{r}) \frac{e^{i\vec{k}\cdot\vec{p}}}{\sqrt{A}} \phi_{n}(z)$$

$$\mathcal{H}_{ba} = \langle b|\mathcal{H}|a\rangle = -\frac{3A_0}{2m}\langle b|\hat{e}\cdot\vec{p}|e^{i\vec{k}_0p\cdot\vec{r}}|a\rangle$$

$$= -\frac{9A_0}{2m}\int d^3\vec{r} \ u_c^*(\vec{r}) \frac{e^{-i\vec{k}_t\cdot\vec{p}}}{\sqrt{A}} \phi_n^*(z) \left[\hat{e}\cdot\vec{p}|e^{i\vec{k}_0p\cdot\vec{r}}\right] u_r(\vec{r}) \frac{e^{i\vec{k}_t\cdot\vec{p}}}{\sqrt{A}} g_m(z)$$

$$= -\frac{gA_0}{2m} \int d^3\vec{r} \ u_c^{\dagger}(\vec{r}) \frac{e^{-i\vec{k}_c \cdot \vec{p}}}{\sqrt{A}} \phi_n^{\dagger}(2) e^{i\vec{k}_0 \cdot \vec{r}} \frac{e^{i\vec{k}_c \cdot \vec{p}}}{\sqrt{A}} \left[\hat{e} \cdot \vec{p} \right] u_v(\vec{r}) g_m(2)$$

- where we have used the dipole approximation again to move $e^{i\vec{E}\cdot\vec{P}}$ outside of the operator, since we know $\vec{\nabla}u_{V} >> \vec{\nabla}(e^{i\vec{E}\cdot\vec{P}})$
- Now apply the product rule:

$$[\hat{e}, \vec{p}] u_v(\vec{r}) g_m(z) = g_m(z) (\hat{e}, \vec{p}) u_v(\vec{r}) + u_v(\vec{r}) (\hat{e}, \vec{p}) g_m(z)$$

In is only a function of
$$\geq$$
, so $\vec{p}g_m(z) = \frac{\hbar}{i}\vec{\nabla}g_m(z) = p_z\hat{Z}g_m(z)$

$$\mathcal{H}_{ba} = -\frac{gA_0}{2m} \int d^3\vec{r} \ \mathcal{N}_c^*(\vec{r}) \frac{e^{-i\vec{k}_c^*\cdot\vec{p}}}{\sqrt{A}} \phi_n^*(z) e^{-i\vec{k}_c^*\cdot\vec{p}} \frac{e^{i\vec{k}_c^*\cdot\vec{p}}}{\sqrt{A}} \left[g_m(z) \left(e^{i\cdot\vec{p}} \right) u_v(\vec{r}) + \ell_z p_z g_m(z) \right]$$

$$= -\frac{8Ao}{2m} \left[\left(\int a^3 \vec{r} \frac{e^{i(\vec{k}_t - \vec{k}_t') \cdot \vec{p}}}{A} \phi_n^*(z) g_n(z) \right) \left(\int \frac{d^3 \vec{r}}{\Omega} u_c^*(\vec{r}) \left(\hat{e} \cdot \vec{p} \right) u_v(\vec{r}) \right) \right]$$

+
$$\left(\int d^{3}\vec{r} \frac{e^{i(\vec{k_{t}}-\vec{k_{t}}')\cdot\vec{p}}}{A} \phi_{n}^{*}(z) \ell_{z} P_{z} g_{m}(z)\right) \left(\int \frac{d^{3}\vec{r}}{\Omega} u_{c}^{*}(\vec{r}) u_{v}(\vec{r})\right)$$

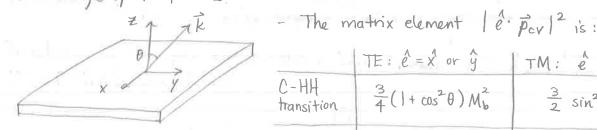
$$=-\frac{gA_0}{2m}\left(\int d^2\vec{p}\frac{e^{i(\vec{k_c}-\vec{k_c}')}\vec{p}}{A}\right)\left(\int d_2\,\Phi_n^{\dagger}(2)\,g_n(2)\right)\left(\int \frac{d^3\vec{r}}{\Omega}\,u_c^{*}(\vec{r})\left(\hat{e}\cdot\vec{p}\right)u_{\nu}(\vec{r})\right)$$

$$\mathcal{H}'_{ba} = -\frac{gA_0}{2m} \delta(\vec{k}_t - \vec{k}_t') (\hat{e} \cdot \vec{p}_{cv}) \cdot \int dz \, \phi_n^*(z) g_m(z)$$

conservation of transverse momentum

momentum matrix element for periodic parts of X confinement direction selection rules

- Polarization dependence of interband transition in quantum wells
- To analyze this, a detailed $\vec{k}\cdot\vec{p}$ methodology is needed to first obtain the dispersion relations for the conduction band, HH band, and LH band
 - This is done first by defining $\vec{k} = k\hat{z}$
 - · Then a coordinate rotation is applied to the resulting wavefunctions to get a general expression for uc, unh, uih
- The matrix elements $|\hat{e} \cdot \hat{p}_{cv}|^2$ are polarization-independent in a bulk semiconductor. In a quantum well, define \hat{z} to be the confinement direction, and let θ be the angle of \vec{k} from \hat{z} :



(R refers to the electron, not the photon)

A 1111	TE: $\ell = \hat{x}$ or \hat{y} $\frac{3}{4}(1+\cos^2\theta)M_b^2$	$TM: \stackrel{\wedge}{e} = \stackrel{\sim}{Z}$ $\frac{3}{2} \sin^2 \theta M_b^2$
C-LH transition	$\left(\frac{5}{4} - \frac{3}{4}\cos^2\theta\right)M_b^2$	$\left(\frac{1}{2} + \frac{3}{2}\cos^2\theta\right)M_b^2$

2Mb

- Notes:

· The sum of the C-HH, C-LH matrix elements is polarization independent

Sum of

C-HH, C-LH

- If $\theta = 0$ (\vec{k} in plane of ∂W)
 - = No C-HH transition for TM
 - Weak C-LH transition for TE

- Band edge:

HH > C: MTE =
$$\frac{3}{2}M_b^2$$
, MTM = O \Rightarrow HH > C responds strongly to TE,

$$LH \rightarrow C: M_{TE}^{LH} = \frac{1}{2}M_b^2, M_{TM}^{LH} = 2M_b^2 \Rightarrow LH \rightarrow C \text{ responds strongly to TM,}$$
weakly to TE

For optical amplifiers, polarization independence is desired. To get this in a quantum well, tensile strain is used, to make the HH and LH transitions roughly balanced

Interband transition selection mes

The interband selection rules are determined by the term,

$$\int_{-\infty}^{\infty} \phi_n^{f}(z) g_m(z) dz \wedge \int_{-\infty}^{\infty} f(z) g_m(z) dz \wedge \int_{-\infty}^{\infty} f(z) dz$$
conduction valence subband

$$\int_{-\infty}^{\infty} \phi_n^{\sharp}(z) g_m(z) dz \sim \int_{-\infty}^{\infty} \sin\left(\frac{n\pi}{L_z}z\right) \sin\left(\frac{m\pi}{L_z}z\right) dz = \delta mn$$

> Interband fransitions are only allowed between the nth subbands of the conduction and valence bands.

Quantum well interband absorption and gain

The general expression is the same as in the bulk case,

$$g(\hbar\omega) = -\alpha(\hbar\omega) = Col\hat{e} \cdot \vec{p}_{cv}|^2 P_r^{2p}(\hbar\omega - E_g) f_g(\hbar\omega - E_g)$$

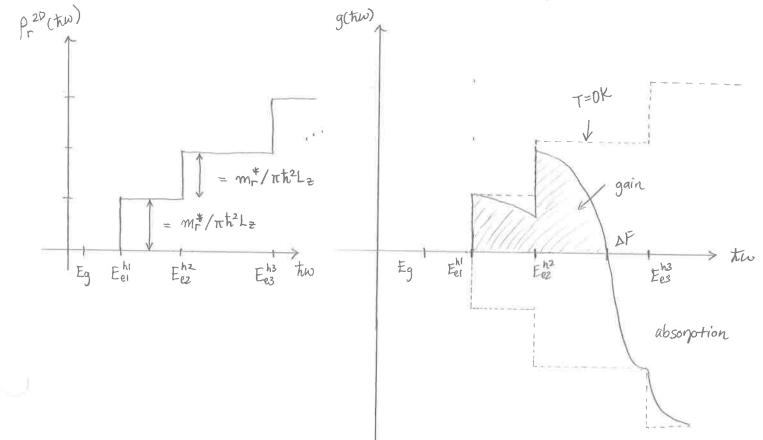
where the differences are in:

(1) The momentum matrix element |ê. pcv|2 has a new polarization dependence and selection rules between subbands

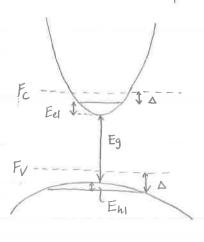
(2) The joint DOS is now given by:

$$P_r^{2D}(\hbar\omega - E_g) = \sum_{m} \sum_{n} \frac{m_r^*}{\pi \hbar^2 L_z} \mathcal{H}(\hbar\omega - E_{en}^{hm}) = \sum_{n} \frac{m_r^*}{\pi \hbar^2 L_z} \mathcal{H}(\hbar\omega - E_{en}^{hn})$$

where $E_{en} = (E_c + E_{en}) - (E_V - E_{hm}) = E_q + E_{en} + E_{hm}$



- · Transparency condition in a quantum well
 - · Consider an undoped, unstrained quantum well, and use the bulk values of electron and hole effective mass (which is not exactly valid)
 - The transparency condition occurs when $\Delta F = \hbar w$. What is the carrier density when this occurs?



$$N = \frac{m_e^* kT}{\pi k^2 L_2} \ln \left(1 + e^{(F_c - (E_g + E_{e1}))/kT} \right)$$

$$P = \frac{m_n^* kT}{\pi k^2 L_2} \ln \left(1 + e^{(-E_{h1} - F_v)/kT} \right)$$

where the energy is referenced to the bulk valence level Ev=0.

Now notice that the transparency condition implies

$$\Delta F = \hbar \omega = E_g + E_{el} + E_{hl}$$

$$F_c - F_V = E_g + E_{el} + E_{hl}$$

$$F_c - (E_g + E_{el}) = -(-E_{hl} - F_V) = \Delta \cdot kT$$

Therefore,
$$N = \frac{me^{kT}}{\pi h^2 L_z} \ln (1 + e^{\Delta})$$

$$P = \frac{mi^{*}kT}{\pi h^2 L_z} \ln (1 + e^{-\Delta})$$

Net reutrality in an undoped semiconductor implies N = P

$$\Rightarrow m_e^* \ln(1 + e^{\Delta}) = m_h^* \ln(1 + e^{-\Delta})$$

Numerically solving this using $m_e^*/m_h^* = \frac{1}{7.5}$, which may be accurate for an unstrained GaAs quantum well, with $m_e^* = 0.067 m_0$:

$$\Delta = 1.41$$
; therefore

$$N_{fr} = \frac{m_e^* kT}{\pi h^2 L_z} \ln(1 + e^{\Delta}) = \frac{1}{L_z} (1.2 \times 10^{12} \text{ cm}^{-2})$$

$$P_{fr} = N_{fr} \approx 1.2 \times 10^{18} \text{ cm}^{-3}$$

$$\Delta = 1.41$$
 implies that F_c lies inside the conduction band and outside the valence band at transparency, due to effective mass asymmetry

This is the necessary current density for gain to occur. Note again that this analysis holds only for an undoped, unstrained material; for the effects of strain and doping, see §6!

Quantum well intersubband transitions

• Intersubband transitions respond to energies of $E_{e2}-E_{e1}=\frac{\hbar^2}{2m_e^2}\left(\frac{2\pi}{L_z}-\frac{\pi}{L_z}\right)$

≈ 168mV for Lz = 10nm GaAS

· Transition matrix element:

The wavefunctions are

$$|a\rangle = u_{c}(\vec{F}) \frac{1}{\sqrt{A}} e^{i\vec{k}\cdot\vec{P}} \phi_{1}(z) \leftarrow 1^{st} e^{-subband}$$

$$|b\rangle = u_{c}(\vec{F}) \frac{1}{\sqrt{A}} e^{i\vec{k}\cdot\vec{P}} \phi_{2}(z) \leftarrow 2^{nd} e^{-subband}$$

To evaluate Hba, it is easier to use the alternate form:

$$\mathcal{H}'ba = -\vec{E} \cdot g\vec{F}ba = -E_0 \hat{e} \cdot \langle b|g\vec{F}|a \rangle \cdot e^{i\vec{K}op\vec{F}} \stackrel{\text{slow}}{=} -gE_0 \hat{e} \cdot \int u_c^*(\vec{F}) \frac{1}{\sqrt{A}} e^{-i\vec{K}\dot{t}\cdot\vec{p}} \phi_2^*(z) \vec{F} u_c(\vec{F}) \frac{1}{\sqrt{A}} e^{i\vec{K}\dot{t}\cdot\vec{p}} \phi_2^*(z) \hat{e}^{i\vec{K}\dot{t}\cdot\vec{p}} \phi_2^*(z) \hat{e}^{i\vec{K}\dot{t}\cdot\vec{p}} \phi_2^*(z) \hat{e}^{i\vec{K}\dot{t}\cdot\vec{p}} \phi_2^*(z) \hat{e}^{i\vec{K}\dot{t}\cdot\vec{p}} \phi_2^*(z) \hat{e}^{i\vec{K}\dot{t}\cdot\vec{p}} \hat{e}^{i\vec{K}\dot{t}\cdot\vec{p$$

- In the above step, the x and y components of \vec{F} vanish because integrating $(x\hat{x} + y\hat{y}) e^{i(\vec{k}t \vec{k}t') \cdot \vec{P}}$ returns 0
- · Note that:

$$\int u_c^*(\vec{r}) u_c(\vec{r}) \frac{d^3\vec{r}}{s2} = 1$$

$$\int \frac{1}{A} e^{i(\vec{k}\vec{t} - \vec{k}\vec{t}') \cdot \vec{p}} d\vec{p} = \delta(\vec{k}\vec{t} - \vec{k}\vec{t}')$$

$$\Rightarrow \mathcal{H}'_{ba} = -gE_0 \hat{e} \cdot \int \phi_2^{\dagger}(z) \phi_1(z) z \hat{z} dz \cdot \delta(\vec{k_t} - \vec{k_t}')$$

$$\mathcal{H}_{ba} = -E_0(\hat{e}\cdot\hat{z})\mu_{21}\delta(\vec{k_t}-\vec{k_t})$$

where
$$\mu_{21} = g \int \phi_2^*(z) \phi_1(z) \neq dz$$

dipole moment between subband envelope functions

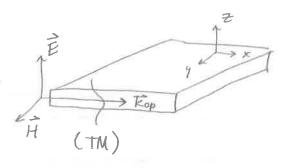
Polarization dependence of intersubband transitions

Note that the dipole moment is always polarized along the confinement direction, in this case $\hat{\Xi}$.

$$\mathcal{H}_{ba} = - E_0 \left(\hat{e} \cdot \hat{z} \right) \mu_{21} \cdot \delta(\vec{k_t} - \vec{k_t})$$

- If the incident field does not have a component along 2, there will be no intersubband matrix element; i.e. responds only to TM polarization
- If \vec{E} is polarized at an angle θ away from \hat{z} , the matrix element diminishes with θ :

$$\mathcal{H}_{ba}' = - E_0 \mu_{21} \cos \theta \cdot \delta(\vec{k_t} - \vec{k_t}')$$



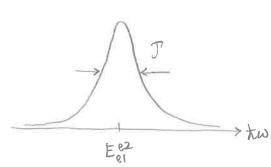
Even more generally:

- * An electric field can induce an intersubband transition if it has a component along one of the confinement directions
- * If there are multiple confinement directions (e.g. quantum wire), intersubband transitions can occur along any confinement direction that is excited by a nonzero component of \hat{E}
 - If E is polarized only along one of the confinement directions, it can only induce transitions between two subbands of that one confinement direction

- Absorption: inter-subband

$$\alpha(tw) = \frac{1}{N_{ph}} \frac{2\pi}{\hbar} \left| -E_0 \hat{e} \cdot \vec{\mu}_{21} \right|^2 g\left(E_{e1} - \hbar w\right) \cdot \frac{2}{V} \sum_{k} \left[f_c \left(E_{e1}\right) - f_c \left(E_{e2}\right) \right]$$
in adent flux matrix energy conservation counting all initial and final states

Energy conservation: without accounting for energy broadening, $g(E_{ei}^{e2}-\hbar\omega)=8(E_{ei}^{e2}-\hbar\omega) \text{ which has an infinite peak}$



$$g(E_{el}^{c2} - k\omega) = \frac{1}{\pi} \frac{F/2}{(E_{el}^{e2} - k\omega)^2 + (F/2)^2}$$

This gives the shape of the absorption spectrum a (thw)

$$\frac{2}{V}\sum_{K_1}f_c(E_{el}) = N_1 \leftarrow \text{subband } 1$$

$$\frac{2}{V} \sum_{k_2} f_V(\bar{E}_{e2}) = N_2 \leftarrow \text{#e-in}$$
subband 2

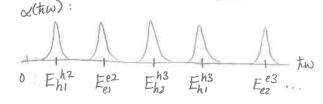
- Then the expression becomes :

$$d(\hbar w) = \frac{\pi w}{\varepsilon_0 \, m_{rc}} \left| \stackrel{1}{e} \cdot \vec{\mu}_{21} \right|^2 g(E_{el}^{e2} - \hbar w) (N_1 - N_2)$$

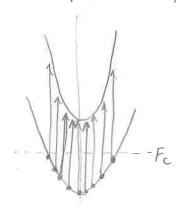
When integrated, $\int g \ d\omega \rightarrow 1$ and we can assume it is narrow around E_{el}^{2l} so the ω out front can be pulled out:

individual peak values depend on w, Mab, DN

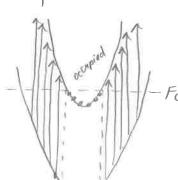
$$\int d(\hbar\omega) d(\hbar\omega) = \frac{\pi E_{el}^{e2}}{\varepsilon_0 m_r c \hbar} |\hat{e} \cdot \vec{\mu}_{2l}|^2 (N_l - N_2)$$



· Absorption as a function of carrier concentration

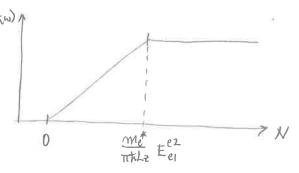


only the first subband is filled



Both subbands filled

once both subbands are filled, the number of possible transitions does not change! So the absorption coefficient tends to saturate



- More rigorously,

$$N_1 - N_2 = N_1$$
 if $N_2 = 0$
 $N_1 = N_2 = \frac{m_e^*}{\pi h^2 L_z} (F_c - E_{e1}) - \frac{m_e^*}{\pi h^2 L_z} (F_c - E_{e2}) = \frac{m_e^*}{\pi h^2 L_z} (E_{e2} - E_{e1})$

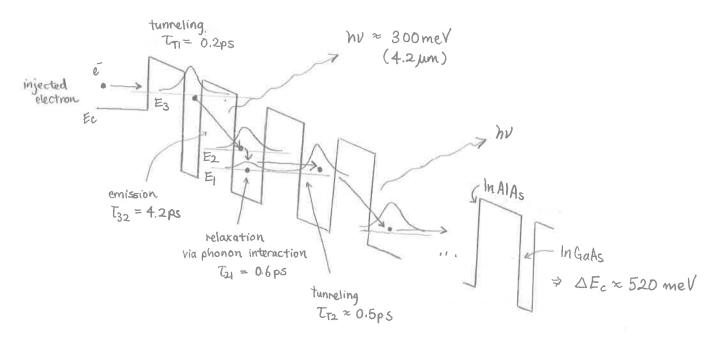
fixed!

Intersubband selection rules

This is governed by
$$\int_{0}^{L} Z \oint_{n}^{*}(Z) \oint_{m} (Z) dZ \text{ is Zero if the integrand is odd}$$

- So as long as In and Im have different parity (i.e. one is even and one is odd), this integral will be nonzero
- Thus, the allowable transitions are

- · Quantum cascade lasers gain in inter-subband transitions
- Form a heterojunction superlattice, in which the energy bands in very closely spaced quantum wells are strongly coupled
- The QCL is based on a 3-level system:



- Due to the slow intersubband radiation process ($T_{32} = 4.2ps$) relative to the speed of filling E₃ and emptying E₂, the population inversion is very efficient
- One injected electron can produce several infrared photons as it travels through the superlattice (cascading, or carrier recycling)
- Yenry useful emitters in the mid-IR and THz range

$$g(\hbar\omega) = C_0 |\hat{e} \cdot \vec{p}_{ov}|^2 p_r(\hbar\omega) \left(f_c(E_e) - f_v(E_h) \right)$$

$$= C_0 |\hat{e} \cdot \vec{p}_{ov}|^2 \left[p_r(\hbar\omega) f_c(E_e) - p_r(\hbar\omega) f_v(E_h) \right]$$

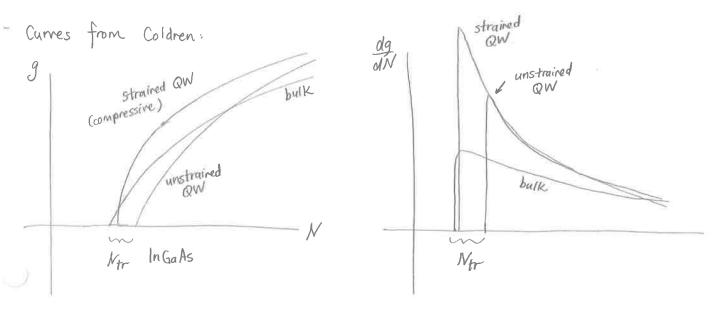
$$= C_0 |\hat{e} \cdot \vec{p}_{ov}|^2 \left[p_r(\hbar\omega) f_c(E_e) + p_r(\hbar\omega) (1 - f_v(E_h)) - p_r(\hbar\omega) \right]$$

$$\approx \frac{dN}{dE} (E_e) \qquad \approx \frac{dP}{dE} (E_h)$$

- Since fr is proportional to the density of states fe, fn in either band, the two terms above correspond approximately to the carrier density (per unit everyy) at the two band edges Ee and Eh

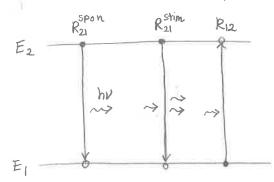
Ehl

- Thus, if the band edge carrier density is a sensitive function of the Fermi level positions, the differential gain becomes larger. This is achieved when two conditions are met:
 - 1) p is steep at the band edges -> use a quantum well
 - 2) The Fermi levels lie close to their respective band edges, since the occupancy functions fc, for change most rapidly at the Fermi levels -> use strain to get symmetric electron and hole bands



(5) Spontaneous emission

· The most famous derivation of the spontaneous emission rate invoke's Einstein's A&B coefficients:



$$R_{21}^{spon} = A_{21} f_{2} (1 - f_{1})$$

$$R_{21}^{stim} = B_{21} f_{2} (1 - f_{1}) \cdot P(E_{21})$$

$$R_{12} = B_{12} (1 - f_{2}) f_{1} \cdot P(E_{21})$$
where $P(E_{21}) = \frac{\# photons}{\text{volume energy}} = \langle mph \rangle \cdot \int_{ph} (E_{21})$

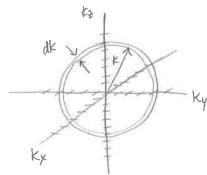
Bose-Einstein Photon DOS

at $h\nu = E_{21}$

distribution

Photon density of states

First find DOS in K-space, then use the photon dispersion relation



- Consider a spherical shell in k-space w/ radius k

$$dN = \frac{4\pi k^2 dk}{(2\pi/L)^3} \times 2 = \frac{4\pi k^2}{\pi^2} dk$$

$$dn = \frac{k^2}{\pi^2} dk$$

Now use $E = \hbar \omega = \frac{\hbar ck}{n}$ $k = n \cdot \frac{E}{\hbar c}, \quad dk = \frac{m}{\hbar c} dE$

$$\Rightarrow dn = \frac{1}{\pi^2} \frac{n^2 E^2}{h^2 c^2} \cdot \frac{n}{hc} dE = \frac{1}{\pi^2} \frac{n^3 E^2}{h^3 c^3} dE$$

$$\int_{\text{ph}}(E) = \frac{8\pi n^3 E^2}{h^3 c^3}$$
 photon DOS

Meanwhile the Bose-Einstein distribution is $\langle mph \rangle = e^{E/kT} - 1$ This represents the expectation value of the number of photons that occupy a state at energy E, derived from statistical mechanics. - In thermal equilibrium,

$$R_{21}^{spon} + R_{21}^{stim} = R_{12}$$

$$A_{21} f_2(1-f_1) + B_{21} f_2(1-f_1) P(E_{21}) = B_{12} f_1(1-f_2) P(E_{21})$$

$$P(E_{21})$$
 $\left[B_{12}f_{1}(1-f_{2})-B_{21}f_{2}(1-f_{1})\right]=A_{21}f_{2}(1-f_{1})$

$$P(E_{21}) = \frac{A_{21}f_2(1-f_1)}{B_{12}f_1(1-f_2) - B_{21}f_2(1-f_1)}$$

Recall that the Fermi-Dirac occupancy is

$$f_{1} = f(E_{1}) = \frac{e^{(E_{1} - E_{F})/kT}}{e^{(E_{1} - E_{F})/kT}}, \quad |-f_{1}| = \frac{e^{(E_{1} - E_{F})/kT}}{e^{(E_{1} - E_{F})/kT} + 1}$$

$$f_{2} = f(E_{2}) = \frac{e^{(E_{2} - E_{F})/kT}}{e^{(E_{2} - E_{F})/kT}}, \quad |-f_{2}| = \frac{e^{(E_{2} - E_{F})/kT}}{e^{(E_{2} - E_{F})/kT} + 1}$$

All the Fermi products have the same denominator, so:

$$P(E_{21}) = \frac{A_{21}e^{(E_1 - E_F)/kT}}{B_{12}e^{(E_2 - E_F)/kT} - B_{21}e^{(E_1 - E_F)/kT}} = \frac{A_{21}}{B_{12}e^{(E_2 - E_1)/kT} - B_{21}}$$

Equate this to the known photon density P(E21):

$$P(E_{21}) = \frac{A_{21}}{B_{12}e^{E_{21}/kT} - B_{21}} = \frac{8\pi n^3 E_{21}^2}{c^3 h^3} \cdot \frac{1}{e^{E_{21}/kT} - 1}$$

$$\left(\frac{A_{21}}{B_{12}}\right) \frac{1}{e^{E_{21}/kT} - \frac{B_{21}}{B_{12}}} = \left(\frac{8\pi n^8 E_{21}^2}{c^3 h^3}\right) \frac{1}{e^{E_{21}/kT} - 1}$$

This equation immediately implies:

1) B12 = B21: the stimulated emission rate equals the absorption rate

2)
$$A_{21} = \left(\frac{8\pi m^3 E_{21}^2}{c^3 h^3}\right) B_{21}$$

proportionality between spontaneous emission and absorption pre-factors

Recall that the absorption coefficient is femi inversion $\alpha(E_{21}) dE = -g(E_{21}) dE = \frac{m}{c} \beta_{21} \left[f_1 - f_2 \right]$

Meanwhile the spontaneous emission rate is

$$\Gamma_{21}(E_{21}) dE = A_{21} f_{2}(1-f_{1})$$

$$= \left(\frac{8\pi m^{3} E_{21}^{2}}{c^{3} h^{3}}\right) B_{21} f_{2}(1-f_{1})$$

$$= \left(\frac{8\pi m^{3} E_{21}^{2}}{c^{3} h^{3}}\right) \frac{c}{n} \frac{\alpha(E_{21}) dE}{f_{1}-f_{2}} f_{2}(1-f_{1})$$

$$\Gamma_{21}(E_{21}) = \frac{8\pi m^{2} E_{21}^{2}}{c^{2}h^{3}} \frac{f_{2}(1-f_{1})}{f_{1}-f_{2}} \alpha(E_{21})$$

where
$$\frac{f_{2}(1-f_{1})}{f_{1}-f_{2}} = \frac{e^{(E_{2}-F_{c})/kT}}{\left[e^{(E_{2}-F_{c})/kT}+1\right]\left[e^{(E_{1}-F_{V})/kT}+1\right]} \cdot \frac{e^{(E_{2}-F_{c})/kT}}{\left[e^{(E_{2}-F_{c})/kT}+1\right]\left[e^{(E_{1}-F_{V})/kT}+1\right]} \cdot \frac{e^{(E_{2}-F_{c})/kT}}{\left[e^{(E_{2}-F_{c})/kT}+1\right]\left[e^{(E_{1}-F_{V})/kT}+1\right]}$$

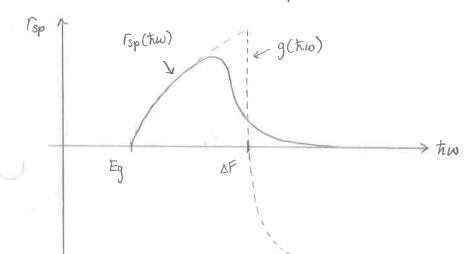
$$= \frac{e^{(E_{1}-F_{V})/kT}}{e^{(E_{2}-F_{c})/kT}-e^{(E_{1}-F_{V})/kT}} = \frac{e^{(E_{2}-F_{c})/kT}-1}{e^{(E_{2}-F_{c})/kT}-1}$$

Therefore:

Note that this also depends on AF, near or above transparency

$$r_{SP}(E) = \frac{8\pi n^2 E^2}{c^2 h^3} \alpha(E) \frac{1}{e^{(E-\Delta F)/ET}-1}$$

This is the Roosbroeck-Shockley relation!



spontaneous emission rate (#/s.m³·ev)

- · For E<< DF, last term > -1
- · For $E \gg \Delta F$, last term $\Rightarrow 0$
- · On this plot, designate emission as positive, absorption as regative
- At transparency, $E = \Delta F$, the Bose-Einstein distribution blows up but this is balanced by a going to 0 due to the Fermi inversion factor

Spontaneous emission lifetime: an alternative expression for rsp is

$$F_{sp}(\hbar w) = \frac{1}{T_{sp}} f_r(\hbar w - E_g) \left[f_c(E_2) \cdot (1 - f_v(E_1)) \right]$$

$$f_e(\hbar w) : \text{probability of emission}$$

Equate this to the previously derived expression:

$$T_{Sp}(\hbar\omega) = \frac{8\pi n^2 E^2}{c^2 h^3} \alpha(\hbar\omega) \frac{1}{e^{(E-\Delta F)/kT} - 1} = \frac{1}{T_{Sp}} P_r(\hbar\omega - E_g) f_e(\hbar\omega)$$

$$\frac{8\pi n^2 (\hbar\omega)^2}{c^2 h^3} g(\hbar\omega) \frac{1}{1 - e^{(E-\Delta F)/kT}} = \frac{1}{T_{Sp}} P_r(\hbar\omega - E_g) f_e(\hbar\omega)$$

$$\frac{8\pi n^2 (\hbar\omega)^2}{c^2 h^3} \left[C_0 l_e^4 P_{cr} l_e^2 P_r(\hbar\omega - E_g) f_g(\hbar\omega) \right] \frac{1}{1 - e^{(E-\Delta F)/kT}} = \frac{P_r(\hbar\omega - E_g) f_e(\hbar\omega)}{T_{Sp}}$$

Recall that

$$\frac{f_2(1-f_1)}{f_1-f_2} = -\frac{f_c(1-f_V)}{f_c-f_V} = -\frac{f_e}{f_g} = \frac{1}{e^{(E-\Delta F)/kT}-1}$$

$$\Rightarrow \frac{8\pi n^2(\hbar \omega)^2}{c^2h^3} \cdot C_0 |\hat{e}\cdot\hat{p}_{cV}|^2 \cdot \left(\frac{f_g}{f_e}\right) = \frac{1}{1+e^{(E-\Delta F)/kT}} = \frac{1}{T_{SP}}$$

So:
$$T_{SP} = \frac{c^2h^3}{8\pi n^2(\hbar \omega)^2} \cdot \frac{1}{c_0 |\hat{e}\cdot\vec{p}_{ev}|^2}$$
multiply constants (
$$T_{SP} = \frac{hc^3 \, \epsilon_0 \, m_0^2}{2\, g^2 \, n \, \omega} \cdot \frac{1}{|\hat{e}\cdot\vec{p}_{ev}|^2} \sim 1 \, ns$$
 for semiconductors (interband)

6 Effects of doping and strain

- · Threshold current: the threshold current is best derived from the rate equations in steady state (see § 9)
 - At transparency or at threshold, $S \approx 0$

$$\Rightarrow \frac{dN}{dt} = \frac{\eta_i I_{th}}{g \gamma} - \frac{N_{th}}{\tau} = 0$$

The current must exactly balance the rate of carrier loss in steady state:

Ith =
$$\frac{gV}{\eta_i}$$
 [APth + BNth Pth + CNth Pth] (n-type active region)

At threshold in a III-V optoelectronic material, the radiative rate usually dominates, so

Let $d=L_Z$ be the confinement direction and the transverse area of the quantum well is $A_t=l\mathcal{W}$

- If the guantum well is undoped, N+h = Pth so

$$J_{th} = \frac{1}{\eta_i} g_{Lz}(BN_{th}) N_{th} = \frac{1}{\eta_i} \frac{g_{Lz}}{\tau_e} N_{th}$$

where BNHh is in units of s^{-1} and $T_e = 1/BNHh$ can be considered the carrier life time

- The transparency current follows the same relationship with N:

In order to reduce Itr, and therefore Ith, we must reduce the product Ntr × Ptr.

· Undoped quantum well:

= In § 4, we found that $N_{tr} = P_{tr} = 1.2 \times 10^{18} \text{ cm}^{-3}$ Let our figure of ment be $\overline{N}_{tr} = \sqrt{N_{tr} P_{tr}} = 1.2 \times 10^{18} \text{ cm}^{-3}$

with
$$m_e^*/m_h^* = 1/7.5$$
, $m_e^* = 0.067 m_o$
 $L_Z = 10 \text{ nm}$

- Effect of doping

· n-doping:

Charge reutrality becomes $N = p + N_D^+$ where we assume $N_D^+ \cong N_D$ This becomes:

$$\frac{m_e^* kT}{\pi h^2 L_z} \ln (1 + e^{\Delta}) = \frac{m_h^* kT}{\pi h^2 L_z} \ln (1 + e^{-\Delta}) + N_D$$

Let No = 1 x 1018 cm-3 and numerically solve the above

$$A = 2.12, \quad N_{tr} = 1.6.2 \times 10^{18} \text{cm}^{-3}$$

$$P_{tr} = 6.2 \times 10^{17} \text{ cm}^{-3} \Rightarrow N_{tr} = 1.0 \times 10^{18} \text{ cm}^{-3}$$

$$(\sim 17\% \text{ smaller than undoped!})$$

· up-doping:

Charge reutrality becomes $N + N_A = P$ where we assume $N_A = N_A$ This becomes:

$$\frac{m_e^* kT}{\pi \hbar^2 L_z} \ln (1 + e^{\Delta}) = \frac{m_h^* kT}{\pi \hbar^2 L_z} \ln (1 + e^{-\Delta}) - N_A$$

Let $N_A = 1 \times 10^{18} \text{ cm}^{-3}$ and numerically solve:

$$\Rightarrow \Delta = 0.878$$
, Ntr = 8.9 x 10¹⁷ cm⁻³
Ptr = 1.9 x 10¹⁸ cm⁻³ $\Rightarrow N_{tr} = 1.3 \times 10^{18}$ cm⁻³
(~8% larger than undoped)

Conclusions:

- To reduce the threshold/fransparency current density, m-doping is desirable
 - This is due to the lower effective mass of the conduction band,
 which populates slower than the valence band for a given shift in Fermi levels
- To get higher differential gain, p-doping is desirable
 - · This is because in the p-doped case, the Fermi levels both shift down in energy and become closer to their respective band edges
- The effect on transparency cument is relatively small in both cases for doping $\sim 10^{18} \, \text{cm}^{-3}$ ($\sim 10-20\%$ difference)
- Both types of doping will significantly increase the majority carrier concentration, and can increase the Auger rate at transparency if the minority carrier concentration does not substantially decrease

- Overall, how does a quantum well laser improve upon a bulk laser?
 - 1) Lower threshold current, since $Jth = \frac{1}{\eta_1} gd$ $\frac{Nth}{Te}$
 - In bulk, $d \sim 100$ nm or more. In a quantum well, $d = Lz \sim 10$ nm $\Rightarrow 10x$ reduction
 - Possibility of strain to further reduce Ntr
 - 2) Higher differential gain: do is enhanced by the steepness of the DOS function. This is desirable for extending modulation bandwidth
 - 3) Reduced frequency chirping under direct modulation

· Benefit of strain

- Consider the hypothetical case where $m_e^* = m_h^* = 0.067 m_o$ We still have: $N = \frac{m_e^* kT}{\pi h^2 L_z} \ln (1 + e^{\Delta})$

Assume undoped:

$$N = P \Rightarrow \ln(1 + e^{\Delta}) = \ln(1 + e^{-\Delta})$$

So:
$$\Delta = 0$$

Therefore,

$$P_{tr} = N_{tr} = \frac{m_e^{t} kT}{Th^2 L_z} \ln 2 = 5.1 \times 10^{17} cm^{-3}$$

There is a 2x reduction in Ntr!

- Effect on threshold current: Ith goes down as Itr goes down

$$J_{tr} = \frac{1}{\eta_1} g L_z \left(AN_{tr} + BN_{tr}^2 + CN_{tr}^3 \right) < \frac{1}{2} J_{tr} \text{ of undoped, unstrained QW}$$

$$2 \times \text{ reduction}$$

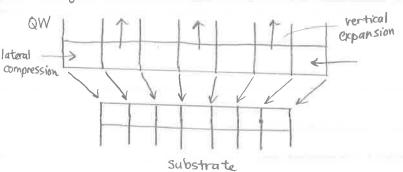
$$4 \times \text{ reduction}$$

$$C \text{ is also reduced by strain}$$

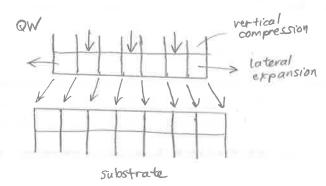
. Note: the Lz dependence in Ntr cancels out in Jtr.

· Effects of strain on band structure

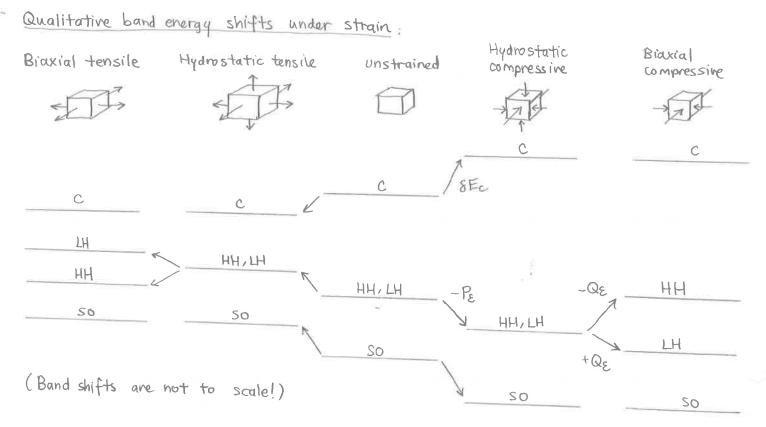
- Compressive strain: the QW has a larger lattice constant than the substrate

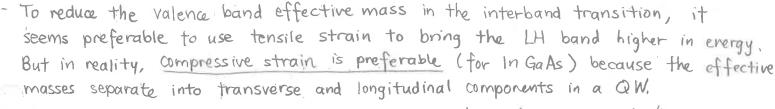


Tensile strain: the QW has a smaller lattice constant than the substrate



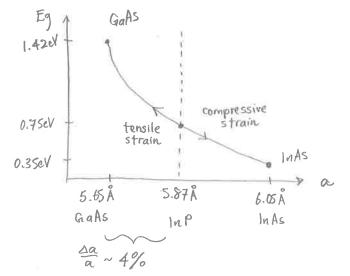
- Directions of strain:
 - Hydrostatic: equal strain in all 3 directions
 - > Shifts energy levels (changes bandgap) but affects all bands equally
 - · Biaxial: strain along 2 of 3 directions (useful for QWs)
 - · Uniaxial: strain only along 1 direction
- Strain is only possible if the quantum well is below a critical thickness; above this thickness, the lattice relaxes to its native state, and generates many structural defects in the process





Strain & bandgap: we introduce strain commonly by adjusting the Ga/In ratio

of In GaAs P grown on In P



- · Notice that by applying tensile strain by increasing the Ga content,
 - Eg decreases due to tensile strain effect
 - Eg increases due to larger Ga content

6.05Å Net effect: Eg increases

$$\mathcal{E} = \frac{q_0 - a(x)}{q_0} = \frac{\Delta a}{q_0}$$

Stress/strain analysis

• Stress: deforming force per unit area applied in a particular direction -> o • Strain: relative change in length caused by deforming force -> c

In general, both or and ε are 3×3 tensors. We will assume $\sigma_{ij}=0$ if $i\neq j$ meaning only normal forces σ_{x} , σ_{y} , σ_{z} act to deform the lattice; no shear or rotation occurs. This means we also only consider ε_{ij} , i=j.

- The stress and strain are related by the stiffness tensor C:

$$\begin{bmatrix} \sigma_{x} \\ \sigma_{y} \\ \sigma_{z} \end{bmatrix} = \begin{bmatrix} C_{11} & C_{12} & C_{13} \\ C_{21} & C_{22} & C_{23} \\ C_{31} & C_{32} & C_{33} \end{bmatrix} \begin{bmatrix} \varepsilon_{x} \\ \varepsilon_{y} \\ \varepsilon_{z} \end{bmatrix} \approx \begin{bmatrix} C_{11} & C_{12} & C_{12} \\ C_{12} & C_{11} & C_{12} \\ C_{12} & C_{12} & C_{11} \end{bmatrix} \begin{bmatrix} \varepsilon_{x} \\ \varepsilon_{y} \\ \varepsilon_{z} \end{bmatrix}$$

- All diagonal elements equal -> C11 - All off-diagonal elements equal -> C12

Biaxid strain: $\delta_{x} = \delta_{y} > 0$ (tensile) $\delta_{x} = \delta_{y} < 0$ (compressive)

Consider
$$\sigma_x = \sigma_y$$
, $\sigma_z = 0$ so the also have $\varepsilon_x = \varepsilon_y$

 Under biaxial strain, the deformation along 2 will be opposite in sign from the deformation along x, y
 (like squeezing effect)

Combining this with band structure models, we get:

$$\delta E_{c} = a_{c}(\xi_{x} + \xi_{y} + \xi_{z}) = a_{c}(2\xi_{xy} - \frac{2C_{12}}{C_{11}}\xi_{xy})$$

$$\delta E_{c} = 2a_{c}(1 - \frac{C_{12}}{C_{11}})\xi_{xy}$$

$$\alpha = a_{c} - a_{v}$$

$$= hydrostatic potential$$

$$Q_{\varepsilon} = -b(1 + 2\frac{C_{12}}{C_{11}})\xi_{xy}$$

$$\delta = shear potential$$

where the definitions are

• Assuming a direct bandgap, the band edges (at the minimum K=0) are:

$$E_{C} = E_{g}(x) + \delta E_{c}(x)$$

$$E_{HH} = -P_{\varepsilon}(x) - Q_{\varepsilon}(x)$$

$$E_{LH} = -P_{\varepsilon}(x) + Q_{\varepsilon}(x)$$

where energy is referenced to EHH at k=0 without strain.

- · Tensile vs. compressive strain
 - · Tensile: SEc < 0, PE < 0, QE > 0
 - * Compressive: $\&E_c>0$, $P_E>0$, $Q_E<0$

- The dispersion relations away from the center of K-space are given by

$$E_{HH}(k) = -P_{\varepsilon}(k) - Q_{\varepsilon}(x) - \frac{h^{2}}{2m_{hh}^{t}} k_{t}^{2} - \frac{h^{2}}{2m_{hh}^{z}} k_{z}^{2}$$

 $E_{LH}(k) = -P_{\varepsilon}(x) + Q_{\varepsilon}(x) - \frac{h^2}{2m_{th}^2} k_t^2 - \frac{h^2}{2m_{st}^2} k_{\varepsilon}^2$

(without valence band mixing)

transverse dispersion longitudinal dispersion

transpareny carrier density

- · Use mt to calculate the 2D DOS (which goes into the gain expression)
- " Use m2 to calculate the energy levels of the QW subbands

Effective masses: HH vs. LH

$$m_{hh}^{t} = \frac{m_o}{\gamma_1 + \gamma_2}$$
, $m_{ah}^{t} = \frac{m_o}{\gamma_1 - \gamma_2}$

 $m_{hh}^2 = \frac{m_0}{\gamma_1 - 2\gamma_2}$, $m_{sh}^2 = \frac{m_0}{\gamma_1 + 2\gamma_2}$

where Pi, Y2 > 0 are band structure parameters

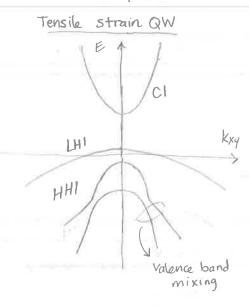
In a quantum well,

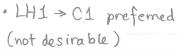
- The heavy hole has a lighter transverse effective mass! > HH → C transition is preferred to reduce N+r
- The light hole has a lighter longitudinal effective mass
- The values of the effective mass change with strain, however

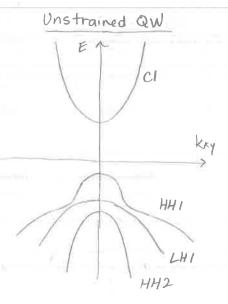
An unstrained quantum well seems to have a DOS advantage over a bulk laser:

- 1) HH has a lower DOS than LH in a QW
- 2) HH has a heavier m2, so it is closer to the conduction band than LH > HH > C transitions are already preferred over LH > C transitions

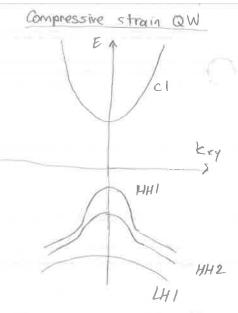
The benefit of strain







· HH1 -> C1 preferred



 HH1 → C1 further preferred (desirable)

- Compressive strain moves the LH bands further from the HH bands, leading to a lower valence band effective mass due to a stronger preference for the HH \rightarrow C transition
 - > lower Ntr, Jth
- A lower mh (more comparable to me) also brings the Fermi levels closer to the band edges on both sides. Since the Fermi occupancy function changes most rapidly at the Fermi level, this means that the carrier density becomes more sensitive to Fermi level shifts
 - > larger differential gain dg/dN

(this can be achieved with p-doping as well, but this increases 1/4r due to the extra holes that need to be supplied)

Polarization dependence

- · Compressive strain: most transitions are HH+C, so the gain is more responsive to TE polarization (also the case with unstrained, but less so)
- · Tensile strain: most transitions are LH+C, so the gain is more responsive to TM polarization; at moderate tensile strain, the gain is independent of polarization (desirable for optical amplifiers)
- * Compressive -> HH transitions -> TE

Tensile -> LH transitions -> TM

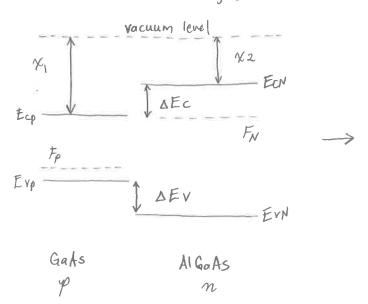
(7) Heterojunctions and the Double Heterostructure

- Attaining the transparency condition $\Delta F = Eg$ is difficult to achieve in practice in a normal ypm-diode, because the current becomes too large.

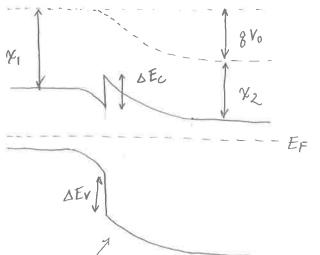
Today, most lasers use the double beterostructure

- Band alignment

No bias, before forming junction:

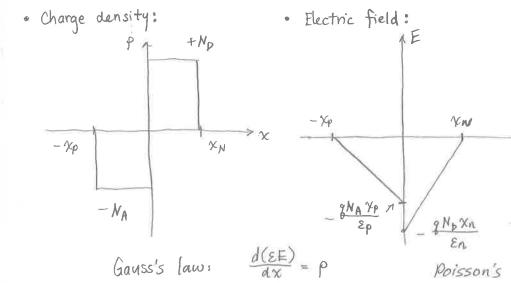


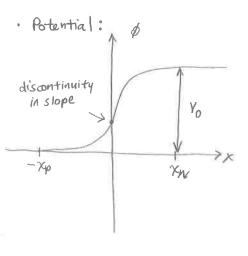
After contact, line up the Fermi levels:



Band bending: e transfers out of N-type in this part, so it bends upward

- Electrostatics





Poisson's eq:
$$\frac{dV}{dx} = -E(x)$$

- The built-in potential is:

$$\int_{Q} Y_{0} = F_{N} - F_{p} = E_{gp} + \Delta E_{c} - (E_{cN} - F_{N}) - (F_{p} - E_{Vp})$$

$$= E_{gp} + \Delta E_{c} - kT \ln \frac{N_{CN}}{N_{D}} - kT \ln \frac{N_{Vp}}{N_{A}}$$

$$qV_{0} = E_{gp} + \Delta E_{c} + kT \ln \frac{N_{A}N_{D}}{N_{CN}N_{VQ}}$$

from Fermi levels
before contact

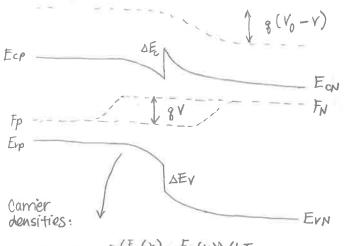
< 0 if non-degenerately doped
(if degenerately doped, this expression is invalid)</pre>

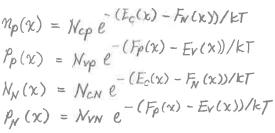
- The valence band $E_V(x)$ follows $-g\phi(x)$, but with a discontinuity of ΔE_V at the junction
- The conduction band $E_c(x) = E_r(x) + E_g(x)$; again, there is a disconfinuity at the junction, due to both ΔE_r and the change in Eg

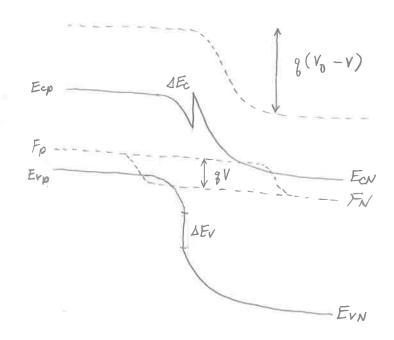
Biased heterojunction

Forward bias: V>0

Reverse bias: V <0







$$\Rightarrow M_{p}(x)P_{p}(x) = N_{cp} N_{vp} e^{(E_{v}(x)-E_{c}(x))/kT} (F_{w}(x)-F_{p}(x))/kT$$

$$= N_{cp} N_{vp} e^{-E_{g}(x)/kT} e^{gV/kT}$$

$$M_{p}(x)P_{p}(x) = n_{ip}^{2} e^{gV/kT}$$
The law of mass action still holds

- Double heterostructure laser basics:

· Communication wavelength:

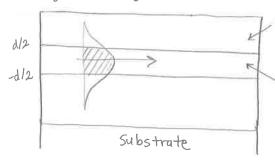
0.35eV

Absorption minimum in silica, optimal for transmission through optical fiber > In Ga As P is commonly used λ = 1.30 μm

Smallest dn/dx: least material dispersion

lattice constant

· Edge - emitting laser:



cladding: large Eg, small n

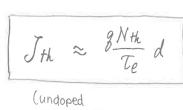
active region: small Eg, large n - At a given wavelength, approximate the gain as a linear function of the camer concentration:

Threshold gain (see § 1)

$$g_{th} = \frac{1}{\int (\alpha_i + \alpha_m)} = a(N_{th} - N_{tr})$$

$$\Rightarrow N_{th} = N_{tr} + \frac{1}{\int a(\alpha_i + \alpha_m)}$$

· Threshold current



approximation)

use confinement to reduce Ith 3d (Ntr + Ta (a; + am))

reduce Nor to lower threshold cument (e.g. by using a QW)

loss increases threshold cument

- The carrier lifetime Te is determined by how long it takes a carrier to

recombination
$$= \frac{N}{\text{Te}}$$
 (cm⁻³s⁻¹)

$$\frac{N}{Te} \approx AN + BN^2 + CN^3$$

recombination

Radiative recombination Current in a heterojunction

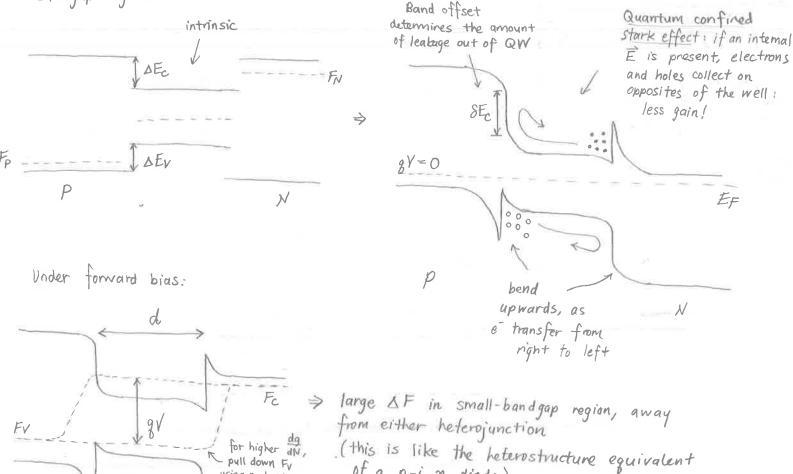
Al though the derivation is lengthy, we can show that the current across a heterojunction is given by:

$$J = g \left(\frac{p_n}{L_n} n_{po} + \frac{p_p}{L_p} p_{No} \right) \left(e^{gV/kT} - 1 \right)$$
electrons in p holes in N

$$n_{po} = n_{ip}^2 / N_A \qquad p_{No} = n_{iN}^2 / N_p$$

Double heterostructure

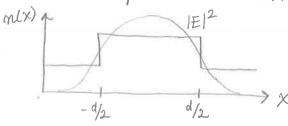
· Carrier confinement : high concentration of electrons and holes in a smallbandgap region



· Optical confirement: Structure is a slab waveguide along x

pull down Fy

using p-doping



E=0 in the well

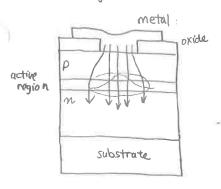
· confinement factor

of a p-i-n diode)

* To inchease I while using small QWs for current confinement, use multiple quantum wells (MQW)

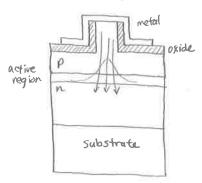
Common DH laser structures

Gain-quided laser



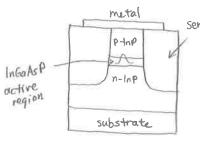
- · Gain guiding: the oxide layer ensures that only a small part of the active region provides gain
 - > weak guiding of the optical mode
- · Cument can spread laterally from oxide to reach the active region (weak current confirement)
 - > larger Ith
- Higher order side modes are easily generated (a bad thing!)

Ridge wareguide laser



- · Reduces lateral current spreading
 - > reduces Ith compared to gain guiding structure
- Provides weak lateral index guiding of the optical mode due to the ridge above the active region
 - Relatively small confirement factor I
 - > relatively high threshold gain needed > Ith T

Buried heterostructure laser

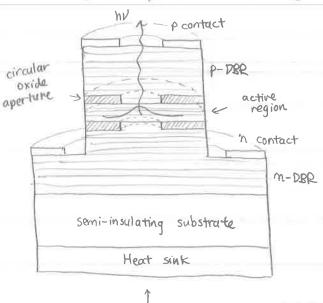


- Semi-in sulating
- · Improves current confinement by surrounding the active region with heterojunctions on all sides, including semi-insulating layers > Ith V
 - optical confinement with larger sn
 - Lower parasitic capacitance : improves bandwidth
- Optical cavities for edge-emitting lasers

Fe:InP

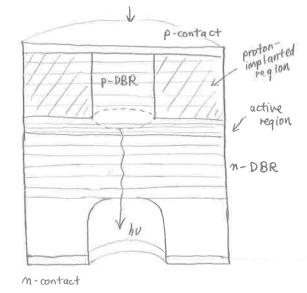
- * Fabry-Perot: two-mirror arrangement (see §(7))
 - Many longitudinal modes within gain spectrum >> limits bandwidth
- Distributed feedback (DFB) laser: mode selectivity by placing a diffractive grating directly above active region
 - Widely used for 1550 nm long-haul optical communications
- · Distributed Bragg reflector (DBR) laser: mode selectivity by introducing periodic index variation at the two ends of laser cavity
 - Like Fabry-Perot with wavelength selective dielectric mimors

Yertical cavity surface-emitting laser (VCSEL)



oxide VCSEL

Proton-implant YCSEL



- Emits out of the top or bottom surface, rather than the edges

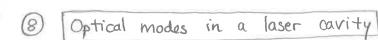
· No need for optically flat facets

. No need for cleaning /dicing ⇒ batch processing ⇒ much cheaper!

· Easily coupled to optical with a circular aperture

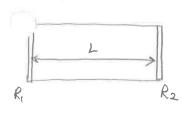
· Small cavity volume

- Low threshold current Ith
- Single mode operation, but large lire midth sh
- Cavity consists of an active region surrounded by DBR stacks with R > 99%
 - Not practical for 1.3-1-55 mm (long haul), as too many DBR pairs would be needed
- Current confinement can be provided in several ways
- · Oxide VCSEL: oxidized AlGaAs apertures provide both lateral cument confinement and index confinement of optical mode
- · Proton implantation: parts of the structure that are implanted with protons become insulating
 - Weaker optical confirement, relies on gain guiding
- The cavity length is usually $\sim \lambda/n$ (one wavelength) which results in a peak of the electric field at the center of the cavity between the DBRs
- The DBR layers can be made to have low resistivity by using graded, highly doped heterojunctions



· Longitudinal modes: modes along propagation direction

field profile



The electric field has the form $\vec{E}(x,y,z) = \vec{E}_0(x,y) e^{i\vec{k}_2z}$ Round trip condition gives the allowed kz:

$$e^{ik_z(2L)} = 1 \Rightarrow 2k_zL = 2\pi m$$

$$k_z = \frac{m\pi}{L}$$

m=1:

$$\frac{\lambda/n}{2} = L$$

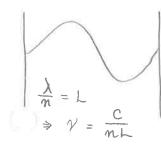
$$\Rightarrow V = \frac{c}{2nL}$$

- Use the dispersion relation for light to find the mode frequencies:

$$k_{z} = \frac{m\pi}{L} = \frac{n\omega}{c} = \frac{2\pi nv}{c}$$

$$V_{m} = m \frac{c}{2nL}$$

m=2:



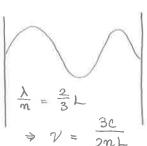
This is obvious simply from inspecting the cavity, as shown on left

- The mode spacing is

$$\Delta V = \frac{c}{2nL} \quad \text{or} \quad \sin \alpha \quad dV = -\frac{c}{\lambda^2} d\lambda$$

$$\Delta \lambda = -\frac{\chi^2}{2nL}$$

m = 3



- If we include material dispersion,

$$\Delta V = V_m - V_{m-1} = \frac{mc}{2m_m L} - \frac{(m-1)c}{2n_{m-1} L}$$

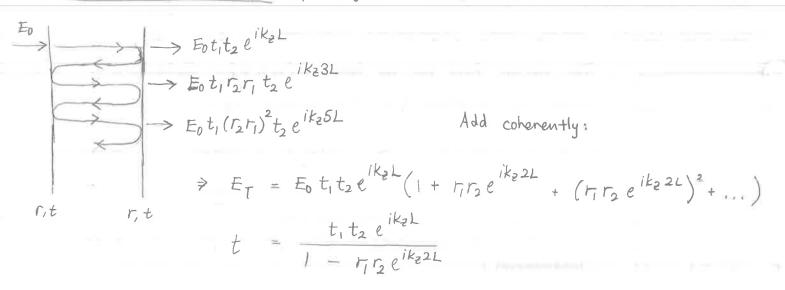
$$= \frac{c}{2n_m L} + \frac{(m-1)c}{2L} \left[\frac{1}{m_m} - \frac{1}{m_{m+1}} \right]$$

$$= \frac{c}{2n_k L} + V_{m-1} n_{m-1} \left[\frac{-\Delta n}{n^2} \frac{1}{\Delta V} \Delta V \right]$$

$$= \frac{c}{2n_k L} - \frac{V}{n} \frac{dn}{df} \Delta V$$

$$\Rightarrow \boxed{\Delta V = \frac{c}{2nL} \left(1 + \frac{v}{n} \frac{dn}{dv} \right)^{-1}} \quad \text{or} \quad \Delta \lambda = \frac{-\lambda^2}{2nL \left(1 - \frac{\lambda}{n} \frac{dn}{d\lambda} \right)}$$

without dispersion



If internal and external media have the same index,

$$T = |t|^{2} = \frac{|t_{1}t_{2}|^{2}}{|1 - r_{1}r_{2}e^{ik_{3}2L}|^{2}} = \frac{|t_{1}t_{2}|^{2}}{(1 - |r_{1}r_{2}|)^{2}} \frac{1}{1 + (\frac{2F}{\pi})^{2} \sin^{2}(k_{z}L)}$$

high reflectivity RIRZ

This gives: $\Delta V = \frac{c}{2nL}$ δV reflectivity R_{1}, R_{2}

$$\frac{|t_1 t_2|^2}{(1 - |r_1 r_2|)^2} \frac{1}{1 + (\frac{2F}{\pi})^2 \sin^2(k_z L)}$$

$$\frac{7}{1 - |r_1 r_2|} = \frac{\pi (R_1 R_2)^{1/4}}{1 - (R_1 R_2)^{1/2}}$$

· The firesse F describes the sharpress of each peak:

$$FWHM: \delta V = \frac{\Delta V}{F}$$

The quality factor is therefore:

$$Q_{FP} = \frac{\nu}{\delta \nu} = \frac{\nu}{\Delta \nu} F$$

$$Q_{FP} = \frac{2mL}{\lambda} \cdot \frac{\pi (R_1 R_2)^{V4}}{1 - (R_1 R_2)^{V2}}$$

Laser modes: overlap the gain spectrum onto Fabry Perot modes. The

M = 1 $\lambda = 1550 \text{ nm}$ L = 1 µm $R_1 = 0.95$ $R_2 = 0.995$ $\sqrt{}$ $Q \approx 144$

Tp & 1.2ps

amplification factor due to gain is:

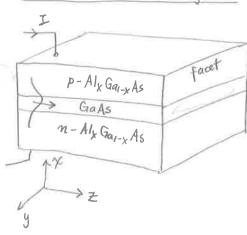
$$Amp = 1 + R_1 R_2 e^{2gL} + (R_1 R_2 e^{2gL})^2 + \dots = 1 - R_1 R_2 e^{2gL} \iff 1 - R_1 R_2 e^{2gL} \implies 1 - R_1 R_2 e^{2$$

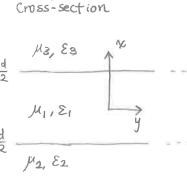
amplification factor gain spectrum g(L)

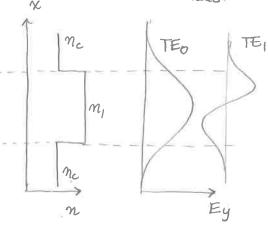
this value will always be clamped slightly below I at gain peak

lasing modes (threshold)

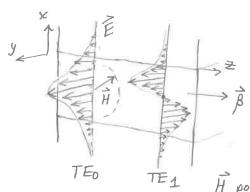




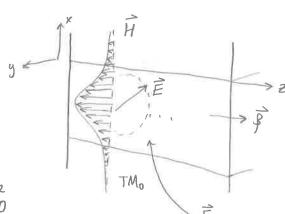




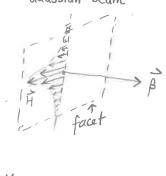
- TE modes:
$$\hat{E} = Ey \hat{y}$$



H points in the XZ plane, Hy = 0



Emerges as TM Polarized Gaussian beam



È points in the XZ plane, Ey = 0

Assume structure is extended infinitely along y: All $\frac{\partial}{\partial y} \rightarrow 0$.

Apply the electromagnetic wave equation

$$\vec{\nabla}^2 \vec{E} = -\omega^2 \mu \epsilon \vec{E} \Rightarrow (\vec{\nabla}^2 + \omega^2 \mu \epsilon) E_y = 0$$

$$(\frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial z^2} + \omega^2 \mu \epsilon) E_y = 0$$

Guess that the solution will be of the form Ey = $\phi(x)e^{ik_z} \neq e^{-i\omega t}$ and that further more,

$$\phi(\chi) = \begin{cases} C_g e^{-\alpha_3}(\chi - \frac{d}{2}) & \chi > \frac{d}{2} \\ C_1 \cos(k_\chi \chi) + D_1 \sin(k_\chi \chi) & -\frac{d}{2} < \chi < \frac{d}{2} \\ C_2 e^{\alpha_2}(\chi + \frac{d}{2}) & \chi < -\frac{d}{2} \end{cases}$$

C, corresponds to even modes: TEO, TE2, ...

Di corresponds to odd modes: TEI, TE3, ...

These can be written as plane waves, so Ey satisfies the wave equation

- Now apply boundary conditions at $x = \pm \frac{d}{2}$

1) Ey must be continuous (transverse to interface)

2) Hx and Hz must be continuous

Use Faraday's law to get \vec{H} $\vec{\nabla} \times \vec{E} = -\mu \frac{\partial \vec{H}}{\partial t}$

$$\begin{vmatrix} \hat{x} & \hat{y} & \hat{z} \\ \frac{\partial}{\partial x} & \frac{\partial}{\partial y} & \frac{\partial}{\partial z} \end{vmatrix} = -\frac{\partial E_{y}}{\partial z} \hat{x} + \frac{\partial E_{y}}{\partial x} \hat{z} = + \mu i \omega \hat{H}$$

$$|\hat{H}| = \frac{1}{i \omega \mu} \left(-i k_{z} E_{y} \hat{x} + \frac{\partial E_{y}}{\partial x} \hat{z} \right)$$

$$|\hat{H}| = -\frac{k_{z}}{\omega \mu} E_{y} + \frac{1}{2} \frac{\partial E_{y}}{\partial x} \hat{z}$$

$$|\hat{H}| = -\frac{k_{z}}{\omega \mu} E_{y} + \frac{1}{2} \frac{\partial E_{y}}{\partial x} \hat{z}$$

* H_X is automatically continuous if Ey is continuous, assuming $\mu = \mu_0$ for all materials

Hz is continuous if 2Ey is continuous

Thus:

At
$$x = -\frac{d}{2}$$
; $C_2 = C_1 \cos\left(-\frac{k_X d}{2}\right) + D_1 \sin\left(-\frac{k_X d}{2}\right)$

$$\alpha_2 C_2 = -k_X C_1 \sin\left(-\frac{k_X d}{2}\right) + k_X D_1 \cos\left(-\frac{k_X d}{2}\right)$$
2

At
$$\chi = +\frac{d}{2}$$
: $C_3 = C_1 \cos\left(\frac{k_x d}{2}\right) + D_1 \sin\left(\frac{k_x d}{2}\right)$

$$-\alpha_3 C_3 = -k_x C_1 \sin\left(\frac{k_x d}{2}\right) + k_x D_1 \cos\left(\frac{k_x d}{2}\right) \qquad (4)$$

We have also 3 additional equations from the wave equation: $(\frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial z^2} + \omega^2_{ME})E_y = 0$

Region 3:
$$-\alpha_3^2 + k_z^2 = \omega^2 \mu_0 \, \epsilon_3$$

Region 1:
$$k_x^2 + k_z^2 = w^2 \mu_0 \epsilon_1$$
 5

Region 2:
$$-x_2^2 + k_2^2 = w^2 \mu_0 \epsilon_2$$

$$\mathcal{E}_2 = \mathcal{E}_3 = \mathcal{E}_c \implies \alpha_2 = \alpha_3 = \alpha$$

$$\Rightarrow -\alpha^2 + k_z^2 = \omega^2 \mu_0 \, \varepsilon_c \quad \textcircled{6}$$

- Even modes: symmetric, C2 = C3

① + ③ :
$$C_1 + C_3 = 2C_2 = 2C_1 \cos(\frac{k_x d}{2}) \Rightarrow C_2 = C_1 \cos(\frac{k_x d}{2})$$

②-④:
$$\alpha_2 C_2 + \alpha_3 C_3 = 2\alpha C_2 = 2k_x C_1 \sin\left(\frac{k_x d}{2}\right) \Rightarrow C_2 = \frac{k_x}{\alpha} \sin\left(\frac{k_x d}{2}\right)$$

Combining these, we get: $\alpha = k_x \tan(k_x \frac{d}{2})$ Even modes

- Odd modes: assymetric: C2 = - C3

$$0 - 3 + C_2 - C_3 = 2C_2 = -2D_1 \sin\left(\frac{k \times d}{2}\right) \Rightarrow C_2 = -D_1 \sin\left(\frac{k \times d}{2}\right)$$

$$\Rightarrow \qquad \boxed{\alpha = -k_x \cot(k_x \frac{d}{2})} \quad \text{odd modes}$$

How do we interpret these characteristic equations?

If a solution exists, a mode exists. Otherwise, the mode does not exist. If multiple solutions exist, multiple modes can propagate.

Let
$$X = k_X \frac{d}{2}$$
, $Y = \alpha \frac{d}{2}$

(5) - (6):
$$K_x^2 + \alpha^2 = \omega^2 \mu_0 (\xi_1 - \xi_c)$$

$$\left(\frac{k_{x}d}{2}\right)^{2} + \left(\frac{xd}{2}\right)^{2} = \left(\frac{d}{2}\right)^{2} w_{M_{0}}^{2} \left(\xi_{1} - \xi_{c}\right)$$

$$\chi^{2} + \chi^{2} = \left(\frac{d}{2}\right)^{2} w^{2} \mu_{0} \varepsilon_{0} \left(\eta_{1}^{2} - \eta_{c}^{2}\right)$$

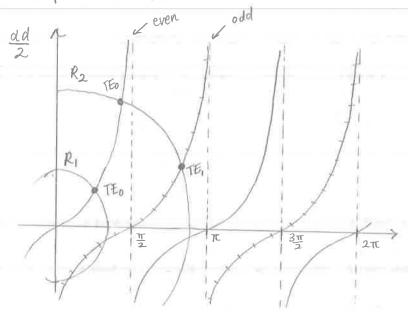
Since RHS is constant (not dependent on k_X or α), this is an equation for a circle in the X-Y plane

Meanwhile, the characteristic equations can be mitten as:

$$Y = X \tan X$$
, $Y = -X \cot X$
(even) (odd)

Plotting the circle together with these expressions gives us a way to find the modes

- Graphical solution:



The two circles correspond to different values of $R^2 = \left(\frac{d}{z}\right)^2 \omega^2 \mu_0 \varepsilon_0 \left(m_i^2 - m_c^2\right)$

- · As R increases, more modes can exist
- · To increase R,
 - Increase the size, d
 - Increase the index contrast n2-n2

Kxd 2

The first even mode always exists. The condition for having a single waveguide mode is

$$R < \frac{\pi}{2} \Rightarrow \frac{d}{2} \omega \sqrt{\mu_0 \, \epsilon_0} \sqrt{m_i^2 - m_c^2} = \frac{d}{2} \frac{\omega}{c} \sqrt{m_i^2 - m_c^2} = \frac{d}{2} k_0 \sqrt{m_i^2 - m_c^2} < \frac{\pi}{2}$$

$$k_0 \frac{d}{2} \sqrt{m_1^2 - m_c^2} < \frac{\pi}{2}$$

Another way to write this is:

$$d\sqrt{m_1^2-m_c^2}<\frac{\lambda_0}{2}$$

small index contrast

$$d^{2}(\eta_{l}^{2} - \eta_{c}^{2}) < \left(\frac{\lambda_{0}}{2}\right)^{2}$$

$$\approx d^{2}(2\eta_{l}) \Delta \eta < \frac{\lambda_{0}^{2}}{4}$$

$$\Rightarrow \left| d^2 \Delta n < \frac{\lambda_0^2}{8m_1} \right|$$

For
$$\lambda = 873$$
 nm
 $M_1 \approx 3.59$ (GaAs)
 $M_2 \approx 3.385$ (Alo.3 Gao.7 As)
 $\Rightarrow d < 0.36$ μ m

Solve for Kz

(5)
$$+ k_z^2 = w^2 \mu_0 \mathcal{E}_1 - k_x^2$$
, (6) $+ k_z^2 = w^2 \mu_0 \mathcal{E}_c + \alpha^2$

- when a mode is at the cutoff of being excited, $\alpha=0 \Rightarrow k_z=n_c k_0$ In the limit of high frequency, kx asymptotes, so $w^2 \mu_0 \, \epsilon_1 >> k_x^2 \Rightarrow k_z=n_1 k_0$ At intermediate frequencies, $n_c k_0 < k_z < n_1 k_0$
- · At the mode's cutoff d=0 (not confined) and & increases with frequency

- Dispersion curves

Mode profile: (TED)

barely confired

KS = WIKO KZ Kz = Mcko 120 to the > Ko ~ W cutoff 1 cutoff 2

Low frequency limit High frequency limit

well confined

of large

The effective index of a mode is defined as Meff = $\frac{kz}{kn}$ For the reasons stated above, every mode starts at Mc and ends at m, as a function of w

= It is customary to define the following 3 parameters:

Normalized frequency
$$V = k_0 d \sqrt{m_1^2 - m_2^2}$$
 frequency $V = k_0 d \sqrt{m_1^2 - m_2^2}$ $\rightarrow 0$ at low V (cutoff = π), 1 at high V Asymmetry $\alpha = \frac{m_1^2 - m_2^2}{m_1^2 - m_2^2}$ $\rightarrow \alpha = 0$ for symmetric $s(ab)$

Re-express single-mode condition: $V < \pi$

Power normalization

$$P = \frac{1}{2} \operatorname{Re} \int_{-\infty}^{\infty} (\vec{E} \times \vec{H}^*) \cdot \hat{z} \, dx = 1 \Rightarrow -\frac{1}{2} \operatorname{Re} \int_{-\infty}^{\infty} \operatorname{Ey} H_{x}^* \, dx = \frac{1}{2} \int_{-\infty}^{\infty} \frac{k_{z}}{w_{\mu}} |\operatorname{Ey}|^{2} dx = 1$$

Optical confinement factor

$$\mathcal{J} = \frac{\int_{-d/2}^{d/2} |\vec{E}(\chi)|^2 d\chi}{\int_{-\infty}^{\infty} |\vec{E}(\chi)|^2 d\chi}$$

This last condition can be used to find the remaining unknown in Specifying Ey(x, 2)

* 57 is small at mode cutoff, and becomes larger with increasing ko or d

> Therefore, for single mode operation, it is desirable to be just below the cutoff of the first excited mode (i.e. near the single mode condition)

TM modes: to solve, use the duality principle and make the following

replacements,

$$\overrightarrow{E} \rightarrow \overrightarrow{H}$$
 $\overrightarrow{H} \rightarrow -\overrightarrow{E}$
and
 $\varepsilon \rightarrow \mu$

TE
TM

The Characteristic equations are $\alpha \frac{d}{2} = \frac{\mathcal{E}_c}{\mathcal{E}_1} \left(\frac{k_x d}{2} \right) \tan \left(\frac{k_x d}{2} \right) \leftarrow \text{even}$ $d\frac{d}{2} = -\frac{\varepsilon}{\varepsilon_1} \left(\frac{k_x d}{z} \right) \cot \left(\frac{k_x d}{z} \right) \leftarrow odd$

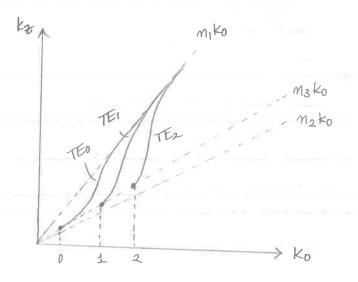
Asymmetric waveguide: n2>n3

· The cutoff condition changes to

where a is the asymmetry parameter

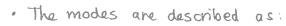
· Notice that now even the TEO mode has a finite cutoff frequency! The Single mode condition changes to:

$$tan^{-1}\sqrt{\alpha} \leq V \leq tan^{-1}\sqrt{\alpha} + \pi$$



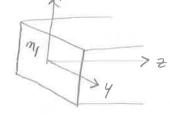
· Now the low frequency limit of Meff is the index of the higher-index cladding

Rectangular waveguide modes

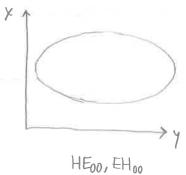


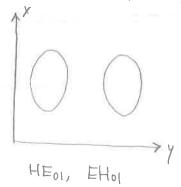
- HEpq: Hx and Ey dominant (like TE for infinite slab)

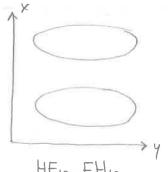
- EHpg: Ex and Hy dominant (like TM for infinite slab)



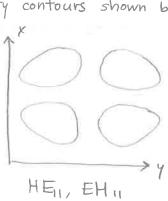
- P, & denote the mode frequency along x and y: intensity contours shown below



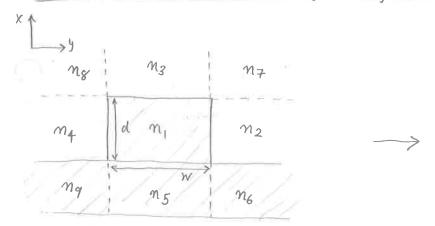




HEIO, EHIO

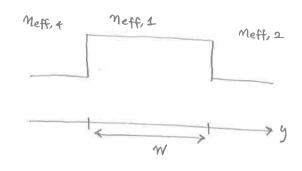


Effective index method for ridge waveguides



- D Solve the asymmetric slab waveguide $M_3 M_1 M_5$, treating it as infinite along y. Find Meff, 1 for this waveguide
- 2) If $m_2 > m_7$, m_6 , use the same method to find m_2 , eff. Otherwise, approximate weak fields in 6, 7 and let m_2 , eff = m_2
- 3 Follow the same procedure as @ for the no-no-no waveguide to find noteff.

Field profiles F(x, y) where the field in y is stitched together from the 3 parts



- A Solve the infinite slab waveguide above to find the field profile G(y) along y.
- © Combine the results to get the overall field E(x,y) = F(x,y)G(y)

$$\left(\frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2}\right)A - i2k\frac{\partial A}{\partial z} = 0$$

The Gaussian beam is a solution to this equation, and is given by

$$A(F) = \frac{A_1}{g(z)} e^{-ik\left(\frac{x^2+y^2}{2g(z)}\right)}$$

Gaussian beam envelope

where
$$g(z) = z + iz_0$$

also written as

$$\frac{1}{g(z)} = \frac{1}{R(z)} - i \frac{\lambda}{\pi W^2(z)}$$

[·] Gaussian beam: the most spatially localized type of wave that can exist in free space

⁻ Origin of the Gaussian beam: consider a paraxial wave $V(\vec{r}) = A(\vec{r}) e^{-ikZ}$ where A is an envelope that varies much less slowly than K. For U to satisfy the Helmholtz equation, A must satisfy:

$$U(\vec{r}) = A_0 \frac{W_0}{W(z)} e^{-\frac{x^2 + y^2}{W^2(z)}} - ikz - ik \frac{x^2 + y^2}{2R(z)} + i\vec{z}(z)$$

where
$$W(z) = W_0 \left[1 + \left(\frac{z}{z_0}\right)^2\right]^{1/2}$$

 $R(z) = z \left[1 + \left(\frac{z_0}{z}\right)^2\right]$
 $\tilde{c}(z) = \tan^{-1} \frac{z}{z_0}$

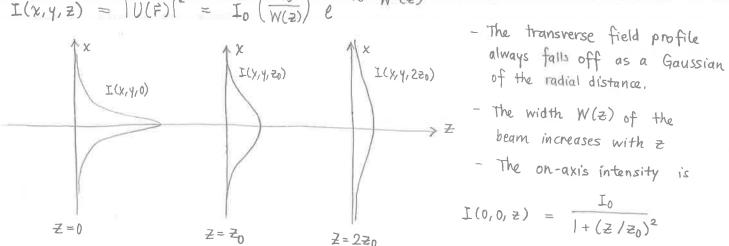
This is the excess delay of $W_0 = \sqrt{\frac{\lambda Z_0}{\pi}}$ a Gaussian beam in companison with a plane wave

Ao and to are determined by boundary conditions; everything can be found

· Properties of the Gaussian beam

1) Intensity

$$I(x,y,z) = |U(\vec{r})|^2 = I_0 \left(\frac{W_0}{W(z)}\right)^2 e^{-2 \frac{x^2 + y^2}{W^2(z)}}$$



- The width W(z) of the beam increases with z
 - The on-axis intensity is

$$I(0,0,2) = \frac{I_0}{1 + (2/2_0)^2}$$

2) The power is

$$P = \int_0^\infty L(x, y, z) dx dy = \frac{1}{2} I_0 \pi W_0^2$$

$$\uparrow \qquad \uparrow$$
peak beam intensity area

At
$$z = z_0$$
, $I = \frac{1}{2}I_0$
As $z \to \infty$, $I = I_0 z_0^2/z^2$

(inverse square law, like a Spherical wave)

We can remite the intensity as

$$L(x,y,z) = \frac{2P}{\pi W^2(z)} e^{-2 \frac{x^2 + y^2}{W^2(z)}}$$

- About 86% of the power is within a circle of radius $\rho = \sqrt{\chi^2 + y^2} = W(z)$
- * About 99% is within 1.5W(Z)

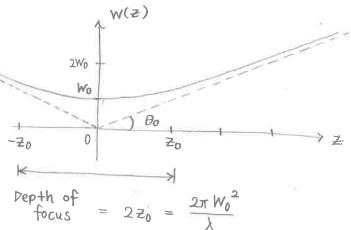
3) Beam radius:
$$W(z) = W_0 \left[1 + \left(\frac{z}{z_0}\right)^2\right]^{1/2}$$

The field drops to e^{-2} at this radius (≈ 0.135)

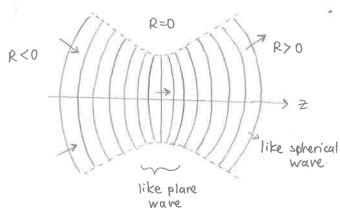
- At z=0, W is at its minimum and is equal to Wo This plane is called the waist and Wo is the waist radius.
- As $\neq \to \infty$, $W(\neq) \approx \frac{W_0}{Z_0} \neq \theta_0 \neq 0$ where the beam divergence is:

$$\theta_0 = \frac{2}{\pi} \frac{\lambda}{2W_0}$$

- Shorter wavelengths diverge less
- · A wider beam diverges less
- · Beams with a greater depth of focus diverge less



4) Wavefronts: due to the phase term $e^{-ik(x^2+y^2)/2R(z)}$ the wavefronts have

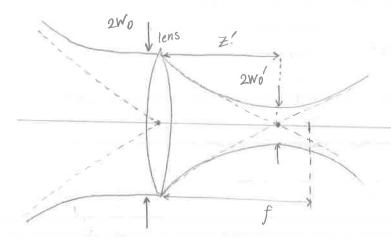


- R(Z) represents the radius of curvature of the wavefront at Z
 - At z=0, R→∞ > planar wavefronts
 - As z > 00, R > z => spherical wavefronts

- The Gaussian beam needs to be specified by 4 parameters:
 - 1) Amplitude Ao
 - 2) Direction, or beam axis
 - 3) Location of waist (z=0)
 - 4) Waist radius Wo or Rayleigh range Zo

Manipulation of Gaussian beams by passive optics

- * A thin lens retains the shape of the Gaussian, but magnifies its waist radius: $W_0' = MW_0$, $\theta_0' = \frac{1}{M}\theta_0$
- * Focusing a Gaussian beam: what is the smallest possible spot size?



- Put a lens directly at the waist of a Gaussian. The new Gaussian has its waist at Z' and has a new diameter 2Wo

$$W_0' = \frac{W_0}{\sqrt{1 + (z_0/f)^2}}$$

$$\Xi' = \frac{f}{1 + (f/z_0)^2}$$

To make the smallest possible spot, use a beam with a large waist Wo and therefore large depth of focus 220. Since this is approximately a plane wave, the lens focuses to a distance f. The waist diameter is:

$$2W_0' \approx \frac{4}{\pi} \lambda F_{\#} = \frac{4}{\pi} \lambda \frac{f}{D} \approx 1.27 \frac{\lambda}{2 \sin \theta}$$

[lens f-number]

D

This roughly corresponds to the diffraction limit,

$$\tan \theta \approx \sin \theta = \frac{b}{2f}$$

$$\frac{f}{D} \approx \frac{1}{2 \sin \theta}$$

- * Passive optics always preserves the quantity $W_0\cdot\theta_0$, which depends only on the wavelength λ . This is essentially a statement of étendue conservation.
- A Bessel beam is given by

$$\mathcal{D}(\vec{F}) = Am Jm(k_T \rho) e^{im\phi} e^{-ik_Z t} \quad \text{where} \quad m = 0, \pm 1, \pm 2, \dots$$

$$\rho = \sqrt{\chi^2 + y^2}, \quad \chi = \rho \cos \phi, \quad y = \rho \sin \phi$$

$$Jm = \text{Bessel function of } 1^{3t} \text{ kind, } m^{th} \text{ order}$$

A Bessel beam has zero beam divergence but infinite rms beam width

Dynamic response of semiconductor lasers The rate equations express the rate of change of N= carrier density and S = photon density camer loss injection rate of rate by carrier loss rate carriers into the recombination by stimulated emission active region ni = IQE = # photons produced # carriers injected $\frac{\gamma_i I}{qV} - \frac{N}{\tau(N)} - \gamma_g g(N) S$ V = active region volume $J = optical confinement factor = \frac{V}{V_P}$ where Vp = photon mode volume $\frac{dS}{dt} = \int V_g g(N)S - \frac{S}{Tp} + \int \beta R_{Sp}$ Rsp = spontaneous emission rate/vol. Vg = group velocity c/nr Tp = photon lifetime Stimulated emission loss rate spontaneous emission rate of photons T = carrier lifetime of photons rate of photons into the lasing mode Me = photon extraction efficiency $\frac{N}{U(N)} = AN + BN^2 + CN^3$, $R_{sp} = BN^2$ fraction of spontaneously emitted photons in the same direction as stimulated emission 12/4 € 10-3-10-4 Graphically: I/g current $(1-\eta_i)\frac{I}{g}$ spontaneous emission (1-B) Rsp V Rsp V RnrV = (AN + CN3)V Light output (incoherent) heat (camier reservoir spontaneous emission into lasing mode Absorption stimulated RIZV R21 V Me To Np Vp Light output (coherent) photon reservoir (1- Te) Mp Yp To (1) Wrong mirror transmission (2) free carrier absorption -> Heat

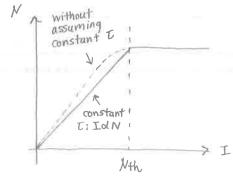
(3) absorption outside active region

• Steady state solutions:
$$\frac{d}{dt} = 0$$

$$\frac{dS}{dt} = \left(\int v_g g(N) - \frac{1}{Tp}\right) S = 0$$

$$\Rightarrow g(N) = \frac{1}{\int v_g Tp} = \frac{\alpha_{total}}{\int r} = g_{th} = g(N_{th})$$

- · Due to photon loss, the gain will be clamped after the laser reaches threshold: otherwise, the field in the cavity increases without bound > no steady state!
 - As a consequence, the carrier density is clamped at Nth above threshold, since the gain is not allowed to increase



- If the current is increased, N will exceed Nth briefly, but all of the excess carrier density will quickly be used up via increased stimulated emission, until a new steady state is reached

$$\frac{dN}{dt} = \frac{\eta_i I}{gV} - \frac{N}{T} = 0 \Rightarrow I = \frac{gV}{\eta_i T} N$$

- Above threshold,
$$N = Nth$$
, $g = gth$

$$\frac{dN}{dt} = \frac{\gamma_i I}{g V} - \frac{N + h}{U} - V_g g + h S = 0$$

$$S = \frac{1}{V_g g_{th}} \left(\frac{\eta_i I}{g V} - \frac{N_{th}}{I} \right) = \frac{1}{V_g g_{th}} \left(\frac{\eta_i I}{g V} - \frac{\eta_i I_{th}}{g V} \right)$$

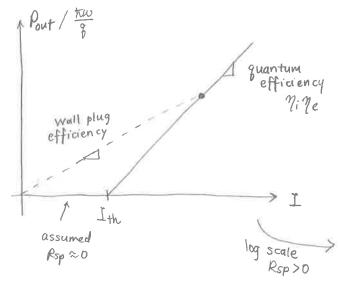
$$S = \frac{1}{V_g g_{th}} \frac{\eta_i}{g V} \left(I - I_{th} \right)$$

The power out is:

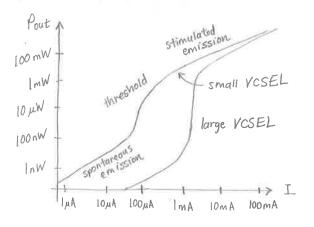
$$P_{out} = V_p S \frac{\alpha_m}{\alpha_m + \alpha_i} \frac{1}{T_p} \hbar \omega = \frac{SV}{T} \frac{\eta_e}{T_p} \hbar \omega = \frac{1}{V_g g_{th}} \frac{\eta_i}{g_{tr}} (I - I_{th}) \frac{V}{T} \frac{\eta_e}{T_p} \hbar \omega$$
extraction efficiency η_e

$$\Rightarrow P_{out} = \frac{\hbar \omega}{g} \eta_i \eta_e (I - I_{th})$$

photon energy / quantum to charge conversion efficiency



. The light output increases linearly with current above threshold under the approximation that Rsp & O



Detailed steady-state analysis: Rsp 70

- · Continue to assume T is constant (this implies that the SRH term = AN dominates the recombination rate)
- · Apply the rate equations:

$$\frac{dS}{dt} = \int V_g g(N) S - \frac{S}{Tp} + \int \beta R_{Sp} = 0$$

$$S(N) \left[\frac{1}{Tp} - \int V_g g(N) \right] = \int \beta R_{Sp}$$

$$D S(N) = \frac{\int \beta R_{Sp}(N)}{\frac{1}{Tp} - \int V_g g(N)}$$

$$\frac{dN}{dt} = \frac{\eta_i I}{gV} - \frac{N}{\tau} - v_g g(N) S(N) = 0$$

- · Below threshold: $S \approx 0$, so from ② we again have $I(N) = \frac{qV}{\eta_i} \frac{N}{T(N)}$
- · Above threshold: recall to = I'vg 9th

$$S(N) = \frac{\beta R_{SP}(N)/\nu_g}{g_{th} - g(N)}$$

- If the photon density stays finite, the threshold gain cannot actually be reached!
- g and N stay clamped just slightly below 9th and Nth

$$S(N) = \frac{\beta Rsp(N)/\gamma_g}{gth - g(N)}$$
 the sportaneous emission βRsp is clearly being amplified!

Use 2 to solve for current:

$$I(N) = \frac{qV}{\eta_i} \frac{N}{t(N)} + \frac{qV}{\eta_i} v_g g(N) S(N)$$

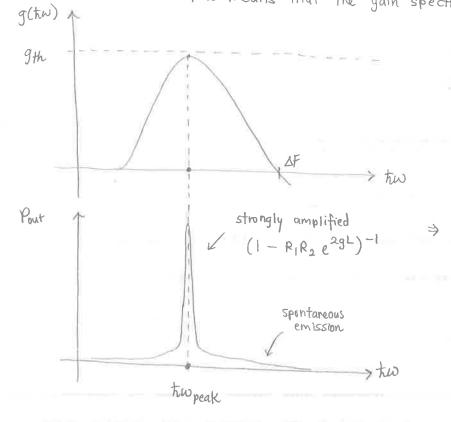
$$\approx I_{th}$$

$$I(N) = I + \frac{gV}{m_i} V_g g(N) \frac{\beta R_{SP}(N)/V_g}{g + h - g(N)}$$

$$\frac{gV}{\eta_i}$$
 $\beta Rsp(N) \frac{g(N)}{g_{th} - g(N)} = I(N) - Ith$

$$\Rightarrow g(N) = g_{th} \cdot \left(\frac{1}{g_{N\beta} R_{sp(N)/m_i} + 1} \right) < g_{th}$$

- * This shows again that the gain never reaches 9th (and same with N) but asymptotically approaches it with increasing current.
- This means that the gain spectrum appears as below



Independent of the gain physics, g is clamped at 9th due to the loss rate of photons from the cavity

Ith

I

Therefore, even without considering the allowable longitudinal lasing modes, it is a good approximation to assume a narrow emission spectrum with

· Differential analysis

- To analyze the modulation characteristics of semiconductor lasers, we need to linearize the rate equations around a steady state solution and treat the modulation as a small signal
- Take the differential of the rate equations:

$$d\left(\frac{dN}{dt}\right) = d\left(\frac{\eta_i I}{gV} - \frac{N}{I(N)} - V_g g(N)S\right)$$

where the differential carrier lifetimes are found from

$$d\left(\frac{N}{\tau}\right) = dN \frac{d}{dN}\left(\frac{N}{\tau}\right) = dN\left(A + 2BN + 3CN^{2}\right)$$

likewise,
$$\frac{1}{T_{\Delta N}} = \frac{d}{dN} (\beta R_{SP}) = \frac{d}{dN} (\beta BN^2)$$

$$\Rightarrow \frac{1}{T_{\Delta N}} = 2\beta BN + \frac{dB}{dN} BN^2$$

* The differential gain dg is found by differentiating g(N,S):

$$dg = \frac{\partial q}{\partial N} dN + \frac{\partial q}{\partial S} dS = \alpha dN - \alpha_p dS$$

where the differential gain values are defined as

$$\alpha = \frac{\partial g}{\partial N}$$
, $ap = -\frac{\partial g}{\partial S}$

Although a and ap are not constant with N and S, we can set them to their values under a steady state bias in a differential analysis.

To know their actual dependences, we need a model for g(N, s). From laser physics, we can approximate the gain as:

$$g(N,S) = \frac{g_0}{1 + \varepsilon S} \ln \left(\frac{N}{N_{+r}} \right) \Rightarrow a_p = -\frac{g_0}{2S} = \frac{\varepsilon}{1 + \varepsilon S} \cdot g$$

$$g_{ain}$$
Saturation

If we cast 3 and 4 into matrix form and do some algebra, we can write them as:

$$\frac{d}{dt} \begin{bmatrix} dN \\ dS \end{bmatrix} = \begin{bmatrix} -\gamma_{NN} & -\gamma_{NS} \\ \gamma_{SN} & -\gamma_{SS} \end{bmatrix} \begin{bmatrix} dN \\ dS \end{bmatrix} + \frac{\eta_i}{\gamma_{SN}} \begin{bmatrix} dI \\ 0 \end{bmatrix}$$

where
$$\gamma_{NN} = \frac{1}{T_{\Delta N}} + \gamma_g a S$$

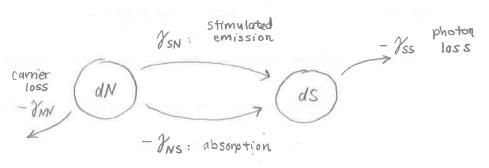
$$\gamma_{NS} = \gamma_g g - a_p \gamma_g S$$

$$\gamma_{SN} = \Gamma_{\gamma_g} a S$$

$$\gamma_{SS} = \Gamma_{\gamma_g} a S$$

where we can approximate the steady state gain as

Interpretation:

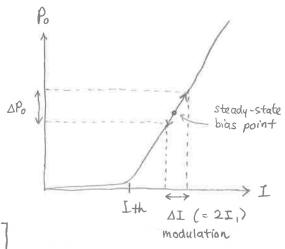


Small signal intensity modulation response

Let
$$dI(t) = I_1 e^{i\omega t}$$
 so $\frac{d}{dt} \rightarrow i\omega$
 $dN(t) = N_1 e^{i\omega t}$
 $dS(t) = S_1 e^{i\omega t}$

The differential rate equations become

$$\begin{bmatrix} \gamma_{NN} + i\omega & \gamma_{NS} \\ -\gamma_{SN} & \gamma_{SS} + i\omega \end{bmatrix} \begin{bmatrix} N_1 \\ S_1 \end{bmatrix} = \frac{\gamma_1 I_1}{8 V} \begin{bmatrix} 1 \\ 0 \end{bmatrix}$$



We solve this by first determining the determinant of the matrix on the LHS, then applying Cramer's rule to solve for NI, SI.

determinant
$$\Delta = (\gamma_{NS} \gamma_{SN} + \gamma_{NN} \gamma_{SS}) - \omega^2 + i\omega (\gamma_{NN} + \gamma_{SS})$$

= $\omega_R^2 - \omega^2 + i\omega \gamma$

where
$$W_R^2 = \gamma_{NS}\gamma_{SN} + \gamma_{NN}\gamma_{SS} \approx \frac{\gamma_g aS}{Tp}$$
 = relaxation resonance frequency and $\gamma = \gamma_{NN} + \gamma_{SS} \approx \kappa_R^2 + \gamma_0$ = damping

where
$$K = \text{Tp} \left[1 + J \frac{\alpha p}{\alpha} \right]$$

$$w_R^2 = \frac{v_g a}{\tau_p} S = \frac{v_g a}{\tau_p} \left(\frac{1}{v_g g_{th}} \frac{\eta_i}{q_V} (I - I_{th}) \right) = \frac{\int v_g a}{V} \frac{\eta_i}{q} (I - I_{th})$$

⇒
$$w_e^2 = \frac{\Gamma \gamma_g a}{g V} \gamma_i (I - I + h)$$

when the photons interact with the carriers. Consider the response to a sudden change in input cument;

N increases:
$$\frac{n_{i}T}{gV}$$
, gain increases

S then increases, due to increased gain: $JVgg(N)S$

N decreases, through increased stimulated emission: $-v_ggS$

S decreases, due to decreased gain

N increases, due to decreased stimulated emission

We describes the natural frequency of this oscillation

Frequency response

Now if we use Cramer's rule to solve the rate equations, we find:

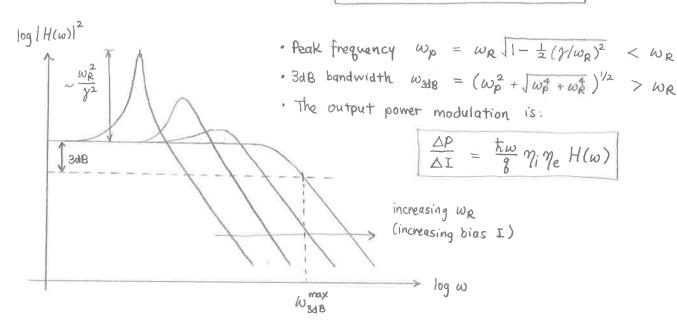
$$N_1 = \frac{\gamma_1 I_1}{g V} \frac{\gamma_{SS} + i \omega}{\omega_R^2} H(\omega)$$

$$S_1 = \frac{\gamma_i I_1}{gV} \frac{\gamma_{SN}}{\omega_R^2} H(\omega)$$

where the modulation transfer function is given by:

steady-state expression for s

$$H(\omega) = \frac{\omega_{R}^{2}}{\omega_{R}^{2} - \omega^{2} + i\omega\gamma}$$



Some notes on frequency response

- · Beyond the resonance peak, which occurs at $wp \approx wR$, the response falls off rapidly, at 40 dB/decade
- · The damping Y x KwR + Yo, increases with the bandwidth wR
 - At low power (small wp), the damping is small and the resonance peak is strong; at high power the resonance flattens out due to damping
- - * The modulation bandwidth increases as we drive the laser with more power I > Ith
 - · The bandwidth benefits directly from:
 - 1) Large differential gain $a = \frac{dg}{dN} \Rightarrow use a strained QW laser!$
 - 2) Large mode confinement factor J > use multiple QWs
 - 3) Small active region volume V
 - 4) Injection quantum efficiency 7; > 100%
 - The limit to w_R is ultimately set by $a = \frac{dg}{dN} = \frac{a_0}{1 + \varepsilon S}$
 - * As the bias I is increased, we increases until the photon density S is $\sim 1/\epsilon$. At that point a begins to decrease quickly as a result of gain compression
 - However, in practice such high photon densities are not usually reached. The limit is then set by the damping γ , which increases as ω_R increases. If γ increases too much, ω_{3dB} will fall below ω_R . The maximum possible 3dB bandwidth occurs at an optimal damping, and is found to be:

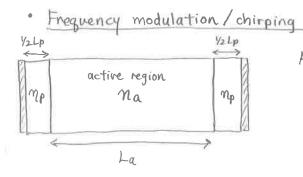
high power limit
$$\rightarrow W_{3d8}^{max} = \frac{\sqrt{2}}{2\pi K} = \frac{\sqrt{2}}{2\pi} \frac{1}{1 + \sqrt{3}ap/a} \times \frac{1}{\sqrt{Tp}} = W_{3d8}^{max}$$

This value represents a limit to the intrinsic modulation capabilities of the laser. It is not surprising that its ultimate limit is $\sim 1/\tau_p$

- Thermal management, mirror damage, etc. are also potential risks associated with high power operation

· Transient intensity modulation response

- In response to a sharp step increase in input current, the system response depends on the amount of damping
 - · If underdamped ($\gamma \ll \omega_R$), the carrier and photon densities will oscillate at the relaxation resonance frequency ω_R
 - The photon density (and power output) will oscillate and settle to the new, higher value
 - The carrier density will oscillate and settle back to Nth
 - · If the damping is large (Y~ WR), the carrier and photon densities will exponentially rise/fall to their steady-state values
 - This regime is not useful in practice, since W_{348}^{max} is at $\sqrt{2}$

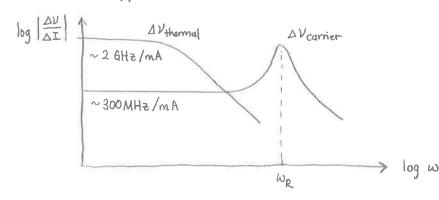


As the current I is modulated,
the carrier density N is modulated, and thus
the active region refractive index ma is modulated, and
the cavity length length is modulated, and finally
the resonant mode frequency is modulated.

We can write this relationship as: $\Delta v = -\frac{T v_g}{\lambda} \frac{d m_a}{d N} \Delta N$ We then find the modulation to be given by:

$$\frac{\Delta V}{\Delta I} = \frac{\alpha}{4\pi} \int v_g a \frac{m_i}{gV} \cdot \frac{\gamma_{ss} + i\omega}{\omega_e^2} H(\omega) \quad \text{where} \quad \alpha = -\frac{4\pi}{\lambda a} \frac{dn}{dN}$$

- Additionally, there is a further frequency modulation arising from the temperature modulation of the laser \Rightarrow ΔV thermal. This tends to dominate at low frequency
- = The two effects are illustrated below:



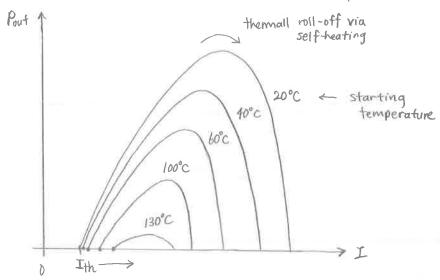
These are important effects to keep in mind when designing a directly modulated laser system!

Noise in lasers:

- Relative intensity noise: time variations in the photon and carrier densities due to random recombination and generation events in the absence of external modulation
 - > Noise floor on the output power magnitude
- · Frequency noise: random fluctuations in output frequency due to
 - (1) Spontaneous emission
 - (2) carrier density fluctuations, through the frequency chirping effect

· Temperature-dependent effects

- Any input power that does not leave the laser as light output is dissipated in the laser: $P_D = P_{in} P_{out}$. The resulting temperature rise of the laser is $P_D Z_T$ where Z_T is the thermal impedance
- The laser's L-I characteristic changes: Nor of T and Ith of et 1/To (approximately)
 - · Above threshold, Mi decreases and oi increases with T, to give a lower quantum efficiency
 - · Eventually the output power decreases with current due to heating effects
 - Auger recombination
 - Electron leakage from active region by thermal excitation

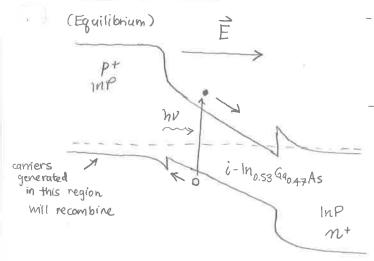


- Warelength matching: a big issue in VCSELs

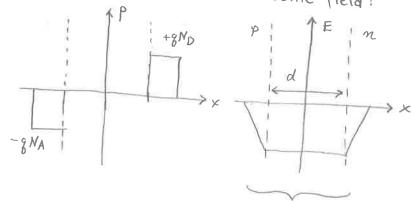
- (1) the gain peak shifts with temperature
- (2) the wavelength of the cavity mode shifts with temperature, due to thermal expansion, etc.
- the wavelength mismatch lowers
 the gain of the cavity mode, so
 Ith increases away from the optimal
 temperature (usually 300 K)

Photo detectors and photo conductors

· p-1-n photodiode



- The light-absorbing region is an intrinsic semiconductor sandwiched between heavily doped, preferably larger bandgap p- and n- regions
- The i-region becomes almost fully depleted of carriers and has a constant electric field:

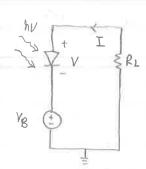


depletion region

The photodiode typically is reversebiased so that a high electric field exists in the i-region, which helps

sweep the carriers to their respective contacts.

- Wavelength sensitivity: by making the 1-region width $d \sim 1/\alpha(\lambda)$, the p-i-n will absorb most of the incident light at the wavelength A
 - This is only practical for λ shorter than the bandgap
 - Photododes of different thicknesses can be stacked so that each layer is sensitive to a different wavelength (shortest & goes on top)
- - = A thinner absorption region provides a larger fmax, but less optical absorption ⇒ a fundamental trade-off between speed and signal strength!
 - A thinner absorption region also increases the diode capacitance $C = \frac{EA}{W_T}$ This will end up limiting the speed, but there may be structures that are able to circumerent this
- The p-i-n provides no gain, but has low noise.

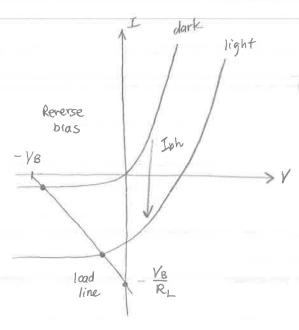


· The photodiode current in the dark is:

which is that of a normal diode

 In the presence of optical absorption, this is modified to

where

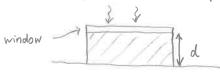


= 3rd quadrant: "photoconductive" mode

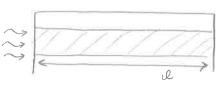
- 4th guadrant: "photovoltaic" mode

. The quantum efficiency depends on structure

- Surface-illuminated p-i-n



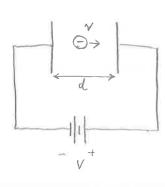
- Waveguide p-i-n



$$\eta = \eta_1(1-R)(1-e^{-\int \alpha d})$$
facet
reflectivity
factor

Paradox: a photon generates an electron and a hole, does it produce 2g of charge?

First, introduce Ramo's theorem



The current produced in the external circuit by a charge moving with relocity V(t) between the plates is:

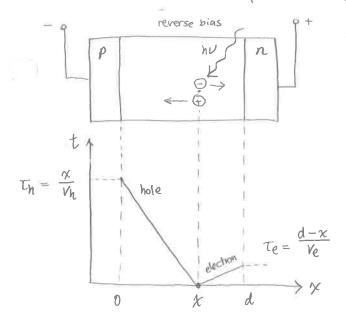
$$i(t) = \frac{q\gamma(t)}{d}$$

- Proof: work done by the field to move the charge dx is $aW = gE dx = g\frac{V}{d} dx$

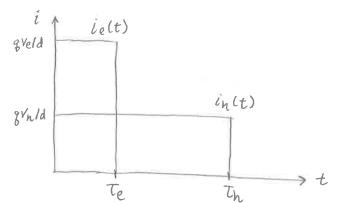
• Equate this to the work done by the power supply $dW = \frac{9V}{d} dx = i(t)V dt \Rightarrow i(t) = \frac{9}{d} \frac{dx}{dt} = \frac{9}{d} v(t)$

* The cument stops when $V(t) \rightarrow 0$; when e^- reaches contact

= Find the response of a photo-generated electron-hole pair



- Find how long it takes each carrier to reach the contact
 Use Ramo's theorem to find the photocument
 - over time



- Integrate the cument to find the total collected charge

$$Q = \int i_e(t) dt + \int i_h(t) dt$$

$$= \frac{g v_e}{d} T_e + \frac{g v_h}{d} T_h = \frac{g v_e}{d} \frac{d - x}{v_e} + \frac{g v_h}{d} \frac{x}{v_h}$$

$$Q = g \rightarrow 1 \text{ charge is detected!}$$

- Now suppose that a pulse of many photons produces electron-hole pairs through the volume of the absorption region. The resulting current i(t) goes to 0 once the last comier reaches its contact
 - · Since in the worst case the carrier must travel the whole distance of and holes are slower, the fransit time Ttr & d/V4.
- The response time is limited both by the transit time and the RC time $T = T_{tr} + T_{RC} = \frac{d}{y_h} + R \frac{\varepsilon A}{d}$

Optimize:
$$\frac{\partial \tau}{\partial d} = \frac{1}{\gamma_h} - \frac{\Re A}{d^2} = 0 \Rightarrow d = \sqrt{\Re A \gamma_h}$$

$$T_{opt} = 2\sqrt{\frac{R \epsilon A}{Y_h}} \Rightarrow \int_{3dB}^{opt} \approx \frac{1}{2\pi T_{opt}}$$

$$f_{3dB}^{\circ pt} = \frac{1}{4\pi} \sqrt{\frac{V_h}{R \varepsilon A}}$$

Typical values TRC ≈ 14ps, Ttr ≈ 20ps > f 3dB = 9.7 GHZ

Bandwidth - efficiency product

· Surface-illuminated p-i-n: assume R=0%, assume transit time limited

$$\eta \times f_{3dB} = \eta_{i}(1 - e^{-\alpha d}) \frac{1}{2\pi} \frac{V_{h}}{d} \approx \eta_{i}(1 - (1 - \alpha d)) \frac{1}{2\pi} \frac{V_{h}}{d}$$

$$\Rightarrow \eta \times f_{3dB} = \frac{\eta_{i} \alpha V_{h}}{2\pi}$$
weakly absorbing limit

The product does not depend on dimensions: a trade-off between efficiency and speed exists! (or rephrased: signal strength vs. speed)

* Waveguide p-i-n: assume R=0%, assume RC-limited

$$\eta \times f_{3d8} = \eta_{i}(1 - e^{-T\alpha L}) \frac{1}{2\pi Rc}, \quad C = \frac{\varepsilon Lw}{d}$$

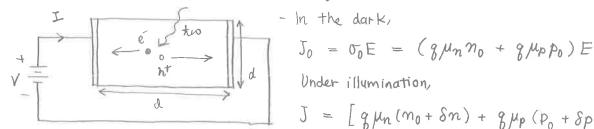
$$\approx \eta_{i} \int \alpha L \frac{1}{2\pi R} \frac{d}{\varepsilon Lw}$$

$$\eta \times f_{3d8} = \frac{\eta_{i} \int \alpha d}{2\pi R \varepsilon w} \quad (independent of L)$$

* In either type of detector, there is a tradeoff between efficiency and bandwidth (through d or I)

· Photoconductors

- A photoconductor's resistivity changes with the absorption of incident light



$$J_0 = \sigma_0 E = (g \mu_n n_0 + g \mu_p p_0) E$$

$$J = [g\mu_n(n_0 + \delta n) + g\mu_p(p_0 + \delta p)]E$$

$$\Delta J = J - J_0 = (g \mu_N \delta n + g \mu_P \delta P) E$$

$$\Delta J = g(\mu n + \mu p) \delta n E$$

or
$$\Delta \sigma = g(\mu_n + \mu_p) \delta n$$

- Both contacts on the photoconductor must be Ohmic
- The optically injected carrier density is given by the steady-state condition:

$$\frac{\partial}{\partial t}(\delta n) = G_0 - \frac{\delta n}{T_n} = 0 \quad \Rightarrow \quad \delta n = G_0 T_n$$
photogeneration carrier rate rate rifetime

The photogeneration rate is
$$G_0 = \frac{P_{opt}}{\hbar w} \cdot \frac{1}{lwd} \cdot \eta$$

photons incident volume same as in photodiode per second (assuming $\eta = \eta_i(1-R)(1-e^{-\alpha d})$

The generated current is:

$$\Delta I = dw \Delta J = lwg(\mu_n + \mu_p) \delta n E$$

 $d \int \int \int A - dW$

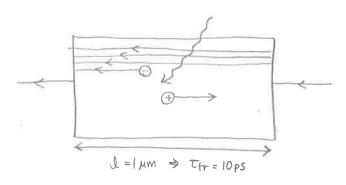
$$\Delta I = dwg\mu n E \cdot \eta \frac{\rho_{opt}}{\hbar w} \frac{1}{dwd} Tn = g \nu_n \eta \frac{\rho_{opt}}{\hbar w} \frac{1}{L} Tn$$

$$\Rightarrow \Delta I = (\eta P_{\text{opt}} \frac{g}{\hbar w}) \cdot \frac{T_n}{T_{\text{tr}}}$$

$$primary photoconductive gain$$
absorbed photons * g

The photoconductive gain
$$\frac{\Delta I}{Iph} = \frac{T_n}{T_{tr}} \leftarrow \frac{recombination}{transit}$$
 time

Interpretation: if many electrons travel across the electrodes before recombining with a hole, there is net gain.



. The mechanism is the same as that in a bipolar fransistor:

$$\beta = \frac{ic}{ib} = \frac{T_{\pm}}{T_{rb}}$$

· The generated electron reaches the electrode before the hole

· Current continuity forces the external circuit to supply anther electron, which also crosses the device before the hole reaches its electrode

· This continues until the electron recombines with the hole

> Many electrons pass through the circuit for every photon absorbed, leading to gain

·
$$|n| ||-V|$$
, gain $\approx \frac{1 \text{ ns}}{10 \text{ ps}} = 100$

The photoconductor's responsibily is

$$R_{\lambda} = \frac{\Delta \Sigma}{P_{opt}} = \eta \frac{g}{\hbar w} \cdot \frac{T_n}{T_{tr}}$$
 which has units of $\frac{A}{W}$

Photoconductor frequency response (small signal)

Consider again the equation

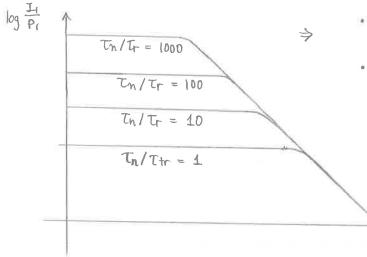
$$\frac{d}{dt}(\delta n) = G_0 - \frac{\delta n}{Tn}$$

If
$$\delta n = N_1 e^{i\omega t}$$
, this becomes $i\omega N_1 = \eta \frac{P_1}{\hbar \omega} \frac{1}{\ell n d} - \frac{N_1}{\ell n}$
 $\delta P = P_1 e^{i\omega t}$
 $(i\omega + \frac{1}{\ell n})N_1 = \eta \frac{P_1}{\hbar \omega} \frac{1}{\ell n d}$
 $N_1 = \eta \frac{P_1}{\hbar \omega} \frac{1}{\ell n d}$

Recall
$$I_1 = dw J_1 = dw g v_e N_1$$

$$I_1 = P_1 \eta \frac{g}{h w} \frac{v_e}{l} \frac{1}{i w + 1/t_h}$$

$$\Rightarrow \frac{I_1}{P_1} = \frac{\eta g}{h w} \left(\frac{t_h}{t_{tr}}\right) \frac{1}{l + i w t_h}$$



- · Greater gain increases the low frequency response, but reduces the 3dB bandwidth
- · The gain-bandwidth product of the device is always conserved

- Photoconductor structure

· A conventional photoconductor has an m-i-n structure, in which electrons can be injected from both sides

> log w

- · To increase In, n-i-p-i superlattice structures are sometimes used to create spatially separated potential wells for electrons and holes
- · To further improve T_n , heterojunction superlattices with $\Delta E_c \approx 0$ but large ΔE_V can be used to create potential wells for holes but allow electrons to flow freely

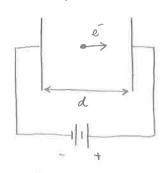
Noise in photodetection

- Noise can arise due to:
 - 1) Random arrival of photons
 - 2) Randomress in the generation of photoelectron-holes
 - 3) Randomress in the gain process (for avalanche photodiodes)
 - 4) Circuit no ise: shot noise and thermal no ise
- The number of photons arriving on a detector follows a Poisson distribution (i.e. the probability that an event occurs within a given time interval is distributed uniformly over that interval)
 - If $\bar{m}=\#$ photons arriving within an interval T, the probability that exactly n photons arrive within that interval is

$$p(n) = \frac{\overline{n}^n \cdot e^{-\overline{n}}}{n!}$$

of course, the mean =
$$\overline{n}$$
 \Rightarrow SNR = $\frac{\overline{n}^2}{\sigma^2} = \overline{n}$

Shot noise: the random arrival times of photons upon the detector translantes into random flow of charge, which produces a noisy current signal. Consistent with the properties above, the noise current is derived as follows:



- We model the generation of charge in a photodetector as a current pulse associated with a moving cruitye between two capacitor plates. From Ramo's theorem, $i_\ell(t) = \frac{3}{d} V(t) \qquad \text{for} \quad 0 \le t \le T_t$

$$i_{\ell}(t) = \frac{8}{d} v(t)$$
 for $0 \le t \le \tau$

Now add together a sequence of charges generated at random times:

$$i_{\tau}(t) = \sum_{i=1}^{N_{\tau}} i_{e}(t-t_{i})$$
 for $0 \le t \le T$

Take the Fourier transform:

$$i_{T}(f) = \sum_{i}^{N_{T}} i_{e}(f) e^{-i2\pi f t_{i}}$$

F.T. of ielt) Shift property of F.T.

Noise power is proportional to 1112:

$$i_{T}(f) i_{T}^{*}(f) = |i_{e}(f)|^{2} \cdot \sum_{i}^{N_{T}} \sum_{j}^{N_{T}} e^{-i2\pi f(t_{i} - t_{j})}$$

$$|i_{f}(f)|^{2} = |i_{e}(f)|^{2} \cdot \left[N_{T} + \sum_{i \neq j}^{N_{T}} \sum_{j}^{N_{T}} e^{-i2\pi f(t_{i} - t_{j})}\right]$$

$$\downarrow_{i=j}^{N_{T}} \text{ terms}$$

$$|i_{t}(f)|^{2} = |i_{e}(f)|^{2} [N_{T} + \sum_{i \neq j} \sum_{j} e^{-i2\pi f(t_{i} - t_{j})}]$$

Now take the ensemble average: since the times t_i are randomly correlated, all the terms $e^{-i2\pi f(t_i-t_j)}$ with $i\neq j$ average to zero

$$\Rightarrow \langle |i_f(f)|^2 \rangle = |i_e(f)|^2 N_T = |i_e(f)|^2 T \cdot \overline{N}$$
Then find the FT $i_e(f)$:

average # electrons / time

$$i_e(f) = \int_0^{\tau_t} i_e(t) e^{i2\pi f t} dt$$

$$\approx \frac{9}{d} \int_0^{\tau_t} \frac{dx}{dt} dt = \frac{9}{d} \int_0^d dx = 9$$

Assumed low frequency $f \ll \frac{1}{\tau_t}$

$$\Rightarrow \langle |i_{\tau}(f)|^2 \rangle = g^2 T \overline{N}$$

The final step is to take out the dependence on the interval T by finding the spectral density of the noise:

$$P_{T} = \frac{1}{T} \int_{-T/2}^{T/2} \langle |i_{T}^{2}(t)| \rangle dt = \frac{1}{T} \int_{-\infty}^{\infty} \langle |i_{T}(f)|^{2} \rangle df - \frac{2}{T} \int_{0}^{\infty} \langle |i_{T}(f)|^{2} \rangle df$$

$$= \int_{0}^{\infty} S(f) df$$

where
$$S(f) = \frac{2}{T} \langle |i_T(f)|^2 \rangle = 2g^2 \overline{N} = 2g (g \overline{N}) = 2g \overline{I}_{R}$$
 bias

Therefore, the noise cument is given by:

$$\langle i_N^2 \rangle = 2g \langle I \rangle \Delta f$$

mean bandwidth current of noise (DC)

Unlike Johnson noise, shot noise is only present when the mean photocument is nonzero, i.e. when photons are actually incident on the detector.

• Thermal noise associated with random thermal motion of charge Although it will not be derived here, this noise is:

$$\langle i_N^2 \rangle = \frac{4kT}{R} \Delta f$$
 $\langle v_N^2 \rangle = 4kTR \Delta f$ thermal current noise thermal voltage noise

Unlike shot noise, the thermal noise in the photodetector does not depend on . The number of photons incident

· Total noise:

$$\langle i_N^2 \rangle_{\text{total}} = 2g \langle I \rangle \Delta f + \frac{4kT}{R} \Delta f$$

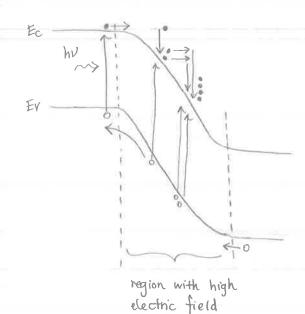
The ratio of signal to noise power in a detector without gain (such as a p-i-n photodiode) is:

$$SNR_{pin} = \frac{(RP)^2}{2g(RP + I_D)\Delta f + \frac{4kT}{R}\Delta f}$$

In a detector with gain, the shot noise term needs to be modified (see the next section on avalanche photodiodes)

Avalanche Photodiode (APD)

A photodiode with an intrinsic carrier multiplication mechanism



- In the absorption region, the large electric field induces avalanche breakdown; carriers moving in this region to quickly gain enough kiretic energy to cause impact ionization, creating a new electron-hole pair
- · In the ideal case, only one type of carrier experiences a multiplication effect. The current enhancement in this case is:

$$\frac{dJ_n}{dx} = \alpha_n J_n(x) \Rightarrow J_n(x) = e^{\alpha_n x} J_{no}$$
 at the edge of the absorption region,

$$J_n(w) = e^{\alpha mW} J_{no}$$
 where $\alpha_n = e^{-impact}$ ionization coefficient (cm⁻¹)

multiplication current without multiplication

In practice, since each impact ionization process generates an electron-hole pair, the generated holes can also cause impact ionization while moving in the opposite direction. In this case, coupled differential equations must be solved:

$$\frac{dJ_n}{dx} = \alpha_n J_n + \alpha_p J_p$$

$$-\frac{dJ_p}{dx} = \alpha_n J_n + \alpha_p J_p$$

Width = W

where In and Ip are functions of the electric field in the depletion region

Note that

$$\frac{d}{dx}(J_m + J_p) = 0 \Rightarrow J = J_m + J_p$$
 is constant, due to current continuity $J_p = J - J_m$

Solve:

$$\frac{dJ_n}{dx} = \alpha_n J_n + \alpha_p (J - J_n) = (d_n - \alpha_p) J_n + \alpha_p J$$

If we assume $J_p(W) = 0$, then we can find that the multiplication factor is:

$$Mm = \frac{1-k}{e^{-(1-k)\alpha_n W}-k}$$
 where $k = \frac{\alpha p}{\alpha n}$

(k = 0 recovers the single-carrier case)

- APD response time

$$T = T_t + T_m \qquad \text{where} \qquad T_m \approx \frac{M_n \, kW}{V_e} + \frac{W}{V_n} \,, \qquad T_t = \frac{W_{abs}}{Y_h}$$

$$\text{transit time} \qquad \text{time of } \\ \text{multiplication} \\ \text{process} \qquad \qquad \text{approximate expressing} \\ \text{for the duration of the} \\ \text{entire process, assuming} \\ \text{large } M_n \qquad \qquad W_{abs} = \text{absorption}$$

$$\text{region width}$$

$$\Rightarrow$$
 $\tau \simeq \frac{M_{n}kW}{V_{e}} + \frac{W + W_{abs}}{V_{h}}$

- The gain-bandwidth product ignoring quantum efficiency is:

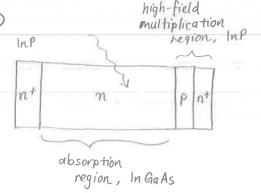
$$G \times BW = M_n \cdot \frac{1}{2\pi T}$$

$$G \times BW = \frac{V_e}{2\pi kW} \cdot \frac{1}{1 + \frac{1}{M_n k} \frac{V_e}{V_h} (1 + \frac{W_{abs}}{W})}$$
Generally dominates

- Conclusions: the factor k = dp/dn describes the extent to which feedback (due to impact ionization caused by hot holes) is present in the derice
 - (1) A larger k gives larger gain
 - (2) A larger K increases response time, reduces bandwidth
 - Overall, the second effect is stronger: the gain-bandwidth product drops with larger K, so this feedback is generally undesirable unless a large gain is the primary goal
 - (3) A larger k also increases moise
 - (4) The feedback process can cause instabilities, leading to uncontrolled avalanche: generally want to avoid this by making the multiplication region thin, W << Wabs

Separate-absorption-multiplication (SAM) APD

- · Absorb photons in a large intrinsic region
- · Allow generated carriers to drift under a moderate electric field
- · Avalanching occurs in a thin multiplication region



- Noise in APDs:

- · Since the gain mechanism involves random processes, noise is generated and then amplified. What is the effect on SNR?
 - · Signal power is multiplied by <M>2
 - * Shot noise power is multiplied by $\langle M^2 \rangle$
- . The quantity < M2 > is re-expressed using the noise factor F

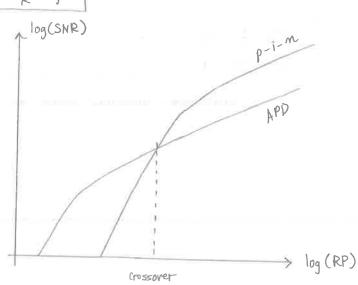
$$\langle M^2 \rangle = \langle M \rangle^2 F$$
 where $F = \frac{\langle M^2 \rangle}{\langle M \rangle^2} = k \langle M_n \rangle + (1-k) \left(2 - \frac{1}{\langle M_n \rangle}\right)$

- * F>1 and quantifies the excess noise caused by the gain process
- Note: F increases with k due to the feedback process, which produces even more noise
- The noise figure is defined as NF = 10 log F (dB)
- · Finally, the APD SNR is:

$$SNR_{APD} = \frac{(RP)^2 \langle M \rangle^2}{2g(RP + I_D)F \langle M \rangle^2 \Delta f + \frac{4kT}{R} \Delta f}$$

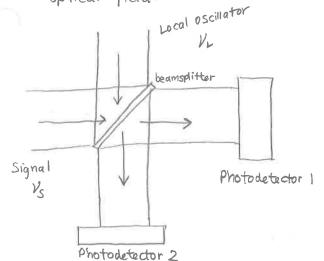
Since the noise factor F has a greater effect on SNR when the current is large, the APD is superior to the PIN only when the photocurrent is sufficiently small

The change in slope occurs once the shot noise > becomes larger than the thermal noise



· Cohenent detection

It is possible to measure both the magnitude and phase of an incident optical field.



Signal field
$$\mathcal{E}_{S} = \mathcal{E}_{S0} e^{i2\pi y_{S}t} e^{i\phi_{S}}$$

where the amplitude and phase Eso and \$ are modulated much more slowly than is

- Local oscillator field $\mathcal{E}_{L} = \mathcal{E}_{L0} e^{i2\pi \nu_{L} t} e^{i\phi_{L}}$
- · Each photodetector will measure the intensity of the optical field.

- An ideal double balanced mixer is one in which:

- (1) The beams are perfect plane waves and are perfectly aligned
- (2) The beamsplitter equally splits the optical power into the 2 paths
- (3) The two paths have relative lengths such that Es and Ex will have a To phase lag in one path, compared to their relative path in the other

Therefore the intensities on the two detectors will be:

$$\begin{aligned} |\xi_{\parallel}|^{2} &= \frac{1}{2} \left[|\xi_{s} - \xi_{L}|^{2} \right] = \frac{1}{2} \left[|\xi_{s0}|^{2} + |\xi_{L0}|^{2} - 2\xi_{s0}\xi_{L0} \cos(2\pi(v_{s} - v_{L})t + (\phi_{s} - \phi_{L})) \right] \\ &\Rightarrow |\xi_{\parallel}|^{2} = \frac{1}{2} \left(|\xi_{s0}|^{2} + |\xi_{L0}|^{2} \right) - \xi_{s0}\xi_{L0} \cos(2\pi\Delta v \cdot t + (\phi_{s} - \phi_{L})) \\ |\xi_{2}|^{2} &= \frac{1}{2} \left(|\xi_{s0}|^{2} + |\xi_{L0}|^{2} \right) + \xi_{s0}\xi_{L0} \cos(2\pi\Delta v \cdot t + (\phi_{s} - \phi_{L})) \end{aligned}$$

If ΔV is small, that is $V_S \approx V_L$, the incident light on the detector will be reatly monochromatic and the current is proportional to the optical intensity by: $I = \eta P_{\rm opt} / h \nu = \eta \left(\frac{1}{2} E |E|^2 A \right) / h \nu \propto |E|^2$

$$\Rightarrow I_{1} = \frac{1}{2}I_{S} + \frac{1}{2}I_{L} + \sqrt{I_{S}I_{L}} \cos(2\pi\Delta V \cdot t + (\phi_{S}(t) - \phi_{L}(t)))$$

$$I_{2} = \frac{1}{2}I_{S} + \frac{1}{2}I_{L} - \sqrt{I_{S}I_{L}} \cos(2\pi\Delta V \cdot t + (\phi_{S}(t) - \phi_{L}(t)))$$
interference term

to direct detection to direct detection of signal power of LO power

- Notice that if we subtract the two detected signals,

$$I_2 - I_1 = 2 \sqrt{I_s I_L} \cos \left[2\pi \Delta Y \cdot t + \phi_s(t) - \phi_L(t) \right] \qquad \text{(heterodyne)}$$

$$\approx 2 \sqrt{I_s I_L} \cos \left(\phi_s(t) - \phi_L(t) \right) \qquad \text{(homodyne, } \Delta V = 0 \text{)}$$

Knowing all the parameters of the local oscillator (\mathcal{E}_{H} , \mathcal{V}_{L} , ϕ_{L}), the magnitude of the signal \mathcal{E}_{SO} can be estimated from $I_{1}(t)$ or $I_{2}(t)$, while the phase $\phi_{S}(t)$ can easily be estimated from $I_{2}(t) - I_{1}(t)$.

Advantages of coherent detection.

- · Measure both the modulated intensity and modulated phase of the optical field
- · The measured field is large even if the signal is weak, due to the presence of a strong local oscillator; this improves the SNR,

$$SNR_{coherent} = \frac{\langle |2\sqrt{I_{s}I_{L}}\cos(\phi_{s} - \phi_{L})|^{2}\rangle}{2g\langle I_{L}\rangle \Delta f} = \frac{2\langle I_{s}\rangle\langle I_{L}\rangle}{2g\langle I_{L}\rangle \Delta f}$$

SNR coherent =
$$\frac{\langle 1s \rangle}{g \Delta f}$$
 where we assumed the local oscillator is much stronger than the signal: $\langle I_L \rangle \gg \langle I_S \rangle$ and is also large enough to overcome thermal noise: $2g \langle I_L \rangle \gg 4kT/R$

- Meanwhile, the SNR of direct detection without gain is:

$$SNR_{direct} = \frac{\langle I_s \rangle^2}{2q \langle I_s \rangle \Delta f + (4kT/R)\Delta f} \rightarrow \frac{\langle I_s \rangle}{2q \Delta f}$$

So the SNR of coherent detection is ~2x for relatively strong signals, and even more for weak signals

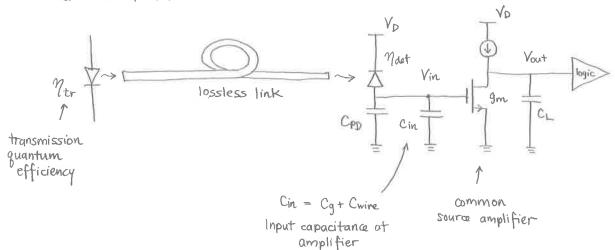
Compare to an APD:

$$SNR_{APD} = \frac{\langle I_S \rangle^2 M^2}{2q \langle I_S \rangle F M^2 \Delta f} = \frac{\langle I_S \rangle}{2q F \Delta f}$$

Note: An APD is not desirable with coherent detection, because it possesses an SNR benefit only when it is not shot noise limited!

So coherent detection is still 2x better than even a noiseless APD (F=1); the local oscillator provides a type of noiseless gain!

Receiver circuit analysis (from Ryan Going's thesis. Ch 2)
Consider the optical link below:



- The power expended emitter is:

$$P_{tx} = \hbar w \cdot \eta_{tr} \cdot N_{ph} \cdot f_{bit}$$
 where $N_{ph} = {}^{\#}photons/bit$
 $f_{bit} = bit rate (s^{-1})$

The power expended at the receiver is:

· It can be shown that the transconductance of the MOS amplifier is:

$$g_m = \frac{I_D}{2(V_{qs} - V_{th})} = \frac{I_D}{2V_{ov}}$$
 where $V_{ov} = V_{qs} - V_{th}$

· We assume that the transistor is designed so that its 3dB bandwidth coincides with the bit rate foit. Assuming some other ideal amplifier properties, we can obtain

$$2\pi f_{bit} = \frac{g_m}{A_V C_L}$$
 where $A_V = \frac{V_{out}}{V_{in}}$ is the gain

The input voltage is the voltage produced on the collection of capacitors Cpp, Cwire, and Cg when charge is generated by the absorption of a photon (assume a photodiode with no gain)

$$\Rightarrow$$
 Vin = $\frac{g \text{Nph } \eta \text{ det}}{CpD + Cwire} + Cg$

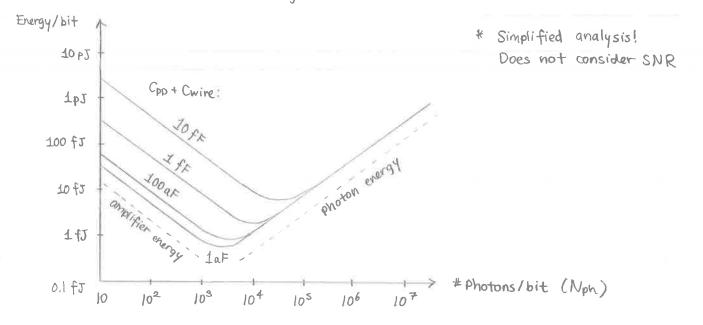
* Putting this together, we get:

$$P_{rx} = V_D \cdot 2V_{ov} g_m = 2V_D V_{ov} 2\pi f_{bit} C_L \frac{V_{out}}{V_{in}}$$

$$= \frac{4\pi V_{ov} V_{out} V_D C_L (C_{pD} + C_{wire} + C_g)}{g N_{ph} N_{det}} \cdot f_{bit}$$

Find the total power in the transmitter and neceiver, then divide by the bit rate to get the energy/bit of the optical link

- * Using the values: Mtr = Mdet = 1 $V_D = Yout = 1V, Yov = 100 \, mV \Rightarrow \text{ realistic values compared}$ $Cg = C_L = 200 \, aF \Rightarrow \text{ to ITRS}, 2013$
- · This leads to the following plot:



Condusions:

- Reducing the number of photons/bit does not always improve the energy/bit of communication:
 - · For large Nph, the transmitter dominates the energy cost, so reducing Nph helps
 - For small Nph, the amplifier energy dominates: the photodiode signal becomes very low, and more electrical power will be needed to amplify the signal
- The photodetector capacitance (including Cwire) must be reduced to decrease the lowest achievable energy/bit
 - · Low CpD -> more voltage per photon
 - · Chine can be reduced by tight integration of the photodiode and amplifier

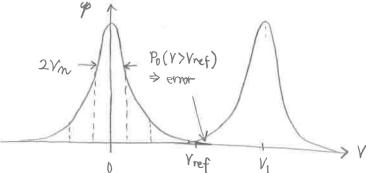
· Bit error rate and SNR

- The bit error rate is broadly defined as the fraction of detected bits that contains an error; a tolerable value is no more than 10^{-9} , a specification of 10^{-15} is common
- Suppose that if a '1' signal is detected, the expected voltage signal is V1. If a '0' signal is detected, the expected signal voltage is 0. Suppose also that the noise is Gaussian, with a probability distribution functions given below:

 $\varphi_{1}(V) = \frac{1}{V_{n}\sqrt{2\pi}} e^{-\frac{(V-Y_{1})^{2}}{2Y_{n}^{2}}}$

$$90(V) = \frac{1}{V_N \sqrt{2\pi}} e^{-\frac{V^2}{2V_N^2}}$$

where v_n^2 is the noise variance.



- In a receiver system, a digital comparator inspects the signal voltage V and decides that a bit is (1) if V > Vref, and (0) if V < Vref, and usually $Vref = \frac{1}{2}V_1$ is chosen.
 - \Rightarrow An error occurs if V > V ref even though the bit is (0) We can find this probability by integrating ϕ_0 :

BER = $\int_{V_{ref}}^{\infty} \varphi_0(V) dV = \int_{V_{ref}}^{\infty} \frac{1}{V_n \sqrt{2\pi}} e^{-\frac{V^2}{2V_n^2}} dV$

Let
$$Z = \frac{V}{V_n}$$
, $dZ = \frac{dV}{V_n}$, so $Z_{ref} = \frac{V_{ref}}{V_n}$

$$\Rightarrow$$
 BER = $\frac{1}{\sqrt{2\pi}}\int_{Z_{ref}}^{\infty} e^{-\frac{Z^{2}}{2}} dz = Q(Z_{ref}) = Q(\frac{V_{ref}}{V_{n}})$

Notice that if V_1 represents the expected voltage when a voltage is present, then $V_{ref}/V_n = V_1/2V_n$ represents half the SNR

$$\Rightarrow |BER = Q(\frac{1}{2}SNR) = \frac{1}{\sqrt{2\pi}} \int_{\frac{1}{2}SNR}^{\infty} e^{-\frac{Z^2}{2}} dz \qquad (Gaussian noise)$$

Quantum noise limit

- If we consider only shot noise, what is the minimum number of photons we can use and still transmit with a specified BER?

$$SNR = \frac{Iph}{\sqrt{2g}Iph\Delta f} = \frac{Iph}{\sqrt{g}Iphfbit} = \sqrt{\frac{Iph/g}{fbit}} = \sqrt{\frac{Nph}{fbit}}$$

$$\Delta f = \frac{1}{2}fbit$$
(Nyguist limit)

We know that if the noise is Gaussian, the required SNR = 6 to get BER = 10^{-9} This means that:

- The shot noise actually follows a Poisson distribution, which is well approximated as Gaussian noise only for large photon numbers. For small photon numbers, the error is significant, and the above is modified to 20 photons/bit (for (11)

11 Optical modulators

- There are various ways to modulate an optical signal other than directly modulating the laser diode itself; these methods are called external modulation
- Motivations for external modulation
 - · Potentially achieve a greater modulation bandwidth (not limited by wx)
 - · Avoid the frequency chirping effect in direct modulation
 - · Makes possible phase modulation -> coherent detection
- The cost is higher power consumption and greater complexity
- Most high-speed optical links > 10 Gbps rely on external modulators

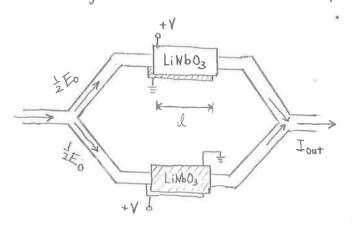
- Electro-optic modulators

- In general, the refractive index of a material can be a function of the field, due to non-linearities:

$$n(E) = M_0 + a_1 E + a_2 E^2 + ...$$

linear electro-optic guadratic electro-optic (Poctok) effect (kerr) effect

- · Cystals with centro-symmetry have $a_1=0$, so the guadratic effect dominates (though it still tends to be weak)
- . Careful engineering is necessary to produce a desired Δn along the direction of propagation of the field
 - e.g. A field of 10^5 V/cm on GaAs can produce $\Delta m \approx 3 \times 10^{-4}$
- Integrated electro-optic intensity modulator:



The two paths experience the same An but opposite in sign

The waveguide mode is split equally at the first fork, then they are combined at the second fork:

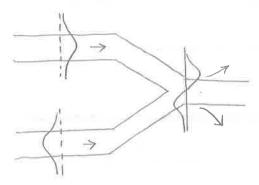
$$\begin{split} I_{\text{Out}} &= \left| \frac{1}{2} E_0 e^{i \frac{2\pi}{\lambda}} (n_0 + \Delta n) \mathcal{L} + \frac{1}{2} E_0 e^{i \frac{2\pi}{\lambda}} (n_0 - \Delta n) \mathcal{L} \right|^2 \\ &= \left| E_0 \right|^2 \left| \frac{1}{2} + \frac{1}{4} e^{i \frac{2\pi}{\lambda}} 2\Delta n \mathcal{L} + \frac{1}{4} e^{-i \frac{2\pi}{\lambda}} 2\Delta n \mathcal{L} \right| \\ &= \left| E_0 \right|^2 \left(\frac{1}{2} + \frac{1}{2} \cos \left(\frac{4\pi}{\lambda} \Delta n \mathcal{L} \right) \right) \end{split}$$

$$I_{out} = |E_0|^2 \cos^2\left(\frac{2\pi}{\lambda} \Delta n \cdot l\right)$$

intensity modulation through Am (V)

$$I_{out} = \frac{1}{2} |E_0|^2 \left[1 + \cos \left(\frac{4\pi}{\lambda} \Delta n(V) d \right) \right] = |E_0|^2 \cos^2 \left(\frac{2\pi}{\lambda} \Delta n(V) d \right)$$

- When V=0, $\Delta n=0$ and $I_{out}=|E_0|^2$
- * When V is such that $\frac{4\pi}{\lambda}\Delta n \cdot l = \pi$, I out = 0. What is the meaning of this destructive interference?

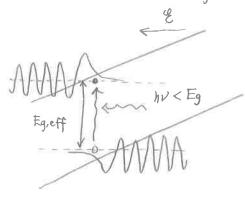


The interfering modes add up to a first-order mode, which is not supported by the output fiber, so it quickly leaks out

· LiNbO3 electro-optic modulators have been demonstrated with speeds in the 10-40 Gbps range

· Franz-Keldysh effect

- When a voltage is applied across a semiconductor absorption region, the electron wavefunctions become Airy functions (triangular potential well solutions to Schrödinger's equation):



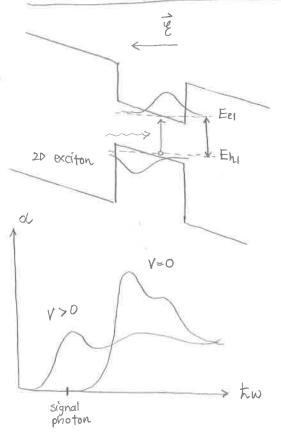
- As a result, due to wavefunction penetration into the gap, it becomes possible for photons below the bandgap energy to be absorbed. The effective bandgap is:

$$\Delta E_g = \left(\frac{9}{m^*}\right)^{1/3} (8 \pi \mathcal{E})^{2/3}$$

This can be considered a photon-assisted tunneling process

- The absorption can be modulated with voltage
 - . Choose a semiconductor with Eg $> h\nu$, where $h\nu$ is the energy of the signal photon
 - If no voltage is applied, the material is transparent to the photon \rightarrow $^{\prime}1^{\prime}$
 - * If a sufficient voltage is applied, the material can absorb the photon via the Franz-Keldysh effect \rightarrow (0)
 - Up to & ≈ 1000 cm-1 is achievable

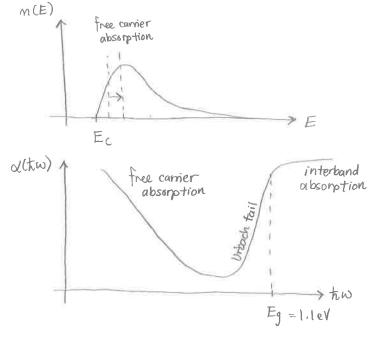
· Quantum-confined Stark effect



- · Under an electric field, the effective bandgap E_{hi}^{el} of a quantum well also decreases, due to the separation of the electron and hole wavefunctions to opposite sides of the quantum well
- The absorption edge effectively shifts to a lower energy, much like the Franz-Keldysh effect in a buk material under an applied voltage
- The generated electron and hole stay separated, so this can be considered a form of exciton absorption; the electron and hole interact through the Coulomb potential

· Free carrier effect in Si

* Free carrier absorption: absorption of photons by carriers, with initial and final states in the same band ⇒ photon energy is converted to kinetic energy



- · Due to the Fermi-Dirac distribution, most of the electrons occupy states within a few KT of the conduction band, with a decaying thermal tail at high energies
 - > number of available free carrier energy transitions decrease with larger photon energy
 - ⇒ a(thw) decreases with photon energy, until the onset of interband absorption
- · Carrier density dependence: the above physics can be captured in a free carrier absorption cross-section:

$$\alpha(\hbar\omega) = \sigma(\hbar\omega) \cdot n$$
 $(cm^{-1}) \quad (cm^{-2}) \quad (cm^{-3})$

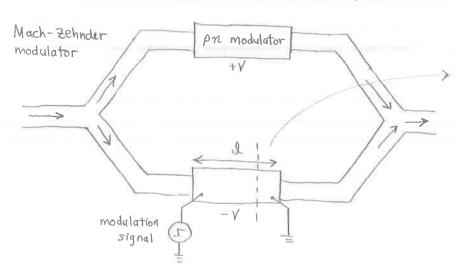
Free carrier index modulation: the Kramers-Kronig relations require that a change in the absorption coefficient of the material be accompanied by a change in the real permittivity, and therefore the refractive index Δn

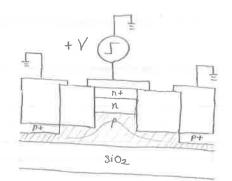
- In Si,
$$\Delta n \approx (-8.8 \times 10^{-22} \text{ cm}^3) \cdot \Delta N_e + (-8.8 \times 10^{-18} \text{ cm}^{3.0.8}) \cdot \Delta N_h^{0.8}$$

Thus, if $\Delta N_e = \Delta N_h = 10^{18} \text{ cm}^{-3}$, $\Delta n \approx -0.003$

This index change is capable of producing a large index change over I ~ 100 um - 1 mm

· Free carrier effect modulator for Si photonics





· By applying a reverse bias to the p-n junction, the depletion region which coincides with the optical mode can be further deprived of comiers (or supplied with camiers if biased in the forward direction)

- There is a fundamental trade-off between index change and loss!
 - · A longer device produces a larger phase shift, but also more loss
 - · A greater carrier density produces a greater index change, but also gives rise to more absorption
- = Carrier density can be modulated at ~ 50 Gbps

Thermo-optic effect

- The refractive index is a function of temperature, due to thermal expansion and the bandgap dependence on temperature
 - In Si, $\frac{dn}{dT} \approx 2 \times 10^{-4} \, \text{k}^{-1} \Rightarrow \Delta n = 0.05$ for $\Delta T = 300 \, \text{k} 270 \, \text{k}$ Larger than the free carrier effect!
- Slow response: time to heat/cool is ~1ms or ~1 µs for very small structures Suitable for tunable filters, but not modulators
- Large power consumption

Semiconductor Devices : selected concepts

· Current flow:

- Drift: Jar = gnv = gnu E where $\gamma = drift$ relocity

where μ is the mobility: $\mu = g \tau/m^*$ with $\tau = mean$ free time

- · M decreases with dopant concentration due to ionized impunity scattering
- · temperature dependence:
 - lattice scuttering increases with temperature > uv
 - ionized impunity scattering decreases with temperature because ions have a stronger effect on carriers with smaller thermal relocity: $V_{th} = \sqrt{\frac{3kT}{m^*}} \Rightarrow \mu \uparrow \text{ with temperature}$
 - In heavily doped samples, the mobility peaks at an intermediate temperature due to these effects
- At high fields, the drift relocity saturates due to strong scattering from optical phonons
- Diffusion: Jaif = g DN Vn (electrons) where D is the diffusivity Jdif = - 9 Dp Vp (holes)
- In equilibrium, the total cument is zero for each carrier type: for electrons, in 1D we have

$$J_{dr}^{n} + J_{dif}^{n} = g \mu_{n} n E + g D_{N} \frac{dn}{dx} = 0$$

Note:

Therefore we have

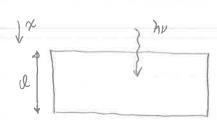
$$g \mu n m E + g D_N \left(-\frac{gE}{kT}\right) n = 0 \Rightarrow \mu n = \frac{g D_N}{kT}$$

or $\frac{D_N}{\mu N} = \frac{kT}{g}$ Einstein relation

 $D_N: cm^2/s$, $\mu_N: cm^2/V.s$

· Recombination/generation

- Optical generation



* Assume monochromatic incident light with $h\nu > E_g$ The optical density decays into the semiconductor as:

$$I = I_0 e^{-\alpha(h\nu) \cdot \chi}$$

Photons generate electron-hole pairs: thus the generation rate is the same for both carriers and has the same spatial dependence:

$$\frac{\partial n}{\partial t}\Big|_{\text{opt}} = \frac{\partial p}{\partial t}\Big|_{\text{opt}} = G_{LO} e^{-\alpha x}$$

- Shockley-Reed-Hall recombination/generation

- * Let $m_0, p_0 = equilibrium$ carrier concentrations, and $\Delta n = m n_0$ $\Delta p = p - p_0$
 - low-level injection: the excess minority carrier concentration is still much smaller than the majority carrier concentration. This constitutes a small perturbation: $\Delta n \ll p_0$ (in p-type), $\Delta p \ll n_0$ (in n-type)
 - . This will be assumed through the analysis
- * The net thermal SRH recombination rate is given by (in an n-type material):

$$\frac{\partial p}{\partial t}|_{SRH} = \frac{\partial p}{\partial t}|_{SRH,R} + \frac{\partial p}{\partial t}|_{SRH,Q}$$

$$= -CpN_Tp + CpN_Tp_0 = -CpN_T(p-p_0)$$
hole recombination rate:
hole generation rate

Fr > Fr. 50 # electrons that

hole recombination rate: EF > ET, so # electrons that occupy the travel simply equals the trap density NT

 The constant cp depends on the trap cross section ofp, the carrier thermal velocity Vth.p, etc.

$$\frac{\partial P}{\partial t}|_{SRH} = -\frac{\Delta P}{Tp}$$
 where $T_p = \frac{1}{C_p N_T}$

- In equilibrium, It set = 0, as expected. In nonequilibrium conditions, this quantity is nonzero, and is of course balanced in steady state by the excitation that drives the system out of equilibrium
- Tp is the minority carrier lifetime. In a p-type material, In is the minority carrier lifetime.

Continuity equations:

$$\vec{\nabla} \cdot \frac{\partial \vec{E}}{\partial t} = \vec{\nabla} \cdot (\vec{E}_{M0} \vec{\nabla} \times \vec{B} - \vec{E} \vec{J}) = \vec{E} \frac{\partial \rho}{\partial t}$$

$$\vec{\nabla} \cdot \vec{J} = - \frac{\partial \rho}{\partial t}$$

If we apply this separately to electrons and holes, we have:

$$\frac{\partial n}{\partial t} = \frac{1}{9} \vec{\nabla} \cdot \vec{J}_N + \frac{\partial n}{\partial t} \Big|_{\text{other}}$$

$$\frac{\partial p}{\partial t} = -\frac{1}{9} \nabla \cdot \vec{J}_p + \frac{\partial p}{\partial t} |_{other}$$

recombination and generation within a volume

Carrier lifetime measurement via optical excitation

Consider an n-type material in 1D, and that the intensity of radiation is uniform within its volume (i.e. weak absorption), and that there is no applied field:

$$\frac{\partial \Delta p}{\partial t} = -\frac{1}{8} \frac{\partial}{\partial x} J_{p,drift} + \frac{\partial p}{\partial t} |_{SRH} + G_{L}$$

$$= -\frac{1}{8} \frac{\partial}{\partial x} \left(-g D_{p} \frac{\partial}{\partial x} \Delta p \right) - \frac{\Delta p}{\tau_{p}} + G_{L}$$

$$\frac{\partial \Delta p}{\partial t} = D_p \frac{\partial^2 \Delta p}{\partial x^2} - \frac{\Delta p}{\tau_p} + G_L$$

$$\frac{1}{\text{diffusion}} \frac{1}{\text{SRH}} \quad \text{optical}$$
recombination generation

Under the assumption of uniform generation, Δp is the same everywhere

$$\Rightarrow \frac{\partial \Delta P}{\partial t} = G_L - \frac{\Delta P}{T_P}$$

Applying the boundary condition that $\Delta p(t=0) = 0$, the solution is:

$$\Delta p(t) = G_L T_P (1 - e^{-t/T_P})$$

- After the light shuts off, Dp decays from its steady - state valve:

$$\Delta p(t) = G_L T_p e^{-t/T_p}$$

A photoconductivity measurement can be used to extract tp:

$$\Delta \sigma(t) = g(\mu_n + \mu_p) \Delta p(t)$$

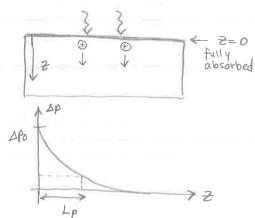
The decay constant of conductivity is easily manifested in the decay of a voltage across a load; this gives the lifetime Tp

- Diffusion length

- · Suppose that light is fully absorbed within an infinitesimal sheet at the semiconductor's surface: what is the distribution of minority carriers?
 - Away from the surface (z>0), $G_L=0$ and in Steady-state $\partial\Delta p/\partial t=0$:

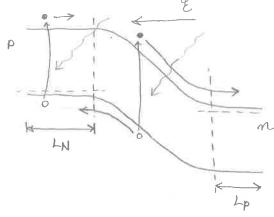
$$\frac{\partial \Delta p}{\partial t} = 0 = D_p \frac{\partial^2 \Delta p}{\partial x^2} - \frac{\Delta p}{T_p}$$

Solve this subject to $\Delta p(z=0) = \Delta p_0$, $\Delta p(z \neq \infty) = 0$



This length refers to the average distance that a minority carrier can travel before recombining

· In photodetectors:



- · In solar cells:
- To fully absorb sunlight, the cell must be sufficiently thick; the diffusion length must be at least this thickness, as otherwise the collection of minority carriers becomes very inefficient (low EQE)

- If a photon is absorbed in the depletion region, it is swept by the large electric field to the region where it is a majority carrier
 - > the external quantum efficiency (# absorbed photons to the # collected carriers) is & 1
 - If a photon is absorbed within a diffusion length outside the depletion region, the generated minority carrier has a good chance of reaching the depletion region, then collected
 - > FQE ≈ 50%
 - If a photon is absorbed more than a diffusion length from the depletion region, the generated minority carrier will almost certainly recombine before reaching the depletion region (via SRH)

Diode ideality factor

· An ideal diode has the following I-V characteristic:

$$I = I_0 \left(e^{gY/kT} - 1 \right)$$

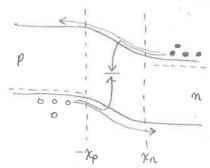
In getting to this equation, several ideal assumptions were made:

- (1) Low injection
- (2) No recombination/generation in the depletion region
- (3) Excess minority concentration → 0 at Ohmic contacts (i.e. quasi-neutral regions are several diffusion lengths thick)
- · Deviating from any of these assumptions changes the ideal characteristic. The "ideality factor" measures these deviations

$$I(V) \propto e^{gV/nkT}$$
 $n=1$ for ideal diode

This concept is relevant under forward bias when the I-V curve is exponential. The value of n can vary with voltage.

- · Assumption (1) is broken at high voltage, (2) is broken at low voltage, and (3) is generally valid.
- Recombination/generation in the depletion region



- Under forward bias, an excess current flows due to SRH recombination in the depletion region

$$I_{SRH} = -gA \int_{-xp}^{xn} \frac{\partial n}{\partial t} |_{SRH} dx$$

A rigorous treatment of SRH gives:

$$\frac{2n}{2t}\Big|_{SRH} = \frac{mp - n_i^2}{T_p(m+n_i) + T_n(p+p_i)}$$

where
$$m_i = n_i e^{(E_T - E_i)/kT}$$
, $p_i = n_i e^{(E_i - E_T)/kT}$

- The result for forward bias is:

$$I_{SRH} = \frac{9A\eta_i}{2T_0} W \frac{e^{\frac{9}{4}V/ET} - 1}{1 + \frac{V_{bi} - V}{kT/8} \frac{\sqrt{\ln t_p}}{2T_0} e^{\frac{9}{4}V/2kT}} \approx \frac{9A\eta_i}{\sqrt{\ln t_p}} \frac{V_{bi} - V}{kT/8} W - e^{\frac{9}{4}V/2kT}$$
depletion width.

The SRH current has ideality factor n=2

- Note: in reverse bias, an excess reverse current is seen that arises from SRH generation in the depletion region

- High cument phenomena:

· Series resistance:

$$I = I_0 e^{gV/kT} \rightarrow I_0 e^{g(V-IR_S)/kT}$$

This causes a high-cument roll-over of the diode I-V curve

· High-level injection: the injected minority carrier concentration becomes comparable to the majority carrier concentration on either side

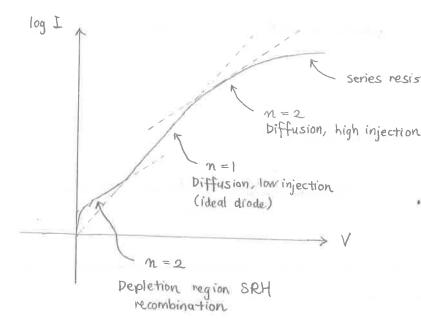
$$\Rightarrow \text{ in } p-\text{type}: \quad m_p p_p = m_p (N_A + m_p) = m_i^2 e^{8V/kT} \quad \text{where it is no longer true that}$$

$$m_p = -\frac{N_A}{2} + \frac{1}{2} \sqrt{N_A^2 + 4 m_i^2} e^{8V/kT}$$

$$m_p = \frac{N_A}{2} \left[\sqrt{1 + \frac{4m_i^2}{N_A^2}} e^{8V/kT} - 1 \right] \approx m_i e^{8V/2kT}$$
and in m -type:
$$p_m = \frac{N_D}{2} \left[\sqrt{1 + \frac{4m_i^2}{N_A^2}} e^{8V/kT} - 1 \right] \approx m_i e^{8V/2kT}$$

$$\Rightarrow J = \left(g \frac{P_N}{L_N} + g \frac{D_P}{L_P}\right) m_i e^{gV/2kT} \longrightarrow Thus, \quad m = 2 \text{ under high level injection}$$

- Combining these effects, a typical I-V curve may look like:



Note: under high injection, there is also a voltage drop induced in the quasi-neutral region

· AC response of diodes

The response is significantly different for a reverse-biased diode (e.g. a photodiode) vs. a forward-biased diode (e.g. an LED /laser)

- Reverse bias: He response is dominated by the junction capacitance

	deplet	rion	
P	1/2	++	n
	<>-	**************************************	

- An AC voltage signal in reverse bias tends to modulate the depletion width ${\sf W}$
- Since the conductance is small in reverse bias, the diode acts as a variable capacitor:

Is as a variable capacitor:
$$C_J = \frac{\mathcal{E}A}{W} \quad \text{where} \quad W \cong \sqrt{\frac{2\mathcal{E}}{gN_D}} (V_{bi} - V)$$
lightly doped side

- If the modulation is small, C_J is constant and the speed of the diode circuit to communicate its signal to a load is: $T = R_L C_J$

- Forward bias :

- The junction capacitance still exists, but there is also significant contribution from minority carrier charge oscillations; this is the origin of diffusion capacitance, which is frequency-dependent
- We must solve the small-signal version of the minority carrier diffusion equation. Consider the m-side of a pt-n junction:

$$\frac{\partial \Delta p}{\partial t} = Dp \frac{\partial^2 \Delta p}{\partial x^2} - \frac{\Delta p}{Tp}$$

$$(\omega \hat{p} e^{i\omega t} = Dp \frac{\partial^2 \hat{p}}{\partial x^2} e^{i\omega t} - \frac{\hat{p}}{Tp} e^{i\omega t}$$

$$0 = Dp \frac{\partial^2 \hat{p}}{\partial x^2} - \hat{p} \left(\frac{1}{Tp} + i\omega\right)$$

$$0 = Dp \frac{\partial^2 \hat{p}}{\partial x^2} - \frac{\hat{p}}{Tp} (1 + i\omega Tp)$$

where pe is the small AC modulation of the minority carrier concentration

The boundary conditions are:
$$\hat{p}(x \rightarrow \infty) = 0$$

and $\tilde{p}(x = x_n) = \frac{n_i^2}{N_D} e^{gV/kT} (e^{g\tilde{V}/kT} - 1)$

where \tilde{V} is the small AC Voltage signal, V is the DC bias

· Solution for current:

where
$$G_0 = \frac{dI}{dV} = \frac{gI_0}{kT} e^{gV/kT}$$
 is the low-frequency conductance $= \frac{I}{kT/g}$

$$\tilde{z} = G_0 \sqrt{1 + i\omega \tau_p} \cdot \tilde{\gamma}$$

Small-signal admittance $Y_D = G_D + i\omega C_D$

- The admittance can be separated into:

• Diffusion conductance
$$G_D = \frac{1}{\sqrt{2}}G_0\left(\sqrt{1+\omega^2G_p^2}+1\right)^{1/2}$$

• Diffusion capacitance
$$C_D = \frac{1}{\sqrt{2}} G_0 \cdot \frac{1}{w} \left(\sqrt{1 + w^2 t_p^2} - 1 \right)^{1/2}$$

- Since $G_0 \sim e^{gV/kT}$, the diffusion capacitance becomes very large as the forward bias is increased

- Frequency dependence:

· Low frequency
$$w \ll \frac{1}{\tau p}$$
: $C_D \simeq G_0 \frac{\tau p}{2} = \frac{I \tau p}{2kT/8}$

· Beyond this frequency, Co decreases with increasing w

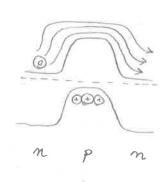
Implications of the diffusion capacitance How much signal voltage is produced by the absorption of N photons?

· Photodetector (reverse biased);

$$\delta \mathcal{N} = C^{2} \nabla \mathcal{N} \Rightarrow \nabla \mathcal{N} = \frac{\partial^{2} C^{2}}{\partial \mathcal{N}}$$

· Phototransistor (forward biased):

- In a phototransistor (npn) the absorbed photons generate holes in the base. The holes attract electrons, and for each hole stored there many electrons can flow across the base without being lost: the ratio is the same as that of photoconductive gain, Tp/Ttr



$$\Rightarrow$$
 $gN = C_{J}\Delta V + \Delta I \cdot T_{tr}$
 $\#e^{-}$ that flow $\#e^{-}$ tholes
across junction, each $\#e^{-}$ stored in
taking 1 transit time $\#e^{-}$ the base

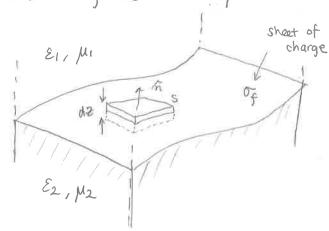
$$qN = (C_J + \frac{ITH}{kT/g})\Delta V \Rightarrow \Delta V = \frac{gN}{C_D + C_J}$$

In the BJT base, Tp > Ttr in calculating Cp. This is because the base side of the pn junction is very narrow.

⇒ because of diffusion capacitanæ, more photons are needed to produce a voltage ΔV; less sensitive!

(13) Electromagnetics and optics: selected concepts

= Electromagnetic boundary conditions



Apply Gauss's law to the pillbox: the sides contribute nothing to flux since the boundary is infinitely thin

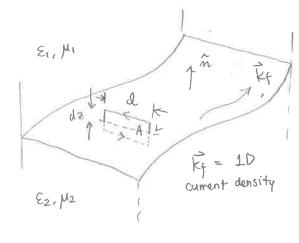
$$\vec{D_1} \vec{A} - \vec{D_2} \vec{A} = \vec{O_f} \vec{A}$$

$$\mathcal{E}_1 \mathcal{E}_1^{\perp} - \mathcal{E}_2 \mathcal{E}_2^{\perp} = \mathcal{O}_{\mathbf{F}}$$

= Use the same reasoning for magnetic field

$$\vec{\nabla} \cdot \vec{B} = 0 \Rightarrow \int_{S} \vec{B} \cdot d\vec{A} = 0$$

$$B_1^{\perp}A - B_2^{\perp}A = 0 \Rightarrow B_1^{\perp} = B_2^{\perp}$$



- Now consider an amperian loop along the surface:

$$\vec{\nabla} \times \vec{\hat{E}} = -\frac{\partial \vec{B}}{\partial t}$$

$$\oint \vec{E} \cdot d\vec{L} = -\frac{d}{dt} \int \vec{B} \cdot d\vec{A}$$

$$\vec{E_1} \cdot \vec{J} - \vec{E_2} \cdot \vec{J} = 0 \leftarrow \text{flux Vanishes}$$
as $dz \to 0$

$$\Rightarrow |\vec{E}_1| = |\vec{E}_2|$$

- Finally, use Ampere's law:

$$\vec{\nabla} \times \left(\frac{\vec{B}}{\mu}\right) = \vec{J}_f + \epsilon \frac{\vec{D}\vec{E}}{\vec{D}\vec{E}} \Rightarrow \vec{D} + \vec{D} = \vec{J}_f + \epsilon \frac{\vec{D}}{\vec{D}} = \vec{D} = \vec{D}$$

$$\Rightarrow H_1 \cdot \vec{l} - H_2 \cdot \vec{l} = I_f = \vec{k}_f \cdot (\hat{n} \times \vec{l})$$
 points in the direction normal to direction normal to the amperian loop

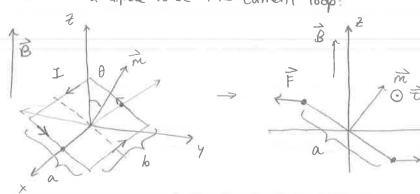
$$\vec{H}_1^{\parallel} - \vec{H}_2^{\parallel} = \vec{K}_f \times \vec{n}$$

$$\frac{1}{\mu_1} \vec{B}_1^{\parallel} - \frac{1}{\mu_2} \vec{B}_2 = \vec{K}_f \times \vec{n}$$

$$\vec{k_f} \times \hat{n}$$
 is the free surface cument density in the plane of the boundary. e.g. if $\hat{n} = \hat{z}$, $\vec{k_f} = k_{fy} \hat{x} - k_{fx} \hat{y}$

Torques and forces on magnetic dipoles

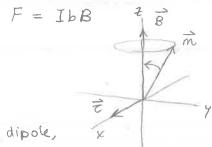
Consider a dipole to be this current loop?



There is no net torque on the sides of length a, because the forces cancel. On the other Sides with length b, the torque

$$\vec{t} = \vec{r} \times \vec{F}$$

 $\vec{t} = aF \sin \theta \hat{x}$



$$\Rightarrow \vec{t} = abIB \sin \theta \hat{x} = mB \sin \theta \hat{x}$$

$$\Rightarrow \vec{\tau} = \vec{m} \times \vec{B} \Rightarrow \text{precession}$$

The magnetic field acts to rotate the magnetic dipole, until it aligns with the field

The potential energy of the dipole in a magnetic field is given by: U = - m. B

Thus, in order to minimize U, the preferred configuration is to have m parallel with B. A preassion dipole tends to dampen and relax to this state

Effect on atomic orbits -> Diamagnetism

· Imagine that in an atom, an electron orbits the nucleus at a radius R and completes a revolution in time $T = 2\pi R/V$

$$I = \frac{-9}{7} = \frac{-97}{2\pi R} \Rightarrow \vec{m} = I\pi R^2$$

I is opposite the
$$\vec{m} = -\frac{1}{2}gvR^{2}$$

motion of electrons

· If we introduce a magnetic field, we need to add a Lorentz force to the classical centripetal force of the orbiting electron

$$m_e \frac{\gamma_B^2}{R} = \frac{1}{4\pi\epsilon_0} \frac{g^2}{R^2} + g\gamma_B B$$

Centripetal Coulomb force component

component

where NB = relocity w/ magnetic field

This leads to:

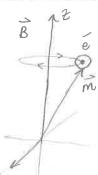
$$\gamma_{B} = \frac{1}{gB} \left(m_{e} \frac{V_{B}^{2}}{R} - \frac{1}{4\pi\epsilon_{0}} \frac{g^{2}}{R^{2}} \right) = \frac{m_{e}}{gBR} \left(V_{B}^{2} - V^{2} \right) \approx \frac{m_{e}}{gBR} 2V\Delta V$$

$$\Rightarrow \Delta V = \frac{1}{2} \frac{gBR}{me} \qquad \text{(change in electron speed)}$$

$$\Delta \vec{m} = -\frac{1}{2} g\Delta V R \hat{z} = -\frac{g^{2}R^{2}}{4m_{e}} \vec{B}$$

- · If a magnetic field is applied, the magnetic dipole moment changes in the direction opposite the direction of the field. This is ultimately because the magnetic field speeds up the electrons, but the electrons orbital motion has a negative dipole moment
- When $\vec{B}=0$, $\vec{M}=0$ in a material since the atoms are randomly oriented. But if $\vec{B}>0$, $\vec{M}<0$ due to this effect: this is diamagnetism

- Paramagnetism



* If we ignore atomic orbital motion, the electron is still a magnetic dipole due to its intrinsic orbital momentum, or spin:

electron spin angular momentum gyromagnetic ratio,
$$\gamma > 0$$

Thus, we still have the torque $\vec{t} = \vec{m} \times \vec{B}$ But note that the electron also has angular momentum \vec{J} , so there will be precession of the magnetic dipole:

$$\vec{\tau} = \frac{d\vec{J}}{dt} = \vec{m} \times \vec{B} \qquad \Rightarrow \qquad \frac{d\vec{m}}{dt} = \gamma \vec{m} \times \vec{B}$$

If we now include energy relaxation, which tends to minimize $U=-\vec{m}\cdot\vec{B}$, and phase relaxation, which reduces M_X and M_Y :

$$\frac{d\vec{m}}{dt} = \gamma \vec{m} \times \vec{B} - \frac{m_z - m_0}{T_1} - \frac{m_x + m_y}{T_2} \leftarrow \text{Bloch equation}$$

$$Precession \qquad Energy \qquad Phase \\ relaxation \qquad relaxation$$

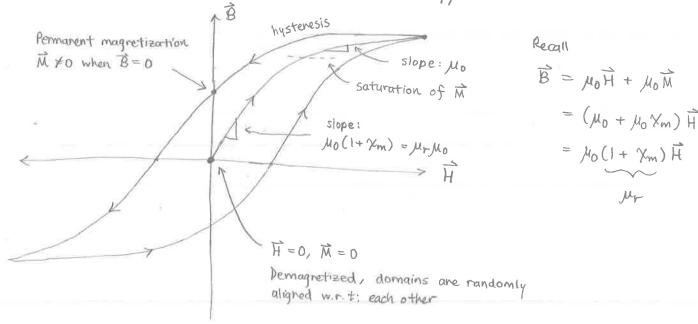
$$\gamma B = \text{precession} \qquad m_z = m_0 \text{ in} \\ \text{frequency} \qquad \text{steady state}$$

$$\text{where } \vec{B} = B\hat{\gamma}$$

- Note that \vec{m} tends to align with \vec{B} , as expected classically. However, due to the Pauli exclusion principle, the electrons in atoms tend to come in pairs with opposite spin \Rightarrow no net torque!
 - This effect is therefore only seen in some metals which have unpaired outer shell electrons, and in these materials is stronger than diamagnetism
 - · As $\vec{B} = 0$, $\vec{M} = 0$, again due to random orientation of electrons
 - · If $\vec{B} > 0$, $\vec{M} > 0 \Rightarrow paramagnetism$ Remains weak, as random collisions disrupt the alignment
 - Protons in the nucleus are also paramagnetic: this is the basis for nuclear magnetic resonance and magnetic resonance imaging

Fernomagnetism

- Exchange interaction: two unpaired valence electrons in adjacent atoms repel each other if they have the same spin, due to the Pauli exclusion principle
 - · The reduced wavefunction overlap gives the parallel-spin state less electrostatic potential energy
 - ⇒ long-range alignment of magnetic dipole moments
- In ferromagnets, the spins of all the unpaired electron within a domain Cwhich can have macroscopic size) are all aligned
 - . The dipoles within a domain respond in unison to a field \vec{B} , and can retain their magnetization after \vec{B} switches off



· Phase changes on reflection

The normal incidence amplitude reflectivity is polarization-independent:

$$r = \frac{m_1 - m_2}{m_1 + m_2} \longrightarrow \pi$$

To phase shift if $m_1 < m_2$ (low \Rightarrow high)

To phase shift if $m_1 > m_2$ (high \Rightarrow low)

To phase shift for reflection from perfect conductor

Total internal reflection

* From Snell's law: $m_1 \sin \theta_1 = m_2 \sin \theta_2$ Let the wave be incident from $m_2 > m_1$ at $\theta_2 = \theta$

$$\sin \theta_1 = \frac{m_2}{m_1} \sin \theta$$

At $\theta = \theta_c$, $\frac{m_2}{m_1} \sin \theta_c = 1$ so $\theta_1 = \frac{\pi}{2}$. If $\theta > \theta_c$, θ_1 is no longer a real number! But let's proceed to examine the wave in m_1

$$E_{\parallel} = e^{i(k_0 m_1 \sin \theta_1) \chi} i(k_0 m_1 \cos \theta_1) \pm e^{ik_0 m_2 \sin \theta_1 \cdot \chi} e^{ik_0 m_1 \cos \theta_1 \cdot Z}$$

where
$$\cos \theta_1 = \cos \left(\sin^{-1}\left(\frac{m_2}{m_1}\sin \theta\right)\right) = \sqrt{1 - \left(\frac{m_2}{m_1}\sin \theta\right)^2}$$

$$\cos \theta_1 = i \sqrt{\left(\frac{m_2}{m_1}\sin \theta\right)^2 - 1}$$

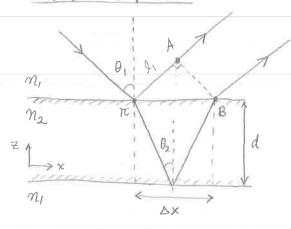
$$E_{1} = e^{ik_{0}m_{2}\sin\theta x} e^{-k_{0} Z \sqrt{\left(\frac{m_{2}}{m_{1}}\sin\theta\right)^{2} - 1 \cdot m_{1}}$$
The peretration depth is:
$$e^{-m_{1}k_{0}S\sqrt{\left(\frac{m_{2}}{m_{1}}\sin\theta\right)^{2} - 1}} = e^{-1}$$

$$\delta \pi R = 2\pi m_{1} \sqrt{\left(\frac{m_{2}}{m_{1}}\sin\theta\right)^{2} - 1} = 2\pi \sqrt{m_{2}^{2}\sin^{2}\theta - m_{1}^{2}}$$

- · There is a phase shift on total internal reflection that is a continuous function of angle; the amount of the phase shift is tedious to calculate
 - Phase shift is 0 at $\theta = \theta_c$
 - Phase shift increases as O increases beyond Oc
 - Phase shift is π at $\theta = \frac{\pi}{3}$

The phase shift might be interpreted as the peretration of the wavefront (NOT the everyy) into the other medium before being reflected back.

Thin film reflection



M2> m1: this is important/

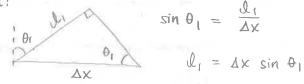
In order for the two reflected waves to interfere constructively, their phase must match (to a factor of 2π) at points A and B

Phase accumulated by reflected ray 1

$$\Delta x = 2d \tan \theta_2$$

where
$$\sin \theta_2 = \frac{m_1}{m_2} \sin \theta_1$$

then:



$$\sin \theta_1 = \frac{Q_1}{\nabla x}$$

$$l_1 = \Delta x \sin \theta_1$$

$$l_1 = 2d \tan \theta_2 \sin \theta_1$$

$$\phi_1 = k_0 n_1 l_1 + \pi$$

$$= k_0 \cdot 2d \tan \theta_2 \cdot m_1 \sin \theta_1 + \pi$$
from top

- Phase accumulated by reflected rate 2

$$\phi_2 = k_0 n_2 \cdot 2 \frac{d}{\cos \theta_2} = k_0 \frac{2dn_2}{\cos \theta_2}$$

The phase difference is:

from top

Surface

reflection,

Since
$$m_1 < n_2$$

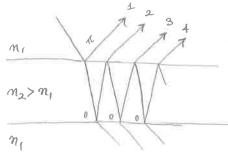
$$\Delta \phi = \phi_2 - \phi_1 = k_0 \frac{2dm_2}{\cos \theta_2} - k_0 2d \tan \theta_2 m_1 \sin \theta_1 - \pi$$

$$= 2k_0 d \left(\frac{m_2}{\cos \theta_2} - m_1 \tan \theta_2 \sin \theta_1\right) - \pi$$

$$= 2k_0 d m_2 \left(\frac{1}{\cos \theta_2} - \tan \theta_2 \sin \theta_2\right) - \pi$$

$$= 2m_2 k_0 d \left(\frac{1 - \sin^2 \theta_2}{\cos \theta_2}\right) - \pi \Rightarrow \Delta \phi = k_0 \left(2m_2 d \cos \theta_2\right) - \pi$$

Subsequent reflected rays have $\Delta \phi = k_0 (2m_2 d \cos \theta_2)$ since all of the internal reflections induce no phase shift in this geometry, and transmitted rays always have no phase shift at the boundary



- Therefore, unless the initially reflected ray is very strong (unlikely), the film's reflectivity depends on whether rays 2, 3, 4, ... interfere constructively;

$$\Rightarrow \Delta \phi = k_0 2m_2 d \cos \theta_2 = 2\pi m$$

$$\lambda_{m} = \frac{1}{m} 2m_{2} d \cos \theta_{2}$$

Reflectivity maxima:

$$\frac{\lambda_m}{n_2} = \frac{2d}{m} \cos \theta_2 \approx \frac{2d}{m} \text{ near normal}$$

· Reflectivity is maximized for d = \frac{1}{2}(\lambda/m_2) (half wavelength)

Metal optics

· Conductivity and the complex permittivity Ampere's law $\rightarrow \vec{\nabla} \times \vec{H} = \vec{J} + \epsilon \frac{\partial \vec{E}}{\partial t}$ where $\epsilon = \epsilon_r \epsilon_0$, ϵ_r is real = $\sigma \vec{E} - i\omega \epsilon \vec{E}$ sinusoidal field $\vec{E} \sim e^{-i\omega t}$ $= \left(\varepsilon + \frac{\sigma}{-i\omega}\right) \left(-i\omega\vec{E}\right)$ $= (\varepsilon_{r} \varepsilon_{0} + i \frac{\sigma}{\omega}) \frac{\partial \overline{\varepsilon}}{\partial t} = \varepsilon \frac{\partial \overline{\varepsilon}}{\partial t}$

where the complex permittivity is
$$\widetilde{\mathcal{E}} = \mathcal{E}_r \mathcal{E}_0 + i \frac{\sigma}{\omega}$$

In metals, imaginary permittivity = conductivity

Skin depth Consider a wave traveling through a metal &, which varies as e ikz $\widetilde{k} = k_0 \widetilde{n} = \frac{\omega}{c} \sqrt{\frac{\widetilde{\epsilon}}{\varepsilon_0}} \frac{\widetilde{n}}{m_0} = \frac{\omega}{c} \sqrt{\frac{\varepsilon_{\mu r}}{\varepsilon_{\mu r}}} + i \frac{\sigma_{\mu r}}{\omega \varepsilon_0}$ where $\tilde{\mu} = \mu r \mu o$ is real

$$\Rightarrow \operatorname{Im} \widetilde{K} = \frac{\omega}{c} \left[\frac{1}{2} \left[(\varepsilon_{r} \mu_{r})^{2} + (\frac{\sigma \mu_{r}}{\omega \varepsilon_{0}})^{2} - \varepsilon_{r} \mu_{r} \right] \right]$$

$$= \frac{\omega}{c} \frac{1}{\sqrt{2}} \left[\varepsilon_{r} \mu_{r} \right] + (\frac{\sigma}{\omega \varepsilon_{r} \varepsilon_{0}})^{2} - \varepsilon_{r} \mu_{r} \right]$$

$$= \frac{\omega}{c} \sqrt{\frac{\varepsilon_{r} \mu_{r}}{2}} \left[\left[1 + (\frac{\sigma}{\omega \varepsilon_{r} \varepsilon_{0}})^{2} - 1 \right]^{\frac{1}{2}} \right]$$

The skin depth is defined where e-Imk. 8 = p-1

$$\Rightarrow S = \frac{1}{\text{Im } \tilde{k}} = \int \frac{2}{\text{Er} \mu r} \frac{C}{\omega} \left[\sqrt{1 + \left(\frac{\sigma}{\omega \epsilon_r \epsilon_0} \right)^2 - 1} \right]^{-1/2}$$

$$= \int \frac{2}{\omega \mu r \mu_0} \sqrt{\frac{1}{\epsilon_r \epsilon_0 \omega}} \left[\frac{\sigma}{\omega \epsilon_r \epsilon_0} \sqrt{1 + \left(\frac{\omega \epsilon_r \epsilon_0}{\sigma} \right)^2 - 1} \right]^{-1/2}$$

$$= \int \frac{2}{\omega \sigma \mu_r \mu_0} \left[\sqrt{1 + \left(\frac{\omega \epsilon_r \epsilon_0}{\sigma} \right)^2 - \frac{\omega \epsilon_r \epsilon_0}{\sigma}} \right]^{-1/2}$$

$$S = \left[\frac{2}{\omega \sigma \mu_{\Gamma} \mu_{0}} \left[\left[1 + \left(\frac{\omega \varepsilon_{1} \varepsilon_{0}}{\sigma}\right)^{2} + \frac{\omega \varepsilon_{\Gamma} \varepsilon_{0}}{\sigma} \right]^{1/2} \right] \right]$$

$$= \left[\frac{2}{\omega \sigma \mu_{\Gamma} \mu_{0}} \left[1 + \left(\frac{\omega \varepsilon_{1} \varepsilon_{0}}{\sigma}\right)^{2} + \frac{\omega \varepsilon_{\Gamma} \varepsilon_{0}}{\sigma} \right]^{1/2} \right]$$

$$= \left[\frac{2}{\omega \sigma \mu_{\Gamma} \mu_{0}} \left[1 + \left(\frac{\omega \varepsilon_{1} \varepsilon_{0}}{\sigma}\right)^{2} + \frac{\omega \varepsilon_{\Gamma} \varepsilon_{0}}{\sigma} \right]^{1/2} \right]$$

$$= \left[\frac{2}{\omega \sigma \mu_{\Gamma} \mu_{0}} \left[1 + \left(\frac{\omega \varepsilon_{1} \varepsilon_{0}}{\sigma}\right)^{2} + \frac{\omega \varepsilon_{\Gamma} \varepsilon_{0}}{\sigma} \right]^{1/2} \right]$$

$$= \left[\frac{2}{\omega \sigma \mu_{\Gamma} \mu_{0}} \left[1 + \left(\frac{\omega \varepsilon_{1} \varepsilon_{0}}{\sigma}\right)^{2} + \frac{\omega \varepsilon_{\Gamma} \varepsilon_{0}}{\sigma} \right]^{1/2} \right]$$

$$= \left[\frac{2}{\omega \sigma \mu_{\Gamma} \mu_{0}} \left[1 + \left(\frac{\omega \varepsilon_{1} \varepsilon_{0}}{\sigma}\right)^{2} + \frac{\omega \varepsilon_{\Gamma} \varepsilon_{0}}{\sigma} \right]^{1/2} \right]$$

$$= \left[\frac{2}{\omega \sigma \mu_{\Gamma} \mu_{0}} \left[1 + \left(\frac{\omega \varepsilon_{1} \varepsilon_{0}}{\sigma}\right)^{2} + \frac{\omega \varepsilon_{\Gamma} \varepsilon_{0}}{\sigma} \right]^{1/2} \right]$$

$$= \left[\frac{2}{\omega \sigma \mu_{\Gamma} \mu_{0}} \left[1 + \left(\frac{\omega \varepsilon_{1} \varepsilon_{0}}{\sigma}\right)^{2} + \frac{\omega \varepsilon_{\Gamma} \varepsilon_{0}}{\sigma} \right]^{1/2} \right]$$

$$= \left[\frac{2}{\omega \sigma \mu_{\Gamma} \mu_{0}} \left[1 + \left(\frac{\omega \varepsilon_{1} \varepsilon_{0}}{\sigma}\right)^{2} + \frac{\omega \varepsilon_{\Gamma} \varepsilon_{0}}{\sigma} \right]^{1/2} \right]$$

$$= \left[\frac{2}{\omega \sigma \mu_{\Gamma} \mu_{0}} \left[1 + \left(\frac{\omega \varepsilon_{1} \varepsilon_{0}}{\sigma}\right)^{2} + \frac{\omega \varepsilon_{\Gamma} \varepsilon_{0}}{\sigma} \right]^{1/2} \right]$$

$$= \left[\frac{2}{\omega \sigma_{\Gamma} \mu_{0}} \left[1 + \left(\frac{\omega \varepsilon_{1} \varepsilon_{0}}{\sigma}\right)^{2} + \frac{\omega \varepsilon_{\Gamma} \varepsilon_{0}}{\sigma} \right]^{1/2} \right]$$

$$= \left[\frac{2}{\omega \sigma_{\Gamma} \mu_{0}} \left[1 + \left(\frac{\omega \varepsilon_{1} \varepsilon_{0}}{\sigma}\right)^{2} + \frac{\omega \varepsilon_{\Gamma} \varepsilon_{0}}{\sigma} \right]^{1/2} \right]$$

$$= \left[\frac{2}{\omega \sigma_{\Gamma} \mu_{0}} \left[1 + \left(\frac{\omega \varepsilon_{1} \varepsilon_{0}}{\sigma}\right)^{2} + \frac{\omega \varepsilon_{\Gamma} \varepsilon_{0}}{\sigma} \right] \right]$$

$$= \left[\frac{2}{\omega \sigma_{\Gamma} \mu_{0}} \left[1 + \left(\frac{\omega \varepsilon_{1} \varepsilon_{0}}{\sigma}\right)^{2} + \frac{\omega \varepsilon_{\Gamma} \varepsilon_{0}}{\sigma} \right] \right]$$

$$= \left[\frac{2}{\omega \sigma_{\Gamma} \mu_{0}} \left[1 + \left(\frac{\omega \varepsilon_{1} \varepsilon_{0}}{\sigma}\right)^{2} + \frac{\omega \varepsilon_{1} \varepsilon_{0}}{\sigma} \right] \right]$$

$$= \left[\frac{2}{\omega \sigma_{\Gamma} \mu_{0}} \left[1 + \left(\frac{\omega \varepsilon_{1} \varepsilon_{0}}{\sigma}\right)^{2} + \frac{\omega \varepsilon_{1} \varepsilon_{0}}{\sigma} \right] \right]$$

$$= \left[\frac{2}{\omega \sigma_{\Gamma} \mu_{0}} \left[1 + \left(\frac{\omega \varepsilon_{1} \varepsilon_{0}}{\sigma}\right)^{2} + \frac{\omega \varepsilon_{1} \varepsilon_{0}}{\sigma} \right] \right]$$

$$= \left[\frac{2}{\omega \sigma_{\Gamma} \mu_{0}} \left[1 + \left(\frac{\omega \varepsilon_{1} \varepsilon_{0}}{\sigma}\right)^{2} + \frac{\omega \varepsilon_{1} \varepsilon_{0}}{\sigma} \right] \right]$$

$$= \left[\frac{2}{\omega \sigma_{\Gamma} \mu_{0}} \left[1 + \left(\frac{\omega \varepsilon_{1} \varepsilon_{0}}{\sigma}\right) + \frac{\omega \varepsilon_{1} \varepsilon_{0}}{\sigma} \right] \right]$$

$$= \left[\frac{2}{\omega \sigma_{\Gamma} \mu_{0}} \left[1 + \left(\frac{\omega \varepsilon_{1} \varepsilon_{0}}{\sigma}\right) + \frac{\omega \varepsilon_{1} \varepsilon_{0}}{\sigma} \right] \right]$$

$$= \left[\frac{2}{\omega \sigma_{\Gamma} \mu_{0}} + \frac{\omega \varepsilon_{1} \varepsilon_{0}}{\sigma}\right]$$

$$= \left[\frac{2}{\omega \sigma_{\Gamma} \mu_{0}} + \frac{\omega \varepsilon_{1} \varepsilon_{0}}{\sigma}\right] \right]$$

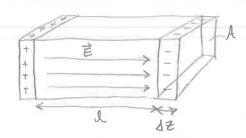
$$= \left[\frac{2}{\omega \sigma_{\Gamma} \mu_{0}} + \frac{\omega \varepsilon_{1} \varepsilon_{0}}{\sigma}\right]$$

low frequency Skin depth

$$\delta \rightarrow \frac{2}{\sigma} \sqrt{\frac{\epsilon_r \epsilon_o}{\mu_r \mu_0}}$$

- Bulk plasmons: oscillations in the free electron gas inside a metal's volume

· Suppose that a metal becomes polarized



· The charge separation creates an electric field:

$$\mathcal{E}_0 \to A = -gnA \Delta Z \leftarrow Gauss's law around box of hegative charge
$$= -\frac{gn}{\mathcal{E}_0} \Delta Z$$$$

· What is the frequency of the oscillation?

$$gE = -\frac{g^2 m}{\epsilon_0} \Delta z = m(\Delta z) \quad \text{by } F = ma$$

$$\frac{ng^2}{m\epsilon_0} \Delta z = \omega^2 \Delta z$$

$$\Rightarrow \omega = \omega_p = \frac{ng^2}{m\epsilon_0}$$

$$\text{bulk plasma frequency}$$

· Consider the oscillation to that of an LC circuit

$$w = \frac{1}{JL_{K}C} \Rightarrow \frac{1}{w_{p}^{2}} = L_{K}C = L_{K} \frac{\varepsilon_{0}A}{\vartheta}$$

$$\Rightarrow L_{K} = \frac{\vartheta}{\varepsilon_{0}A} \frac{1}{w_{p}^{2}} = \frac{\vartheta}{\varepsilon_{0}A} \frac{m\varepsilon_{0}}{mg^{2}}$$

$$\frac{1}{k_{0}} = \frac{1}{k_{0}} \frac{m\varepsilon_{0}}{mg^{2}}$$

This is the effective inductance associated with the inertia of electrons: at very high frequencies, the response of the metal is delayed because the electrons need to move

· This can also be derived from the AC conductivity of a metal, from the Drude model

Reflection of circularly polarized light from a metal surface

· Consider the RCP incident light below:

$$\overrightarrow{E}_{I}(z,t) = \frac{E_{0}}{\sqrt{2}} \cos(kz - wt) \hat{\chi} + \frac{E_{0}}{\sqrt{2}} \sin(kz - wt) \hat{y}$$

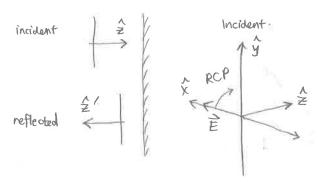
RCP meta

• The electric field must vanish at the surface of a perfect conductor, so $\vec{E}_R = -\vec{E}_I$:

$$\vec{E}_{R}(z,t) = -\frac{F_{0}}{\sqrt{2}}\cos(-kz-\omega t)\hat{x} - \frac{F_{0}}{\sqrt{2}}\sin(-kz-\omega t)\hat{y}$$

This ensures zero field at z = 0 while Ex propagates along - 2

· Now the subtle part: we need to do a coordinate transformation:



refected

\hat{y}'

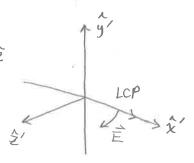
\hat{z}'

\hat{E}

Incident coordinate System is right-handed $\hat{x} \times \hat{y} = \hat{z}$

① First, flip propagation direction so $\hat{z}' = -\hat{z}'$ while $\hat{x}' = \hat{x}$, $\hat{y}' = \hat{y}$ But now $\hat{x}' \times \hat{y}' = -\hat{z}'$.

② Next, flip the x-axis so
$$\hat{x}' = -\hat{x}$$
, $\hat{y}' = \hat{y}$, $\hat{z}' = -\hat{z}$
Now we have $\hat{x}' \times \hat{y}' = \hat{z}'$



· With this transformation, the reflected wave is:

$$\vec{E}_{R} = \frac{E_{0}}{\sqrt{2}} \cos(kz'-\omega t)\hat{x}' - \frac{E_{0}}{\sqrt{z}} \sin(kz'-\omega t)\hat{y}'$$

This describes a left circularly polarized wave, so there is a handedress reversal of circularly polarized light when reflected from a conductor